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Curves

1.1 AFFINE SPACE

Let E^n be an *n*-dimensional affine space and V^n be its associated *n*-dimensional vector space. Thus there is a mapping $E^n \times V^n \to E^n$, which assigns to each pair $(p; \vec{v})$, where $p \in E^n$ is a point of affine space E^n and $\vec{v} \in V^n$ is a vector of associated vector space V^n , the point of affine space E^n , which will be denoted by $p + \vec{v}$. This mapping satisfies the conditions:

- \mathbf{A}_1) for each point $p \in E^n$ it holds $p + \vec{0} = p$,
- \mathbf{A}_2) for each pair of points $(p;q) \in E^n \times E^n$ there exists a unique vector $\vec{v} \in V^n$ such that $p + \vec{v} = q$,
- \mathbf{A}_3) for each point $p \in E^n$ and two vectors $\vec{v}, \vec{w} \in V^n$ it holds

$$(p+\vec{v}) + \vec{w} = p + (\vec{v} + \vec{w}),$$

where $\vec{v} + \vec{w}$ at the right-hand side means the addition of vectors in a vector space V^n .

The vector \vec{v} determined by two points p, q of affine space E^n according to the axiom \mathbf{A}_2 will be denoted by \vec{pq} .

Let $\vec{v}_1, \vec{v}_2, \ldots, \vec{v}_m$, where $1 \leq m < n$, be linearly independent vectors of associated vector space V^n and p be a point of affine space E^n . A point of E^n and m linearly independent vectors of V^n determine the *m*-dimensional plane in E^n . This plane passes through a point p and is defined as follows

$$\mathscr{P} = \{ p + \sum_{\alpha=1}^{m} t_{\alpha} \vec{v}_{\alpha} : p \in E^{n}; \ \vec{v}_{1}, \vec{v}_{2}, \dots, \vec{v}_{m} \in V^{n}, \ t_{1}, t_{2}, \dots, t_{m} \in \mathbb{R} \}.$$
(1.1)

Particularly, any pair $(p; \vec{v})$, where p is a point and \vec{v} is a non-zero vector, determines the straight line

$$\mathscr{L} = \{ p + t \, \vec{v} : p \in E^n, \ \vec{v} \in V^n, t \in \mathbb{R}, \vec{v} \neq \vec{0} \}.$$

$$(1.2)$$

The straight line \mathscr{L} passes through a point p and \vec{v} is its directional vector.

Let $\{\vec{e_1}, \vec{e_2}, \ldots, \vec{e_n}\}$ be a basis for associated vector space V^n . Any vector \vec{v} of V^n can be written in the form

$$\vec{v} = \sum_{i=1}^{n} v^{i} \vec{e}_{i},$$
 (1.3)

where the real numbers v^1, v^2, \ldots, v^n are called the coordinates of a vector \vec{v} . The right-hand side of (1.3) can be written in a more concise form by means of the Einstein summation convention over repeated subscript and superscript. Hence we can write (1.3) in the equivalent form $\vec{v} = v^i \vec{e}_i$ and in what follows we will regularly use this convention to write formula more compactly. The formula (1.3) allows us to identify a vector space V^n with the *n*-dimensional vector space \mathbb{R}^n by identifying a vector \vec{v} with its coordinate vector $(v^1, v^2, \ldots, v^n) \in \mathbb{R}^n$. Assuming this identification, we will write $\vec{v} = (v^1, v^2, \ldots, v^n)$. Then the addition of vectors and the multiplication by real numbers in V^n can be expressed as follows

$$\vec{v} + \vec{w} = (v^1 + w^1, v^2 + w^2, \dots, v^n + w^n),$$

$$a \vec{v} = (av^1, av^2, \dots, av^n), \ a \in \mathbb{R},$$

where $\vec{v} = (v^1, v^2, \dots, v^n), \vec{w} = (w^1, w^2, \dots, w^n).$

Let $\mathfrak{R} = \{O; \vec{e_1}, \vec{e_2}, \dots, \vec{e_n}\}$ be an affine frame in the space E^n , where $O \in E^n$ is an origin of an affine frame \mathfrak{R} and $\{\vec{e_1}, \vec{e_2}, \dots, \vec{e_n}\}$ is a basis for the vector space V^n . Affine frame \mathfrak{R} determines the system of coordinates in the space E^n . Indeed given a point $p \in E^n$ we can assign to it the position vector $\vec{r} \in V^n$, where the initial point of \vec{r} is at the origin O and the end point is at the point p. Then we can expand the position vector \vec{r} in terms of a basis

$$\vec{r} = \sum_{i=1}^{n} p^i \, \vec{e_i},$$

and the coordinates p^1, p^2, \ldots, p^n of the position vector are the coordinates of a point p. It is easy to see that if p^1, p^2, \ldots, p^n are the coordinates of a point p and q^1, q^2, \ldots, q^n are the coordinates of a point q, then the vector \overrightarrow{pq} , determined by points p, q, has the coordinates $q^1 - p^1, q^2 - p^2, \ldots, q^n - p^n$.

Thus an affine frame \mathfrak{R} defines a coordinate system in affine space E^n and the corresponding coordinates will be denoted by x^1, x^2, \ldots, x^n . In what follows we will consider coordinates x^1, x^2, \ldots, x^n from two different points of view. The first point of view on coordinates x^1, x^2, \ldots, x^n is traditional, i.e. we consider x^1, x^2, \ldots, x^n as an ordered sequence of n real numbers which determines a single point in an affine space E^n . The second point of view will be used in the framework of an algebra of smooth functions on an affine space E^n and according to this approach we will consider x^i as a function whose domain is an affine space E^n (or its open subset) and the value at a point p is the *i*th coordinate of a point p, i.e. $x^i(p) = p^i$. In this case, we will call x^i the *i*th coordinate function. Obviously the coordinate system defined by an affine frame \mathfrak{R} allows us to identify an affine space E^n with the space \mathbb{R}^n .

If we choose a different frame $\Re' = \{O'; \vec{e}'_1, \vec{e}'_2, \dots, \vec{e}'_n\}$ in an affine space E^n then we get a different coordinate system whose affine coordinates will be denoted by

(1.4)

 $\tilde{x}^1, \tilde{x}^2, \ldots, \tilde{x}^n$. Let us find how the coordinates x^1, x^2, \ldots, x^n determined by an affine frame \mathfrak{R} can be expressed in terms of coordinates $\tilde{x}^1, \tilde{x}^2, \ldots, \tilde{x}^n$ determined by an affine frame \mathfrak{R}' . Let $A = (A_j^i)$ be a transition matrix from a basis $\{\vec{e}_1, \vec{e}_2, \ldots, \vec{e}_n\}$ to a basis $\{\vec{e}_1', \vec{e}_2', \ldots, \vec{e}_n'\}$, i.e. $\vec{e}_i' = A_i^j \vec{e}_j$. The position vector \vec{r} of a point $x \in E^n$ in a frame \mathfrak{R} can be written as $\vec{r} = x^i \vec{e}_i$ and, analogously, the position vector \vec{r}' of the same point x in a new frame \mathfrak{R}' can be written as $\vec{r}' = \tilde{x}^i \vec{e}_i'$. We have

$$\vec{e}'_j = A^i_j \vec{e}_i, \ \vec{r} = \vec{r}' + \overrightarrow{OO'}, \ \overrightarrow{OO'} = a^i \vec{e}_i.$$

Hence

$$x^{i} \vec{e}_{i} = \tilde{x}^{j} \vec{e}_{j}' + a^{i} \vec{e}_{i} = A_{j}^{i} \tilde{x}^{j} \vec{e}_{i} + a^{i} \vec{e}_{i} = (A_{j}^{i} \tilde{x}^{j} + a^{i}) \vec{e}_{i},$$

 $x^i = A^i_i \, \tilde{x}^j + a^i.$

and

Formula (1.4) is called an affine transformation of coordinates. Note that an *n*th order matrix $A = (A_i^i)$ in (1.4) is a regular matrix, that is, Det $A \neq 0$.

So far, we have considered a space that has only an affine structure. If we want to use the concepts of the length of a vector, a distance between two points, an angle between two vectors, we must assume that an affine space E^n has a Euclidean structure. Let V^n be a Euclidean space. A Euclidean structure in V^n is given by a scalar (or inner) product of vectors which we will denote by $\langle \vec{v}, \vec{w} \rangle$. This inner product is real-valued, symmetric, bilinear and positive-definite. The length of a vector \vec{v} will be denoted by $\|\vec{v}\|$, i.e. $\|\vec{v}\|^2 = \langle \vec{v}, \vec{v} \rangle$. The angle α between two *n*-dimensional vectors \vec{v}, \vec{w} is defined by

$$\cos \alpha = \frac{\langle \vec{v}, \vec{w} \rangle}{\|\vec{v}\| \|\vec{w}\|}.$$

In what follows an affine space E^n such that its associated vector space V^n is endowed with a scalar product will be called an *affine Euclidean space*. In affine Euclidean space we can use orthonormal bases. In what follows we will usually assume that a coordinate system in an affine Euclidean space is constructed by means of an orthonormal frame $\Re = \{O; \vec{e_1}, \vec{e_2}, \dots, \vec{e_n}\}$, where $\langle \vec{e_i}, \vec{e_j} \rangle = \delta_{ij}$. The coordinate system induced by an orthonormal frame will be called a *Cartesian coordinate system*. An advantage of a Cartesian coordinates is that formulae for a length of a vector, for a distance between two points and for an angle have a simple form. For instance the scalar product of vectors \vec{v}, \vec{w} , the square of the length of a vector \vec{v} and the cosine of the angle between vectors \vec{v}, \vec{w} can be written in Cartesian coordinates as follows

$$\langle \vec{v}, \vec{w} \rangle = \sum_{i=1}^{n} v^{i} w^{i}, \quad \|\vec{v}\|^{2} = \sum_{i=1}^{n} (v^{i})^{2}, \quad \cos \phi = \frac{\sum_{i=1}^{n} v^{i} w^{i}}{\sum_{i=1}^{n} (v^{i})^{2} \sum_{j=1}^{n} (w^{j})^{2}}.$$

The distance between two points p, q of an affine Euclidean space will be denoted by ||q - p|| and $||q - p|| = ||\vec{pq}||$, and the distance between points p, q can be expressed in terms of Cartesian coordinates as follows

$$||q-p|| = \sqrt{\sum_{i=1}^{n} (q^i - p^i)^2}.$$

The transformation law for Cartesian coordinates has the same form (1.4) as the transformation law for affine coordinates, but in the case of Cartesian coordinates a matrix A in this formula is orthogonal, i.e. $A \in O(n)$. It should be mentioned that if $A \in SO(n)$ then an orientation of Cartesian coordinate system will remain the same.

The Euclidean structure of an affine Euclidean space E^n can be used to equip this space with a topology. We briefly recall how this topology in E^n is defined. An open ball $B_r(p)$ of radius r and centered at a point p is the following set of points

$$B_r(p) = \{ q \in E^n : ||p - q|| < r \}.$$

Now a subset $U \subset E^n$ is said to be an *open subset* in E^n if for any point $q \in U$ there exists an open ball $B_r(q)$ such that $B_r(q) \subset U$. Hence we have the collection \mathcal{T} of open subsets in E^n and it can be proved that this family satisfies all axioms of a topological space. Thus an affine Euclidean space E^n endowed with the collection of open subsets \mathcal{T} is a topological space. It is worth to mention that an affine Euclidean space E^n as topological space is a Hausdorff space with a countable base.

Let \mathscr{P} be an *m*-dimensional plane (1.1), passing through a point p and determined by linear independent vectors $\vec{v}_1, \vec{v}_2, \ldots, \vec{v}_m$. If p^1, p^2, \ldots, p^n are coordinates of a point $p, v_k^1, v_k^2, \ldots, v_k^n$ are coordinates of kth vector \vec{v}_k , then \mathscr{P} can be described by the parametric equation

$$x^{i} = p^{i} + t_{1}v_{1}^{i} + t_{2}v_{2}^{i} + \ldots + t_{m}v_{m}^{i}, i = 1, 2, \ldots, n,$$

where x^1, x^2, \ldots, x^n are coordinates of an arbitrary point of *m*-dimensional plane \mathscr{P} . Particularly, the parametric equation of a straight line \mathscr{L} (1.2) has the form

$$x^{i} = p^{i} + t v^{i}, \ i = 1, 2, \dots, n.$$

For the purposes of differential geometry it is useful to introduce a notion of a vector at a point of Euclidean space. A vector at a point p of Euclidean space E^n is a pair $(p; \vec{v})$, where $p \in E^n$ and $\vec{v} \in V^n$. A vector at a point $(p; \vec{v})$ will be denoted by \mathbf{v}_p , i.e. $\mathbf{v}_p = (p; \vec{v})$. It should be emphasized here that we do not consider a vector at a point $\mathbf{v}_p = (p; \vec{v})$ from the point of view of the affine structure of space E^n , that is, we do not identify a pair $(p; \vec{v})$ with the point $p + \vec{v}$. A vector at a point p will be also called a *tangent vector to Euclidean space* at a point p. The set of all tangent vectors to E^n at a point p will be denoted by T_pE^n . Hence $T_pE^n = \{p\} \times V^n$. If U is a subset of Euclidean space E^n then all vectors at points of U is the direct product $U \times V^n$, which will be denoted by T(U). It is obvious that $T(U) = \bigcup_{p \in U} T_pE^n$, where \bigcup stands for disjoint union of all $T_pE^n, p \in U$. We can define a projection $\pi : T(U) \to U$ of this direct product onto subset U by the formula $\pi(\mathbf{v}_p) = p$, where $p \in U$.

The vector space structure of V^n induces the vector space structure in the set of tangent vectors $T_p E^n$ if for two tangent vectors $\mathbf{v}_p = (p; \vec{v}), \mathbf{w}_p = (p; \vec{v})$ we define

$$\mathbf{v}_p + \mathbf{w}_p = (p; \vec{v} + \vec{w}), \quad a \, \mathbf{v}_p = (p; a \, \vec{v}), \quad a \in \mathbb{R}.$$

Similarly the Euclidean structure of V^n induces the Euclidean structure in the vector

space $T_p E^n$ if we define the inner product of two tangent vectors $\mathbf{v_p}, \mathbf{w_p}$ by the formula

$$<\mathbf{v}_p,\mathbf{w}_p>=$$

The Euclidean vector space $T_p E^n$ will be called a *tangent space to affine Euclidean* space E^n at a point p. It is worth to mention that a tangent vector $\mathbf{v}_p = (p; \vec{v})$ determines by means of affine structure of E^n the straight line, which passes a point p in the direction of a vector \vec{v} and $p + t \vec{v}$ is the parametric equation of this straight line, which in Cartesian coordinates takes the form

$$x^{i} = p^{i} + v^{i} t, \ i = 1, 2 \dots, n,$$

where p^i are the coordinates of a point p, v^i are the coordinates of a vector \vec{v} and t is a parameter that runs through the whole real line. The tangent space $T_p E^n$ at each point p is *n*-dimensional vector space. We can construct the orthonormal basis for this vector space by means of an orthonormal basis $\vec{e_1}, \vec{e_2}, \ldots, \vec{e_n}$ of the Euclidean space V^n if we define $\mathbf{e}_{p,i} = (p; \vec{e_i}), i = 1, 2, \ldots, n$. Then any tangent vector $\mathbf{v}_p = (p; \vec{v})$ can be written

$$\mathtt{v}_p = v^i \, \mathtt{e}_{p,i}$$

1.2 CURVILINEAR COORDINATE SYSTEMS

In the previous section we showed that an affine frame \Re in an affine space E^n defines an affine coordinate system in this space. However an affine coordinate system is a special case of a coordinate system in affine space. Other well-known coordinate systems are polar coordinates in a plane, spherical and cylindrical coordinates in three-dimensional space. Coordinate systems of this kind are called curvilinear coordinates in space. The aim of this section is to give a general definition of a curvilinear coordinate system.

Let E^n be an affine space and \mathfrak{R} be an affine frame in this space. An affine frame \mathfrak{R} induces the coordinate system in E^n whose affine coordinates will be denoted by x^1, x^2, \ldots, x^n . An affine coordinate system determines a mapping from affine space E^n to an *n*-dimensional arithmetic space \mathbb{R}^n . This mapping assigns to each point of an affine space an ordered set of *n*-real numbers, which we call coordinates of a point *p*. Let us denote this mapping by $\phi : E^n \to \mathbb{R}^n$. Then for a point $x \in E^n$ we have $\phi(x) = (x^1, x^2, \ldots, x^n) \in \mathbb{R}^n$, where x^1, x^2, \ldots, x^n are affine coordinates of a point $x \in E^n$ with respect to an affine frame \mathfrak{R} . Mapping ϕ has two very important properties which are actually the essence of the concept of a coordinate system. First this mapping is a bijection, that is, a one-to-one (injective) and onto (surjective) mapping. Hence ϕ has the inverse which will be denoted by $\phi^{-1} : \mathbb{R}^n \to E^n$. Second the mapping ϕ and its inverse are continuous mappings, i.e. ϕ is a homeomorphism between topological spaces E^n and \mathbb{R}^n . Evidently ϕ depends on a choice of an affine frame \mathfrak{R} , but its property to be a homeomorphism is invariant with respect to a choice of an affine frame.

By analogy with an affine coordinate system we can now give a general definition of a coordinate system in an affine space E^n . Let U be an open set in an affine space

 E^n and ψ be a bijection from a set U onto an open subset W of an n-dimensional arithmetic space \mathbb{R}^n , that is, $\psi: U \subset E^n \to W \subset \mathbb{R}^n$, $\psi(U) = W$. Thus a mapping ψ maps each point x of an open subset U to an ordered set of n real numbers (x^1, x^2, \ldots, x^n) . It is important that if $x \neq y$ and $\psi(x) = (x^1, x^2, \ldots, x^n), \psi(y) = (y^1, y^2, \ldots, y^n)$ then $(x^1, x^2, \ldots, x^n) \neq (y^1, y^2, \ldots, y^n)$. Hence a mapping ψ defines a one-to-one parametrization of points in an open subset U of affine space E^n and therefore could be considered as a coordinate system in U. However this is not enough for the purposes of differential geometry. Differential geometry uses differential calculus to study curves and surfaces. A central concept here is the concept of a smooth function, that is, a function that has partial derivatives of any order and they are continuous functions. It is important for purposes of differential geometry to have an invariant concept of a smooth function which does not depend on a choice of coordinate system. Thus we must impose additional conditions on a mapping ψ and for this we need a notion of a diffeomorphism.

Let V, V' be an open subsets of an *n*-dimensional arithmetic space \mathbb{R}^n and $g: V \to V'$ be a mapping.

Definition 1.1. A mapping $g: V \to V'$ is called a diffeomorphism from V to V' if it satisfies the following conditions

- 1. g is a smooth bijection,
- 2. $g^{-1}: V' \to V$ is also a smooth mapping.

Let $\phi : U \subset E^n \to V \subset \mathbb{R}^n, \phi(U) = V$ is an affine coordinate system in an open subset U defined with the help of an affine frame.

Definition 1.2. A bijection $\psi : U \subset E^n \to W \subset \mathbb{R}^n$ is called a coordinate system in an open subset U of an affine space C^{∞} -compatible with coordinate system ϕ if the mapping $\chi = \phi \circ \psi^{-1} : W \subset \mathbb{R}^n \to V \subset \mathbb{R}^n$ is a diffeomorphism from W to V.

Obviously χ is a bijection because it is a composition of two bijections ϕ and ψ^{-1} . Hence Definition (1.2) requires a smoothness of f and χ^{-1} . In what follows we will tacitly assume that the smoothness condition of a coordinate system is satisfied and usually omit the word " C^{∞} -compatible". Let $f: U \to \mathbb{R}$ be a real-valued function on an open subset $U \subset E^n$ and $\phi: U \to \mathbb{R}^n$, $\phi(U) = V$ be an affine system of coordinates. We will call a function f a smooth function if the real-valued function of n-variables $\tilde{f} = f \circ \phi^{-1}: V \to \mathbb{R}$ is a smooth function. It is important here that a notion of smoothness does not depend on a choice of coordinate system (C^{∞} -compatible with ϕ) then a function f is a smooth function if and only if it is a smooth function in a coordinate system ψ , i.e. $\tilde{f} = f \circ \psi^{-1}: W \to \mathbb{R}$ is a smooth function if and only if it is a smooth function in a coordinate system ψ , i.e. $\tilde{f} = f \circ \psi^{-1}: W \to \mathbb{R}$ is a smooth function of smooth function. This follows from the relations $\tilde{f} = \tilde{f} \circ \chi^{-1}$, $\tilde{f} = \tilde{f} \circ \chi$, where $\chi = \phi \circ \psi^{-1}$, and the fact that a composition of smooth mappings is a smooth mapping.

Let us denote by x^1, x^2, \ldots, x^n the affine coordinates in U defined by ϕ and by $\tilde{x}^1, \tilde{x}^2, \ldots, \tilde{x}^n$ the coordinates in U defined by a coordinate system ψ . Then it is easy

to see that a diffeomorphism $\chi: W \to V$ is uniquely determined by a set of *n* realvalued smooth functions f^1, f^2, \ldots, f^n whose domain is $W \subset \mathbb{R}^n$. By other words we can write

$$\chi = (f^1, f^2, \dots, f^n), \ \chi : (\tilde{x}^1, \tilde{x}^2, \dots, \tilde{x}^n) \in W \mapsto (x^1, x^2, \dots, x^n) \in V,$$

where

$$x^{1} = f^{1}(\tilde{x}^{1}, \tilde{x}^{2}, \dots, \tilde{x}^{n}), \ x^{2} = f^{2}(\tilde{x}^{1}, \tilde{x}^{2}, \dots, \tilde{x}^{n}), \dots, x^{n} = f^{n}(\tilde{x}^{1}, \tilde{x}^{2}, \dots, \tilde{x}^{n}),$$
(1.5)

The formulae (1.5) will be referred to as a transition from an affine coordinate system ϕ to a coordinate system ψ and we will write them in a more compact and symbolic way $x^i = f^i(\tilde{x}^j)$. If the functions f^1, f^2, \ldots, f^n in (1.5) are not linear then a coordinate system ψ will be called a *curvilinear coordinate system*.

In (1.5) x^1, x^2, \ldots, x^n are affine coordinates on an open set U. However we can look at this formula from a more general point of view, that is, assume that x^1, x^2, \ldots, x^n is a system of curvilinear coordinates. In this case formula (1.5) describes the transition from one coordinate system (not necessarily affine) to another. Therefore we will call (1.5) a transition from one coordinate system to another.

We can consider $\chi: V \subset \mathbb{R}^n \to W \subset \mathbb{R}^n$ as a transformation in an *n*-dimensional arithmetic space \mathbb{R}^n . Let us consider the Jacobian $\mathfrak{J}(\chi)$ of a transformation χ

$$\mathfrak{J}(\chi) = \operatorname{Det}\left(\frac{\partial x^i}{\partial \tilde{x}^j}\right). \tag{1.6}$$

Proposition 1.1. The Jacobian of a transition from one coordinate system to another $\chi: W \to V$ is nonzero at each point of W.

Proof. Due to the fact that $\chi: W \to V$ is a diffeomorphism it is a bijection and hence it is a reversible transformation, that is, we have an inverse transformation χ^{-1} : $V \to W$. This means that in formula (1.5) we can express coordinates $\tilde{x}^1, \tilde{x}^2, \ldots, \tilde{x}^n$ in terms of coordinates x^1, x^2, \ldots, x^n , i.e.

$$\tilde{x}^1 = g^1(x^1, x^2, \dots, x^n), \ \tilde{x}^2 = g^2(x^1, x^2, \dots, x^n), \ \dots, \tilde{x}^n = g^n(x^1, x^2, \dots, x^n).$$

According to the definition of a diffeomorphism the functions g^1, g^2, \ldots, g^n are smooth functions. These functions satisfy

$$f^{i}(g^{1}(x^{1}, x^{2}, \dots, x^{n}), g^{2}(x^{1}, x^{2}, \dots, x^{n}), \dots, g^{n}(x^{1}, x^{2}, \dots, x^{n})) \equiv x^{i}$$

Differentiating the both sides of this formula with respect to x^j and taking into account that a function f^i at the left-hand side of this formula is a composite function of variables x^1, x^2, \ldots, x^n we get

$$\frac{\partial x^i}{\partial \tilde{x}^k} \ \frac{\partial \tilde{x}^k}{\partial x^j} = \delta^i_j.$$

Recall that we are using the Einstein summation convention and according to this convention there is a summation over k at the left-hand side of the above formula.

If we now look at the above formula from the point of view of matrix calculus we see that on the left-hand side of this formula there is a product of two matrices $\left(\frac{\partial x^i}{\partial x^j}\right)$, $\left(\frac{\partial \tilde{x}^k}{\partial x^j}\right)$ and on the right-hand side there is an identity matrix. This implies that matrix $\left(\frac{\partial \tilde{x}^k}{\partial x^j}\right)$ is the inverse of matrix $\left(\frac{\partial x^i}{\partial \tilde{x}^k}\right)$, that is, the matrix $\left(\frac{\partial x^i}{\partial \tilde{x}^k}\right)$ is invertible, which means it is regular, that is, the determinant of this matrix is nonzero. But the determinant of matrix $\left(\frac{\partial x^i}{\partial \tilde{x}^k}\right)$ is the Jacobian of a transition from one coordinate system to another (1.6).

Now we will give a few examples of curvilinear coordinates.

Example 1.1. The first example of a curvilinear coordinate system is a polar coordinate system on a plane. Let E^2 be a two-dimensional affine Euclidean space which will be referred to as a plane. We construct a coordinate system on a plane E^2 which is different from an affine coordinate system constructed with the help of a frame. Fix a point O in a plane E^2 and a directed straight line L which passes through a point O. A point O will be referred to as a pole of a polar coordinate system and L as a polar axis of this system. Given a point p of a plane E^2 , which is different from a pole O, we define its polar coordinates as follows. The distance between the pole O and a point p be the first polar coordinate of a point p and it will be denoted by r. Now we



Figure 1.1 Polar coordinates

rotate the polar axis L counterclockwise until it passes through a point p. The angle obtained by this rotation is the second polar coordinate of a point p, it is denoted by φ and is called a polar angle of a point p.

Obviously a polar angle in the case of a pole O is not defined. Thus a polar coordinate system defines a mapping which assigns to each point of a plane (with the exception of the pole O) the pair of real numbers (r, φ) . Let $U = E^2 \setminus L$, \mathbb{R}^2 be a two-dimensional arithmetic space, whose coordinates are denoted by $r, \varphi, W =$ $\{(r, \varphi) \in \mathbb{R}^2 : r > 0, 0 < \varphi < 2\pi\}$ and $\psi : U \to W$ be a mapping defined by polar coordinates. Then ψ is a bijection from $U \subset E^2$ onto open subset $W \subset \mathbb{R}^2$. If we endow a plane E^2 with the orthonormal frame $\{O, \vec{e_1}, \vec{e_2}\}$ whose origin coincides with the pole and positive direction of x-axis coincides with the polar axis L then the Cartesian coordinates x, y can be expressed in terms of polar coordinates as follows

$$x = r \cos \varphi, \quad y = r \sin \varphi. \tag{1.7}$$

Conversely the polar coordinates can be expressed in terms of Cartesian coordinates. Let \mathbb{R}^2 be a second copy of the 2-dimensional arithmetic space whose coordinates are Cartesian coordinates x, y and V be an open subset of \mathbb{R}^2 defined by $V = \mathbb{R}^2 \setminus Ox$, where $Ox = \{(x, y) \in \mathbb{R}^2 : x \ge 0, y = 0\}$ is the x axis. Then the polar coordinates can be expressed in terms of Cartesian coordinates as follows

$$r = \sqrt{x^2 + y^2}, \quad \varphi = \begin{cases} \arccos \frac{x}{\sqrt{x^2 + y^2}} & y \ge 0, \\ 2\pi - \arccos \frac{x}{\sqrt{x^2 + y^2}} & y < 0. \end{cases}$$

Thus we see that the functions in these formulae determine mutually inverse bijections of two subsets V and W. We leave it as an exercise to a reader to show that these bijections are smooth.

The above description of polar coordinates shows that a polar coordinate system can be considered as a coordinate system in the sense of Definition (1.2). However in practical matters, for example, in the theory of plane curves it is convenient to use polar coordinates in a broader sense. In this approach we will assume that in order to find a polar angle of a point we can rotate a polar axis around the pole counterclockwise (positive angle) or clockwise (negative angle) any number of times. In this case a polar angle φ can be any real number and we will assume that polar coordinates $(r, \varphi \pm 2\pi k)$, where k is any integer, determine the same point of a plane.

Example 1.2. One of the most well-known coordinate systems in a 3-dimensional space is a cylindrical coordinate system. We fix a plane \mathfrak{P} in three-dimensional affine Euclidean space E^3 , a point O on this plane and two directed straight lines L, L' (i.e. endowed with fixed positive direction) that pass through point O. Thus the point O divides the directed straight lines L, L' into positive and negative parts. Line L' is perpendicular to the plane \mathfrak{P} , and line L lies in the plane \mathfrak{P} . Considering point O as the pole of a polar coordinate system, and line L as the polar axis we get the polar coordinate system on the plane \mathfrak{P} described in detail in the previous example. Let us denote the coordinates of this polar coordinate system as r, φ . Consider some point p of an affine Euclidean space E^3 different from the point O. Let r, φ be the polar coordinates of the orthogonal projection p' of a point p onto the plane \mathfrak{P} , and z be the length of its orthogonal projection onto the directed straight line L' taken with plus if the projection of p lies on the positive part of L' and with minus otherwise. The set of numbers (r, φ, z) is called the cylindrical coordinates of a point p and the straight line L' is called a z-axis.

If we now equip a 3-dimensional affine Euclidean space E^3 with a Cartesian coordinate system such that an origin of this coordinate system will be the point O, a positive part of x-axis will be the polar axis L and a z-axis will be the line L' (with a positive direction defined on it) then the Cartesian coordinates x, y, z are expressed in terms of cylindrical coordinates r, φ, z as follows

$$x = r\cos\varphi, \ y = r\sin\varphi, \ z = z.$$
 (1.8)



Figure 1.2 Cylindrical coordinates

This formula will be referred to as a transition from cylindrical coordinates to Cartesian ones. If we choose the open subset $U = E^3 \setminus L$ of an affine Euclidean space E^3 as a domain of a cylindrical coordinate system then a cylindrical coordinate system bijectively maps this subset to the open subset W of a 3-dimensional arithmetic space \mathbb{R}^3 , where

$$W = \{ (r, \varphi, z) \in \mathbb{R}^3 : r > 0, 0 < \varphi < 2\pi, -\infty < z < \infty \}.$$
(1.9)

Similar to the case of polar coordinates we can consider an open subset V of 3dimensional arithmetic space \mathbb{R}^3 (whose coordinates are denoted by x, y, z) defined by $V = \mathbb{R}^3 \setminus Ox$, where

$$Ox = \{ (x, y, z) \in \mathbb{R}^3 : x \ge 0, y = 0, z = 0 \}.$$

Then the formula (1.8) defines the smooth bijection from W to V and the following formula (the transition from Cartesian coordinates to cylindrical ones)

$$r = \sqrt{x^2 + y^2 + z^2}, \quad \varphi = \begin{cases} \arccos \frac{x}{\sqrt{x^2 + y^2}} & y \ge 0, \\ 2\pi - \arccos \frac{x}{\sqrt{x^2 + y^2}} & y < 0, \end{cases} \quad z = z$$

defines its inverse which is also smooth. Hence a cylindrical coordinate system can be considered as a curvilinear coordinate system in a 3-dimensional affine Euclidean space E^3 .

Example 1.3. A spherical coordinate system in a 3-dimensional affine Euclidean space E^3 is constructed similarly to a cylindrical coordinate system, that is, we fix a plane \mathfrak{P} , a point O on this plane and two directed straight lines L, L^prime passing through a point O where line L lies in a plane \mathfrak{P} and line L' is perpendicular to \mathfrak{P} . Evidently the point O and the directed straight line L defines the polar coordinate system in the plane \mathfrak{P} . Let p be an arbitrary point of affine Euclidean space E^3 different from O. The first spherical coordinate of a point p is the distance between O

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and p. Let us denote this first spherical coordinate by r. Now let p' be the orthogonal projection of a point p to the plane \mathfrak{P} . Then the second spherical coordinate of a point p is the polar angle of the orthogonal projection p' in the polar coordinate system in the plane \mathfrak{P} defined by O (pole) and directed line L (polar axis). Let us denote this



Figure 1.3 Spherical coordinates

second spherical coordinate by φ . The third spherical coordinate of a point A is the angle obtained by rotation of the positive part of line L' the shortest way around the point O until it passes through a point p. This third spherical coordinate will be denoted by ϑ . The set of three real numbers (r, φ, ϑ) is called spherical coordinates of a point p. A spherical coordinate system can be considered as a curvilinear coordinate system in E^3 in the sense of Definition (1.2) if its domain the open subset $U = E^3 \setminus (L \cup L')$ of E^3 , that is, we remove from 3-dimensional affine Euclidean space E^3 the points of straight lines L, L'. Then a spherical coordinate system induces the bijection $p \in U \to (r, \varphi, \vartheta) \in W$ of U to the open subset $W \subset \mathbb{R}^3$, where

$$W = \{ (r, \varphi, \vartheta) \in \mathbb{R}^3 : r > 0, 0 < \varphi < 2\pi, 0 < \vartheta < \pi \}.$$

If we endow a 3-dimensional space E^3 with a Cartesian coordinate system by placing its origin to the point O, choosing directed straight line L as x-axis, directed straight line L' as z-axis and by constructing y-axis to be passing through O, perpendicular to x-axis, z-axis (and directed so that Cartesian coordinate system will be right-handed) then we can pass from spherical coordinates to the Cartesian coordinates by means of the formula

$$x = r\cos\varphi\sin\vartheta, \quad y = r\sin\varphi\sin\vartheta, \quad z = r\cos\vartheta.$$
 (1.10)

The formula (1.10) defines the smooth bijection from the open set W to the open set $V \subset \mathbb{R}^3$, where $V = \mathbb{R}^3 \setminus V'$ and

$$V' = \{ (x, y, z) \in \mathbb{R}^3 : x \ge 0, y = z = 0 \lor x = y = 0, -\infty < z < \infty \}.$$

Just as in the case of cylindrical coordinates one can find a formula for the transition from Cartesian to spherical coordinates and show that it induces the inverse smooth bijection from V to W. We leave all this as an exercise.

1.3 ALGEBRA OF SMOOTH FUNCTIONS ON EUCLIDEAN SPACE

Let U be an open subset of Euclidean space E^n , where we do not exclude the possibility that this set may coincide with the whole space E^n . Let $f: U \to \mathbb{R}$ be a real-valued function, where U is a domain of this function. In what follows we will assume (unless otherwise stated) that the Euclidean space E^n is equipped with a Euclidean frame $\{O; \vec{e_1}, \vec{e_2}, \ldots, \vec{e_n}\}$, which determines the Cartesian coordinates in E^n . A function f depends on a point x of an open set U, but we can identify a point xwith its coordinates x^1, x^2, \ldots, x^n , and then a function f can be considered as a function of n real variables $f(x^1, x^2, \ldots, x^n)$. In order to simplify notations we will often write f(x), implying that a point x is identified with its coordinates x^1, x^2, \ldots, x^n .

Smooth functions play an important role in differential geometry of Euclidean space E^n . A function $f: U \to \mathbb{R}$ is called a smooth function if it has continuous derivatives of all orders. The set of all smooth functions over an open subset U will be denoted by $C^{\infty}(U)$. What operations can we perform with smooth functions? Adding the values of two smooth functions $f, g \in C^{\infty}(U)$ at every point $x \in U$, we get the new function f + q, which is called the sum of functions f, q. Thus the value of the function f + q at a point x is f(x) + q(x) and this is called the pointwise addition of smooth functions. It is easy to show that the sum f + q is also a smooth function, i.e. $f + q \in C^{\infty}(U)$. Hence the set of smooth functions is closed under the pointwise addition of smooth functions. Similarly, multiplying the value of a smooth function f at each point of U by a real number a, we get the function af and this operation is called the multiplication of functions by real numbers. Analogously, we can define the pointwise product fg of two smooth functions by the formula (fg)(x) = f(x)g(x). It is also easy to show that af and fg are smooth functions, i.e. the set of smooth functions is closed under the pointwise multiplication of functions as well as the multiplication of smooth functions by real numbers.

What algebraic structure in the set of smooth functions is determined by the described above operations? If we consider only pointwise addition of smooth functions, then the set of all smooth functions is an additive Abelian group. It is easy to show that the pointwise addition of functions together with the multiplication of functions by real numbers yield the structure of vector space over the field of real numbers \mathbb{R} . Hence $C^{\infty}(U)$ is a vector space. If we add the pointwise multiplication of functions to these two algebraic operations, then the vector space $C^{\infty}(U)$ of smooth functions becomes an associative unital commutative algebra. An associative unital algebra is a vector space \mathscr{A} (with an addition of elements u, v denoted by u+v and a multiplication of elements by real numbers denoted by a u, where a is a real number) equipped with a mapping $(u, v) \in \mathscr{A} \times \mathscr{A} \to u \cdot v \in \mathscr{A}$, which is

- 1. bilinear, i.e. $(au + bv) \cdot w = au \cdot w + bv \cdot w$, $w \cdot (au + bv) = aw \cdot u + bw \cdot v$, where u, v, w are elements of a vector space \mathscr{A} and a, b are real numbers,
- 2. associative, i.e. $u \cdot (v \cdot w) = (u \cdot v) \cdot w$,
- 3. there exists in \mathscr{A} the element e such that $u \cdot e = e \cdot u = u$ for any element $u \in \mathscr{A}$.

A mapping $(u, v) \in \mathscr{A} \times \mathscr{A} \to u \cdot v \in \mathscr{A}$ is called a multiplication of algebra \mathscr{A} and element *e* is called an *identity element* of \mathscr{A} . In what follows we will use the term algebra to mean an associative unital algebra. If the product of any two elements of algebra does not depend on the order in which we multiply these elements, i.e. $u \cdot v = v \cdot u$, then the algebra is called *commutative*.

It is easy to see that the pointwise multiplication of smooth functions is bilinear with respect to the pointwise addition of smooth functions and the pointwise multiplication of smooth functions by real numbers. It is also associative and commutative. The constant function 1, whose value at any point $x \in U$ is 1, i.e. 1(x) = 1, is the identity element for pointwise multiplication of smooth functions. Thus the set of all smooth functions $C^{\infty}(U)$ over an open subset U of Euclidean space E^n is a commutative algebra.

If we consider only two operations with functions, pointwise addition and multiplication, then the set of all smooth functions $C^{\infty}(U)$ becomes a commutative associative ring with the unit element and this structure is often used in literature.

An orthonormal frame $\mathfrak{E} = \{O; \vec{e_1}, \vec{e_2}, \ldots, \vec{e_n}\}$ determines the Cartesian coordinate system in the affine Euclidean space E^n . This means that each point x has the coordinates x^1, x^2, \ldots, x^n . These coordinates are real numbers and they are uniquely determined by a point x and two different points have different coordinates. For the purposes of differential geometry, it is useful to look at coordinates as functions. For these functions, we will use the same notations, which we use for coordinates, i.e. these functions will be denoted by x^1, x^2, \ldots, x^n . This will not lead to confusion, because every time we specify, we consider x^i as a coordinate or as a function. If we consider x^i as a function, then its value at a point x is the *i*th coordinate of x, that is, $x^i(x) = x^i$. Obviously, every function x^i is a smooth function, because we have

$$\frac{\partial x^i}{\partial x^j} = \delta^i_j$$

The functions x^1, x^2, \ldots, x^n are called *Cartesian coordinate functions*.

1.4 PARAMETRIZED CURVES IN EUCLIDEAN SPACE

In this section we assume that *n*-dimensional Euclidean space E^n is equipped with an orthonormal frame $\mathfrak{E} = \{O; \vec{e_1}, \vec{e_2}, \ldots, \vec{e_n}\}$, which determines the Cartesian coordinates x^1, x^2, \ldots, x^n . A parametrized curve in E^n is a mapping

$$\xi: I \to E^n, \ \xi: t \in I \mapsto (x^1(t), x^2(t), \dots, x^n(t)) \in E^n,$$

where t is a parameter and $I \subset \mathbb{R}$ is an interval, which may be either finite or infinite. We will denote a parametrized curve by $(\xi = \xi(t), I)$ or simply by ξ . A parametrized curve can also be written in the form

$$\xi(t) = (x^{1}(t), x^{2}(t), \dots, x^{n}(t)), \qquad (1.11)$$

which will be called a *parametric equation of a parametrized curve* ξ . We can assign to each point of the affine space E^n its position vector. Thus, if we consider the position

vector $\vec{\xi}(t)$ of a point $\xi(t)$ of a parametrized curve ξ then we get the vector-valued function $\vec{\xi} = \vec{\xi}(t)$, which attributes to a real number $t \in I$ the vector $\vec{\xi}(t)$. Hence we can consider a parametric equation of a parametrized curve as a vector-valued function

$$\vec{\xi} = \vec{\xi}(t) = (x^1(t), x^2(t), \dots, x^n(t)).$$

The domain of $\vec{\xi} = \vec{\xi}(t)$ is an interval I and $\vec{\xi}(t)$ takes values in the vector space V^n . This vector-valued function $\vec{\xi}(t)$ is useful, because in this case we can the operations with vectors and the affine structure of E^n . For instance a mapping $\xi : \mathbb{R} \to E^n$ defined by the parametric equation

$$\xi(t) = p + t \,\vec{v},\tag{1.12}$$

where p is a point and \vec{v} is non-zero vector, is a straight line passing through a point p and \vec{v} is its directional vector. The vector-valued function of this equation $\vec{\xi}(t) = \vec{p} + t \vec{v}$, where \vec{p} is the position vector of a point p. The coordinate form of the equation (1.12) is

$$\xi(t) = (p^1 + v^1 t, p^2 + v^2 t, \dots, p^n + v^n t), \qquad (1.13)$$

where p^1, p^2, \ldots, p^n are coordinates of a point p and $\vec{v} = (v^1, v^2, \ldots, v^n)$ are coordinates of a vector \vec{v} .

If every function $x^i(t)$ in (1.11) is a smooth function, then a parametrized curve ξ will be called a *smooth parametrized curve*. In what follows, we will always assume the smoothness of a curve and the word "smooth" will be omitted. The set of points of Euclidean space E^n determined by a mapping ξ , when a parameter t runs through all the values of an interval I, is called the *image of a parametrized curve* ξ and the image of a parametrized curve ξ and the image of a parametrized curve ξ will be denoted by Im ξ . In what follows we will often refer to the image Im ξ of a parametrized curve (ξ, I) simply as a curve and a mapping $\xi : I \to E^n$ as its parametrization.

The basic notion of the differential geometry of curves is the notion of a tangent line to a curve. Let $p = \xi(t_0)$, $q = \xi(t_0 + \Delta t)$, where $t_0, t_0 + \Delta t \in I$, be two points of a parametrized curve ξ . The straight line that passes through these two points p, q is called a *secant* of a parametrized curve ξ . If Δt tends to zero, i.e. point q approaches to a point q along a parametrized curve ξ , then limiting position of a secant is called the *tangent line* to a parametrized curve at a point p. The tangent line at a point p is a linear approximation of a parametrized curve ξ at this point. In order to find this linear approximation, we can expand a vector-valued function $\xi = \xi(t)$ as power series in t at a point t_0 . We get

$$\vec{\xi}(t) = \vec{\xi}(t_0) + \vec{\xi}'(t_0)\,\Delta t + \mathcal{O}(\Delta t),\tag{1.14}$$

where

$$\vec{\xi'}(t_0) = \frac{d\vec{\xi}(t)}{dt}\Big|_{t=t_0} = \left(\frac{dx^1}{dt}\Big|_{t=t_0}, \frac{dx^2}{dt}\Big|_{t=t_0}, \dots, \frac{dx^n}{dt}\Big|_{t=t_0}\right),$$

is the derivative of a vector-valued function and $\mathcal{O}(\Delta t)$ are the higher order terms. From (1.14) it follows that the parametric equation of a tangent line can be written in the form $\xi(t_0) + \tau \, \vec{\xi'}(t_0)$, where τ is the parameter of a tangent line. Comparing this equation with (1.12) we conclude that the vector $\vec{\xi'}(t_0)$ can be taken as a directional vector of tangent line. A vector $(\xi(t_0); \vec{\xi'}(t_0))$ at a point $p = \xi(t_0)$ of a parametrized curve is called a *velocity vector* of a parametrized curve. The velocity vector of a parametrized curve ξ at a point $p = \xi(t_0)$ will be denoted by $\xi'(t_0)$. Hence

$$\xi'(t_0) = \left(p; \frac{dx^1}{dt}\Big|_{t=t_0}, \frac{dx^2}{dt}\Big|_{t=t_0}, \dots, \frac{dx^n}{dt}\Big|_{t=t_0}\right).$$
(1.15)

Thus according to this definition $\xi'(t_0) \in T_p E^n$.

Similarly the acceleration vector $\xi''(t_0)$ of a parametrized curve ξ at a point $p = \xi(t_0)$ is defined with the help of second order derivatives, that is,

$$\xi''(t_0) = \left(p; \frac{d^2 x^1}{dt^2}\Big|_{t=t_0}, \frac{d^2 x^2}{dt^2}\Big|_{t=t_0}, \dots, \frac{d^2 x^n}{dt^2}\Big|_{t=t_0}\right).$$
(1.16)

A point $p = \xi(t_0)$ of a parametrized curve $(\xi = \xi(t), I)$ is called a *regular point* if the velocity vector $\xi'(t_0)$ at this point is a non-zero vector, that is, $\|\xi'(t_0)\| \neq 0$. A parametrized curve $(\xi = \xi(t), I)$ is called a *regular parametrized curve* if each point of this parametrized curve is a regular point. Since we will study only regular parametrized curves the condition of regularity will be tacitly assumed and the word "regular" will be usually omitted.

A point $p = \xi(t_0)$ of a parametrized curve ξ is called a *point of rectification* if the acceleration vector $\xi''(t_0)$ and the velocity vector $\xi'(t_0)$ at this point are collinear vectors, i.e. $\xi'(t_0)||\xi''(t_0)$. In what follows we will exclude from the consideration the points of rectification of a parametrized curve, that is, we will assume that in addition to the condition of regularity of a parametrized curve $\xi'(t) \neq 0$ we have also the condition of non-collinearity of the velocity vector and the acceleration vector $\xi'(t_0) \not\parallel \xi''(t_0)$.

Naturally, if we consider the image of a parametrized curve $(\xi = \xi(t), I)$, then the set of points Im ξ can be parametrized in many different ways. By other words, we can have two different parametrized curves $(\xi = \xi(t), I)$ and $(\eta = \eta(\tau), J)$ with the same images, i.e. Im $\xi = \text{Im } \eta$. Indeed, consider a smooth function $t = t(\tau)$, whose domain is an interval $J = (c, d) \subset \mathbb{R}$ and a range is an interval I = (a, b). Assume that the derivative $t'(\tau)$ with respect to τ is non-zero at each point $\tau \in J$. Clearly, $t = t(\tau)$ is either strictly increasing $(t'(\tau) > 0)$ or strictly decreasing function $(t'(\tau) < 0)$. The product $\xi \circ t : J \to E^n$, where t is considered as the mapping $t : \tau \in J \mapsto t(\tau) \in I$ determined by a function $t = t(\tau)$, is the parametrized curve $(\eta = \eta(\tau), J)$, where $\eta = \xi \circ t$. Then Im $\xi = \text{Im } \eta$. The parametrized curve $(\eta = \eta(\tau), J)$ is called a *reparametrization* of a parametrized curve $(\xi = \xi(t), I)$ by means of a function $t = t(\tau)$.

The coordinates of a velocity vector depend on a parametrization of a curve. Therefore it is natural to raise a question of how the coordinates of a velocity vector transform under a reparametrization of a curve. Let (ξ, I) and (η, J) be two different parametrizations of a curve, where $\eta = \xi \circ t$ and $t = t(\tau)$ is a function $t : J \to I$, whose derivative with respect τ is non-zero at each point $\tau \in J$. Let

$$\xi(t) = (x^1(t), x^2(t), \dots, x^n(t)), \quad \eta(\tau) = (y^1(\tau), y^2(\tau), \dots, y^n(\tau)),$$



Figure 1.4 Function $t = t(\tau)$

where $t \in I$, $\tau \in J$, be the parametric equations of parametrized curves ξ and η . Then $y^i(\tau) = x^i(t(\tau))$. Making use of the rule of differentiation of a composite function we get

$$\frac{d}{d\tau}(y^{i}(\tau)) = \frac{d}{d\tau}(t(\tau)) \frac{d}{dt}(x^{i}(t))\big|_{t=t(\tau)}$$

Thus if we pass from one parametrization ξ to another η with the help of a function $t = t(\tau)$ then a velocity vector will transform according to a formula

$$\eta'(\tau) = t'(\tau) \,\xi'(t)\big|_{t=t(\tau)}, \ t'(\tau) = \frac{d}{d\tau}(t(\tau)), \tag{1.17}$$

which is called a chain rule for a parametrized curve. Differentiating both parts of the formula (1.17) one more time we obtain the formula for a transformation of acceleration vector when a curve $\xi(t)$ is reparametrized $\eta(\tau) = \xi(t(\tau))$

$$\eta''(\tau) = \left(t'(\tau)\right)^2 \xi''(t)\big|_{t=t(\tau)} + t''(\tau) \xi'(t)\big|_{t=t(\tau)}.$$
(1.18)

Hence we proved the following statement.

Proposition 1.2. Let $(\xi = \xi(t), I)$ be a (regular) parametrized curve and $(\eta = \eta(\tau), J)$ be its reparametrization by means of a function $t = t(\tau), t : J \to I$. Then a velocity vector and an acceleration vector transform as follows

$$\begin{aligned} \eta'(\tau) &= t'(\tau) \,\xi'(t) \big|_{t=t(\tau)}, \\ \eta''(\tau) &= (t'(\tau))^2 \,\xi''(t) \big|_{t=t(\tau)} + t''(\tau) \,\xi'(t) \big|_{t=t(\tau)}. \end{aligned}$$

From this proposition it follows that the condition of regularity of a parametrized curve is invariant under a reparametrization. Indeed the first formula in Proposition (1.2) shows that the reparametrized curve η is regular because of $\xi'(t) \neq 0$ and

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 $t'(\tau) \neq 0$. It also follows from Proposition (1.2) that the condition $\xi'(t_0) \not\models \xi''(t_0)$ (which excludes from consideration the points of rectification) does not depend on a parametrization of a curve. It follows from the first formula of Proposition (1.2) that the vector $\eta'(\tau)$ is collinear with the vector $\xi'(t)$. Next it is easy to see that the vector $(t'(\tau))^2 \xi''(t)$ is a non-zero vector. Indeed we have $t'(\tau \neq 0)$. The acceleration vector $\xi''(t)$ can not be zero vector because this would contradict with the assumption $\xi'(t_0) \not\models \xi''(t_0)$. Now the second formula of Proposition (1.2) shows that acceleration vector $\eta''(\tau)$ is obtained by adding the non-zero vector $(t'(\tau))^2 \xi''(t)$ to vector, which is collinear to $\eta'(\tau)$. Thus the resulting vector $\eta''(\tau)$ cannot be collinear with vector $\eta'(\tau)$.

Let $(\xi = \xi(t), I)$ be a parametrized curve and $t_0, t_1 \in I$, where $t_0 < t_1$. Then the length of an arc of a parametrized curve ξ between two points t_0, t_1 is given by the integral

$$s = \int_{t_0}^{t_1} \|\xi'(t)\| dt.$$
(1.19)

The integrand is an infinitesimal element of an arc length and is often denoted by $ds = \|\xi'(t)\| dt$. We will also use this notation and call ds the arc length differential. By virtue of the fact, mentioned above, that the image of a parametrized curve does not change upon reparametrization, it can be proved that the length of an arc of a parametrized curve is invariant with respect to reparametrization (Exercise 3.2).

A parametrized curve $(\xi = \xi(t), I)$ is called a *unit-speed parametrized curve* if for any point $t \in I$ we have $\|\xi'(t)\| = 1$, i.e. the velocity vector at each point of a parametrized curve is a unit vector.

Proposition 1.3. If $(\xi = \xi(s), I)$ is a unit-speed parametrized curve, then

$$<\xi''(s),\xi'(s)>=0$$

Proof. We have $\|\xi'(s)\| = 1$ or $\langle \xi'(s), \xi'(s) \rangle = 1$. Differentiating the left-hand side of this relation with respect to s, we obtain

$$(\|\xi'(s)\|^2)' = \langle \xi''(s), \xi'(s) \rangle + \langle \xi'(s), \xi''(s) \rangle = 2 \langle \xi''(s), \xi'(s) \rangle.$$

The derivative of the right-hand side is equal to zero and we get

$$<\xi''(s),\xi'(s)>=0.$$

Theorem 1.1. For any parametrized curve $(\xi = \xi(t), I)$ there exists a unit-speed parametrization.

Proof. Fix a point $t_0 \in I$ and define the function

$$s(t) = \int_{t_0}^t \|\xi'(u)\| \, du, \quad t \in I.$$
(1.20)

The domain of this function is an interval I. Let us denote the range of the function s = s(t) by J. Then $s : t \in I \mapsto s(t) \in J$. It follows from the theory of functions that s = s(t) is a smooth function. We can calculate the derivative of s = s(t) by means of the general formula

$$F'(t) = \int_{g(t)}^{f(t)} F'_t(u,t) \, du + f'(t) \, F(f(t),t) - g'(t) \, F(g(t),t), \tag{1.21}$$

where

$$F(t) = \int_{g(t)}^{f(t)} F(u,t) \, du,$$

and F(u, t), f(t), g(t) are smooth functions. In the case of the function s = s(t) the first and last terms in (1.21) vanish because the integrand and the lower limit of integration in (1.20) do not depend on the parameter t. Thus

$$s'(t) = \frac{ds}{dt} = \frac{dt}{dt} \|\xi'(t)\| = \|\xi'(t)\|.$$
(1.22)

Since we consider only regular parametrized curves for any $t \in I$ it holds $s'(t) \neq 0$. Hence there exists uniquely determined inverse function for s = s(t), which is also smooth. The inverse function will be denoted by t = t(s) and $t : s \in J \mapsto t(s) \in I$. The derivative of the inverse function can be expressed in terms of the derivative of the function s(t) as follows

$$t'(s) = \frac{1}{s'(t)}\Big|_{t=t(s)}.$$
(1.23)

Now we can construct a reparametrization of a parametrized curve $(\xi = \xi(t), I)$ with the help of the inverse function t = t(s). Define the parametrized curve $(\eta = \eta(s), J)$, where $\eta(s) = \xi(t(s))$. Making use of the chain rule for parametrized curves (3.1) and the formula (1.22) for the derivative of the function s = s(t), we find

$$\|\eta'(s)\| = t'(s) \left(\|\xi'(t)\|\right)\Big|_{t=t(s)} = \frac{1}{s'(t)}\Big|_{t=t(s)} \left(\|\xi'(t)\|\right)\Big|_{t=t(s)} = 1.$$

Thus we see that the parametrized curve $(\eta = \eta(s), J)$ has the unit-speed parametrization and this ends the proof.

From the proof of Theorem 1.1 it is clear that in order to construct a unit-speed parametrization of a parametrized curve one uses the length of arc as a parameter. This parameter is called *natural parameter* and unit-speed parametrization is called *natural parametrization*.

Example 1.4. Let us show, using a catenary as an example, how a unit-speed parametrization is constructed. We assume that a plane E^2 is equipped with an orthonormal frame that defines a Cartesian coordinate system, whose coordinates we will denote by x, y. The curve, that is called catenary, is a plane curve and usually given by the equation

$$y = a \cosh\left(\frac{x}{a}\right),\tag{1.24}$$

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where a > 0 is a real number. This curve is of mechanical origin. If we have a sufficiently heavy, flexible, inextensible cable that sags under the action of gravity, being fixed at the ends at the same level, then the shape of this cable is the catenary. Equation (1.24) is easy to write in a parametric form. To do this, we can choose



Figure 1.5 Catenary

the variable x as a parameter t. Then a catenary can be written in the form of a parametrized curve as follows

$$\xi(t) = (t, a \cosh\left(\frac{t}{a}\right)), \ I = \mathbb{R}.$$

Differentiating with respect to t, we find the velocity vector and its length

$$\xi'(t) = (\xi(t); 1, \sinh{(\frac{t}{a})}), \ ||\xi'(t)|| = \sqrt{1 + \sinh^2{(\frac{t}{a})}} = \cosh{(\frac{t}{a})}$$

From this we obtain the arc length function of the catenary

$$s(t) = \int_0^t \cosh\left(\frac{u}{a}\right) du = a \int_0^t \cosh\left(\frac{u}{a}\right) d\left(\frac{u}{a}\right) = a \sinh\left(\frac{u}{a}\right) \Big|_0^t = a \sinh\left(\frac{t}{a}\right).$$

Note that the lower limit of integration in this formula shows that we have chosen the vertex of the catenary as the reference point for the arc length, that is, the point of its intersection with axis y. Now we have to find the inverse function for $s = a \sinh(\frac{t}{a})$. We have

$$t = a \sinh^{-1}(\frac{s}{a}) = a \ln \frac{s + \sqrt{s^2 + a^2}}{a}$$

Here \sinh^{-1} stands for the inverse function of \sinh and we used the formula $\sinh^{-1} x = \ln(x + \sqrt{x^2 + 1})$. Thus, we obtain a natural parametrization of the catenary

$$\beta(s) = (a \sinh^{-1}(\frac{s}{a}), a \cosh(\sinh^{-1}(\frac{s}{a}))).$$
(1.25)

The second function of this parametrization can be simplified by means of $\cosh(\sinh^{-1} x) = \sqrt{1 + x^2}$, the application of which gives the final result

$$\beta(s) = (a \sinh^{-1}\left(\frac{s}{a}\right), \sqrt{s^2 + a^2}).$$

Example 1.5. In this example, consider the natural parametrization of a helix. A helix is a spatial curve that is defined by the parametric equation $(\xi = \xi(t), I)$, where

$$\xi(t) = (a \, \cos t, a \, \sin t, bt), \ I = \mathbb{R}$$

and a > 0, b > 0 are constant real numbers. First of all, let us calculate the arc length function of a helix. We have

$$\begin{aligned} \xi'(t) &= (\xi(t); -a \sin t, a \cos t, b), \\ ||\xi'(t)|| &= \sqrt{a^2 + b^2}, \\ s(t) &= \int_0^t \sqrt{a^2 + b^2} \, du = \sqrt{a^2 + b^2} \, t \end{aligned}$$

In this case, the inverse function is very easy to find, in the last formula we should simply express t through s. We have

$$t = \frac{s}{\sqrt{a^2 + b^2}}.$$

Hence the unit-speed parametrization of a helix is

$$\eta(s) = (a \cos \frac{s}{\sqrt{a^2 + b^2}}, a \sin \frac{s}{\sqrt{a^2 + b^2}}, \frac{b s}{\sqrt{a^2 + b^2}}).$$
(1.26)

1.5 VECTOR FIELDS ALONG A PARAMETRIZED CURVE

A concept of a vector field plays an important role in modern differential geometry. Vector fields play an important role in applications of differential geometry in various fields of physics and mechanics. For example, the distribution of velocities of flow of a fluid at some fixed moment of time can be described by means of a geometric notion of a vector field.

Let U be an open subset of n-dimensional Euclidean space E^n . Then a vector field \mathbf{X} in U is a mapping $\mathbf{X}: U \to T(U)$ such that $\pi \circ \mathbf{X} = \mathrm{id}_U$, where $\pi: T(U) \to U$ is the projection and id_U is the identity mapping of U. Thus a vector field is a mapping which assigns to each point p of U a vector at a point p. This vector at a point p will be denoted by \mathbf{X}_p . A subset U will be called a domain of a vector field \mathbf{X} . It is evident that a vector field \mathbf{X} is uniquely determined by n real-valued functions X^1, X^2, \ldots, X^n defined on U, that is, $X^i: U \to \mathbb{R}$, where $\mathbf{X}_p = (p; X^1(p), X^2(p), \ldots, X^n(p))$, and $X^i(p)$ is a value of an *i*th function at a point p. A function X^i will be referred to as an *i*th component of a vector field \mathbf{X} . If there is a coordinate system in an open subset U with coordinates x^1, x^2, \ldots, x^n the the components of a vector field \mathbf{X} are functions of coordinates $X^i(x^1, x^2, \ldots, x^n)$, that is, functions of n real variables. In what follows we will assume that all components X^1, X^2, \ldots, X^n are smooth functions, that is, $X^1, X^2, \ldots, X^n \in C^\infty(U)$. By other words, we will consider only smooth vector fields omitting the word "smooth". Obviously a vector field \mathbf{X} is completely determined by the following vector-valued function

$$\overrightarrow{\mathbf{X}} = (X^1, X^2, \dots, X^n) : U \to \mathbb{R}^n, \quad \overrightarrow{\mathbf{X}}_p = (X^1(p), X^2(p), \dots, X^n(p)).$$

Example 1.6. Let $f \in C^{\infty}(U)$ be a smooth function. Define a vector field grad f by the following formula

$$(\operatorname{grad} f)_p = \left(p; \frac{\partial f}{\partial x^1}\Big|_p, \frac{\partial f}{\partial x^2}\Big|_p, \dots, \frac{\partial f}{\partial x^n}\Big|_p\right),$$

where $p \in U$. The vector field ∇f is called the gradient of a function f. Obviously the gradient of a smooth function is a smooth vector field. The corresponding vectorvalued function of the gradient of a function will be denoted $\overrightarrow{\nabla} f$, i.e.

$$\overrightarrow{grad} f = \left(\frac{\partial f}{\partial x^1}, \frac{\partial f}{\partial x^2}, \dots, \frac{\partial f}{\partial x^n}\right)$$

We can perform the following operations with vectors: addition, multiplication by real numbers, inner product and in three-dimensional space also a cross product. We can extend in obvious way (pointwise) all these operations to vector fields, i.e. for example we can define the sum of two vector fields $\mathbf{X} + \mathbf{Y}$ as the vector field, whose value at a point $p \in U$ is the vector $\mathbf{X}_p + \mathbf{Y}_p$. Analogously we can define a vector field \mathbf{X} multiplied by a real number a and the inner product of two vector fields $< \mathbf{X}, \mathbf{Y} >$, where $< \mathbf{X}, \mathbf{Y} >$ is the function. It is easy to show that if \mathbf{X}, \mathbf{Y} are smooth vector fields, then $\mathbf{X} + \mathbf{Y}, a \mathbf{X}$ are also smooth vector fields and $< \mathbf{X}, \mathbf{Y} >$ is the smooth function. Let us denote the set of all smooth vector fields on an open set $U \subset E^n$ by $\mathcal{D}(U)$. It is clear that the operation of addition of two vector fields and their multiplication by real numbers turn the set $\mathcal{D}(U)$ into the vector space over \mathbb{R} and it should be mentioned that this vector space is infinite dimensional.

The operation of multiplying vector fields by real numbers can be slightly changed, assuming that a number depends on a point of open set U, i.e. it is a function defined on U. In this case we get a new operation of multiplying vector fields by smooth functions. Given a smooth function $f \in C^{\infty}(U)$ and a vector field \mathbf{X} , we define the vector field $f \mathbf{X}$, whose value at a point p is the vector $f(p) \mathbf{X}_p$. It is easy to show that this operation has the following algebraic properties:

- 1. $f(\mathbf{X} + \mathbf{X}) = f\mathbf{X} + f\mathbf{Y},$
- $2. \ (f+g) \mathbf{X} = f \mathbf{X} + g \mathbf{X},$
- 3. $(fg) \mathbf{X} = f(g\mathbf{X}),$
- 4. 1X = X.

These properties show that multiplication of vector fields by smooth functions determines on the Abelian group of vector fields $\mathcal{D}(U)$ the structure of module over the ring of smooth functions $C^{\infty}(U)$.

The concept of a vector field is very useful in the theory of curves. A vector field along a curve is a very natural concept and is analogous to a vector field in Euclidean space. A vector field along a parametrized curve ($\xi = \xi(t), I$) is a mapping, which

assigns to each point of a parametrized curve a vector at this point. A vector field along a parametrized curve $(\xi = \xi(t), I)$ will be denoted by X(t). Hence,

$$\mathbf{X}(t) = (\xi(t); X^{1}(t), X^{2}(t), \dots, X^{n}(t)),$$

where $X^{i}(t)$ is a real-valued function of one variable t, whose domain is $I \subset \mathbb{R}$. In what follows we will assume that for any i = 1, 2, ..., n a function $X^{i}(t)$ is smooth... The set of all smooth vector fields along a parametrized curve ($\xi = \xi(t), I$) will be



Figure 1.6 Vector field along a curve

denoted by D_{ξ} .

Thus, if at each point of a parametrized curve we have a vector that smoothly depends on a point of parametrized curve, then the set of all these vectors is the vector field along a parametrized curve. For example, at each point of a parametrized curve, we defined the velocity vector and the set of all velocity vectors determines the vector field, which will be called a *tangent vector field* or *vector field of velocity* of a parametrized curve. The velocity vector field will be denoted by $\xi'(t)$. Analogously, we can define an *acceleration vector field* of a parametrized curve with the help of acceleration vector. The acceleration vector field of a parametrized curve will be denoted by $\xi''(t)$.

One more important operation with functions, which can be extended to vector fields, is a differentiation. This operation is very important and will lead us to a concept of covariant differentiation of a vector field. We extend differentiation from functions to vector fields as follows: If $\mathbf{X}(t)$ is a vector field along a parametrized curve $\xi = \xi(t)$, then its derivative with respect to t is a new vector field $\mathbf{X}'(t)$, whose component functions are the derivatives of component functions of $\mathbf{X}(t)$ with respect to t. Thus,

$$\mathbf{X}'(t) = (\xi(t); (X^1(t))', (X^2(t))', \dots, (X^n(t))'),$$

where $\mathbf{X}(t) = (\xi(t); X^1(t), X^2(t), \dots, X^n(t))$. From this definition it can be seen that differentiation of a vector field $\mathbf{X}(t)$ essentially reduces to differentiation of its vector-valued function $\mathbf{\vec{X}}(t)$, i.e. $\mathbf{X}'(t) = (\xi(t); \mathbf{\vec{X}}'(t))$. It follows from this definition that

differentiation is a linear operation on vector fields, i.e. $(a \mathbf{X} + \mathbf{Y})' = a \mathbf{X}'(t) + \mathbf{Y}'(t)$, where a is a real number. It is a simple exercise to prove the following properties of differentiation of vector fields

$$(f(t) X(t))' = f'(t) X(t) + f(t) X'(t), \quad (< X(t), Y(t) >)' = < X'(t), Y(t) > + < X(t), Y'(t) > + < X(t), Y(t) > + < X(t), Y'(t) > + < X(t), Y(t) > + < X(t), Y'(t) > + < X(t), Y(t) > + < X(t), Y(t) > + < X(t), Y'(t) > + < X(t)$$

Proposition 1.4. Let X(t) be a vector field along a parametrized curve $(\xi = \xi(t), I)$ of constant length, i.e. for any $t \in I$ it holds ||X(t)|| = c, where $c \ge 0$ is a real number. Then

 $<\mathbf{X}(t),\mathbf{X}'(t)>=0,$

i.e. the derivative of a constant length vector field X(t) is orthogonal to X(t).

Proof. For any $t \in I$ we have $\langle \mathbf{X}(t), \mathbf{X}(t) \rangle = c^2$. Differentiating the both sides of this relation and making use of the property of differentiation of inner product of two vector fields, we obtain

$$(<\mathbf{X}(t),\mathbf{X}(t)>)'=<\mathbf{X}'(t),\mathbf{X}(t)>+< X(t),\mathbf{X}'(t)>=2<\mathbf{X}(t),\mathbf{X}'(t)>=0,$$

and it follows immediately from this that the inner product of X(t) and X'(t) is zero. We note that this proposition is a generalization of Proposition 1.3.

1.6 CURVATURE OF A PARAMETRIZED CURVE

A curvature of a parametrized curve at some point of this curve can be measured by means of its deviation from a tangent line at this point. Let us consider a unit-speed parametrized curve ($\eta = \eta(s), I$). Assuming that η is parametrized by a natural parameter s we do not lose a generality of a structure considered below since in Section (1.4) it is proved that any (regular) parametrized curve can be parametrized with the help of a natural parameter (Theorem 1.1). Hence a velocity vector $\eta'(s)$ is a unit vector at any point of η , i.e. $\|\eta'(s)\| = 1$. Let p be some point of a parametrized curve η corresponding to a value $s_0 \in I$ of natural parameter s, i.e. $p = \eta(s_0)$. Let us give the value s_0 of natural parameter s some increment of Δs so that $s_0, s_0 + \Delta s \in I$ and denote the point of a parametrized curve η corresponding to $s_0 + \Delta s$ by q, i.e. $q = \eta(s_0 + \Delta s)$. Then point q can be considered as the displacement of the point palong a parametrized curve η . The velocity vectors are vectors at a point and in order to have their vector part we can represent them as follows

$$\eta'(s_0) = (p; \vec{\eta}'(s_0)), \eta'(s_0 + \Delta s) = (q; \vec{\eta}'(s_0 + \Delta s)).$$

Let us denote the angle between vectors $\vec{\eta}'(s_0)$ and $\vec{\eta}'(s_0 + \Delta s)$ by $\Delta \phi$. Note that velocity vector η' changes when the point p moves along the curve η to the point qand it happens because η is curved. However, due to the fact that we use the natural parametrization, the velocity vector remains unit vector, that is, there is no change in its length. This means that all the change in the velocity vector η' occurs due to the change in its direction. The figure (1.7) clearly shows that the direction of a velocity vector changes the faster, the more curved a line. Thus, we can measure how curved a line η is using the rate of change of the angle $\Delta \phi$ when the point moves along a line.

Definition 1.3. The curvature $\kappa(s)$ of a parametrized curve $\eta = \eta(s)$ at a point $s_0 \in I$ is the limit

$$\kappa(s_0) = \lim_{\Delta s \to 0} \frac{\Delta \phi}{\Delta s}.$$

If vectors $\vec{\eta}'(s_0), \vec{\eta}'(s_0 + \Delta s)$ are applied from a point p, then the end points of these vectors lie on the unit circle with the center at a point p. Let us denote $\Delta \eta'(s_0) = (p; \vec{\eta}'(s_0 + \Delta s) - \vec{\eta}'(s_0))$. Then it holds

$$\lim_{\Delta s \to 0} \frac{\Delta \phi}{\|\Delta \eta'(s_0)\|} = 1.$$

Indeed, we have $\|\Delta \eta'(s_0)\| = 2 \sin \frac{\Delta \phi}{2}$. Thus

$$\lim_{\Delta s \to 0} \frac{\Delta \phi}{\|\Delta \eta'(s_0)\|} = \lim_{\Delta \phi \to 0} \frac{\Delta \phi}{2 \sin \frac{\Delta \phi}{2}} = \lim_{\Delta \phi \to 0} \frac{\Delta \phi/2}{\sin(\Delta \phi/2)} = 1$$

Now, we can express curvature as follows

$$\kappa(s_0) = \lim_{\Delta s \to 0} \frac{\Delta \phi}{\Delta s} = \lim_{\Delta s \to 0} \frac{\Delta \phi}{\|\Delta \eta'(s_0)\|} \frac{\|\Delta \eta'(s_0)\|}{\Delta s}$$
$$= \lim_{\Delta s \to 0} \frac{\Delta \phi}{\|\Delta \eta'(s_0)\|} \lim_{\Delta s \to 0} \frac{\|\Delta \eta'(s_0)\|}{\Delta s} = \lim_{\Delta s \to 0} \frac{\|\Delta \eta'(s_0)\|}{\Delta s} = \|\eta''(s)\|\Big|_{s=s_0}$$

Thus we proved that the curvature $\kappa(s)$ of a unit-speed parametrized curve $\eta = \eta(s)$ at a point s is equal to the length of acceleration vector at this point.



Figure 1.7 Curvature of a line

1.7 PLANE CURVES AND THEIR CURVATURES

In this section we consider the curvature of a parametrized curve in the case of an affine Euclidean plane E^2 . Throughout this section we assume that the Euclidean

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plane E^2 is endowed with an orthonormal right-handed oriented frame $\{O; \vec{e_1}, \vec{e_2}\}$ and the Cartesian coordinates induced by this frame will be denoted by x, y. Our aim in this section is to define a signed curvature of a plane curve and to derive formula for its calculation in Cartesian coordinates.

Before we move on to considering the curvature of a plane curve we will consider ways to define a curve on a plane. The first and basic way to define a curve is parametric, that is, a curve is given by a parametric equation $\xi(t) = (x(t), y(t))$, where $t \in I$. However there are other ways to define a curve on a plane E^2 . Let y = f(x) be a smooth function and $I \subset \mathbb{R}$ be its domain. Then the graph of a function f, i.e. the set of points

graph
$$f = \{(x, f(x) \in E^2 : x \in I)\},$$
 (1.27)

can be considered as a curve on a plane E^2 . This way to define a plane curve can be considered as a particular case of a parametrized curve. Indeed taking the first coordinate x as a parameter t we can define a curve (1.27) by means of the parametric equation $\xi(t) = (t, f(t)), t \in I$. On the other hand if we have a parametrized curve $\xi(t) = (x(t), y(t))$ then we can present this curve at least locally as a graph of a function. Let $p = \xi(t_0)$ be a point of a curve ξ . Since ξ is a regular curve at least one of the derivatives $x'(t_0), y'(t_0)$ at a point p has to be non-zero. Assume $x'(t_0) \neq 0$. Then there exists a neighbourhood of t_0 where t can be expressed by means of the x(t) function's inverse as t = t(x). Then the graph of the function y = y(t(x)) will locally (in a neighbourhood of t_0) coincide with a curve ξ . Thus we see that the way to define a plane curve by means of a parametric equation is locally equivalent to the way to define a curve as a graph of a smooth function.

Let a plane curve is defined as a graph of a smooth function $f : I \to \mathbb{R}$. Let $x_0 \in I$. In order to find the equation of a tangent line to a curve y - f(x) we expand a function f(x) in a Taylor series in a neighbourhood of a point $x = x_0$

$$y = y_0 + f'(x_0)(x - x_0) + \mathcal{O}(\Delta x),$$

where $y_0 = f(x_0)$, $\Delta x = x - x_0$ and $\mathcal{O}(\Delta x)$ are terms of second and higher order. Omitting all terms of second and higher order we get the equation of a tangent line to a curve y = f(x) at a point $x = x_0$

$$y - y_0 = f'(x_0)(x - x_0).$$
(1.28)

One can write (1.28) in the form of a parametric equation $x = x_0 + t$, $y = y_0 + f'(x_0)t$ which shows that $\mathbf{v} = (p; 1, f'(x_0))$ can be taken as a velocity vector of a curve y = f(x) at a point $p = (x_0, y_0)$.

A plane curve can be also defined with the help of an equation F(x, y) = c, where F(x, y) is a smooth function of two variables and c is a real number. In this case a curve is the set of all points of a plane E^2 whose coordinates satisfy an equation F(x, y) = c. Let us denote this set of solutions by \mathfrak{C} . Additionally we will assume that the set of solutions of an equation F(x, y) = c is non-empty, i.e. $\mathfrak{C} \neq \emptyset$, and at each point of a curve we have $(F'_x)^2 + (F'_y)^2 > 0$. Let $p = (x_0, y_0)$ be a point of a curve F(x, y) = c such that $F'_y|_p \neq 0$. Then according to the implicit function theorem

there exists a neighbourhood of a point p, where y can be expressed as a smooth function of x, i.e. y = f(x), and

$$\frac{df}{dx} = -\frac{F'_x}{F'_y}.$$

Thus the set of solutions of equation F(x, y) = 0 is a curve on a plane, which is locally, that is, in a neighbourhood of each point, can be represented as a graph of a function.

Assume that a plane curve is defined by an equation F(x, y) = c and according to the implicit function theorem in a neighbourhood of a point $p = (x_0, y_0)$ of this curve we have y = f(f), where $y_0 = f(x_0)$ and $F(x, f(x)) \equiv c$. Then

$$dF(x,y) = F'_x(p) \, dx + F'_y(p) \, dy = 0 \implies F'_x(p) + F'_y(p) \, f'(x_0) = 0$$

The left-hand side of this equation can be considered as a scalar product of two plane vectors $(p; F'_x(p), F'_y(p))$ and $(p; 1, f'(x_0))$. But the latter vector is a velocity vector of a curve at a point p. Hence the vector $(p; F'_x(p), F'_y(p))$ is perpendicular to the velocity vector of a curve and can be taken as a normal vector of a curve F(x, y) = 0 at a point p. A straight line which passes a point of a curve and is perpendicular to a tangent line to curve at this point is called a normal line of a curve. From previous considerations it follows that equation of the normal line of a curve F(x, y) = c at a point $p = (x_0, y_0)$ of this curve is

$$F'_{x}(p)(x - x_{0}) + F'_{y}(p)(y - y_{0}) = 0.$$
(1.29)

It is worth to mention that a function F(x, y) standing at the left-hand side of an equation of a curve induces the gradient vector field ∇F in a domain of a function F(x, y). The restriction of this vector field to a curve \mathfrak{C} gives the vector field along this curve and this vector field is perpendicular to the velocity vector field of a curve. Hence $\nabla F|_{\mathfrak{C}}$ is a normal vector field along a curve F(x, y) = c.

The curvature of a plane curve can be defined with the help of Definition 1.3, when n = 2. However, in the case of a plane curve, we can use the right-handed orientation of a basis to endow the curvature of a parametrized curve with a sign. This can be made by means of a complex structure of a plane E^2 . Let p be a point of a plane. We rotate this point around the origin O counterclock-wise by right angle, get the point q and this determines the transformation of a plane E^2 , which will be denoted by J. Hence, $J : E^2 \to E^2$ and J(p) = q. This transformation is called the *complex structure of a plane* E^2 . The transformation J can be applied to vectors of the associated vector space V^2 . Indeed, given a vector $\vec{v} \in V^2$, we apply it from the origin, then rotate around the origin counterclock-wise by right angle and denote the obtained vector by $J(\vec{v})$. It is easy to see that $J : V^2 \to V^2$ is a linear transformation, i.e. we have $J(a\vec{v} + b\vec{w}) = a J(\vec{v}) + b J(\vec{w})$. This linear mapping has the following properties:

1.
$$J^2 = -\mathrm{id}_{E^2}$$

2.
$$< J(\vec{v}), J(\vec{w}) > = < \vec{v}, \vec{w} >;$$

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3. $< J(\vec{v}), \vec{v} >= 0.$

Evidently, if we apply the transformation J to the vectors of basis, then $J(\vec{e}_1) = \vec{e}_2, J(\vec{e}_2) = -\vec{e}_1$. Hence, if $\vec{v} = v^1 \vec{e}_1 + v^2 \vec{e}_2$, then

$$J(\vec{v}) = J(v^1 \vec{e}_1 + v^2 \vec{e}_2) = v^1 J(\vec{e}_1) + v^2 J(\vec{e}_2) = -v^2 \vec{e}_1 + v^1 \vec{e}_2.$$

Thus, the transformation J sends a vector with coordinates v^1, v^2 to the vector with coordinates $-v^2, v^1$. This implies that J maps a point of E^2 with coordinates x, y to the point with coordinates -y, x, i.e. J(x, y) = (-y, x).



Figure 1.8 Complex structure of a plane

Consider a unit-speed parametrized plane curve $\eta = \eta(s)$. According to Proposition 1.3, the acceleration vector $\eta''(s)$ is perpendicular to the velocity vector $\eta'(s)$. Hence, the acceleration vector and the vector $J(\eta'(s))$ are collinear.

Definition 1.4. Let $\eta = \eta(s)$ be a unit-speed parametrized plane curve. Then the signed curvature $\kappa(s)$ of a unit-speed parametrized plane curve ξ is defined by

$$\eta''(s) = \kappa(s) J(\eta'(s)).$$

From this definition we easily get the formula for the curvature of a unit-speed parametrized plane curve

$$\kappa(s) = <\eta''(s), J(\eta'(s)) > .$$
 (1.30)

Taking into account that the acceleration vector $\eta''(s)$ and $J(\eta'(s))$ are collinear vectors and $||J(\eta'(s))|| = 1$, we conclude $|\kappa(s)| = ||\eta''(s)||$ and this shows that Definition 1.4 is consistent with the previously given general Definition 1.3. Hence, in the case of a parametrized plane curve the absolute value of the curvature is equal to the length of acceleration vector. This is the reason why the curvature of a parametrized plane curve is called *signed curvature*.

The calculation of curvature in a unit-speed parametrization is often a difficult task, because first we need to find this unit-speed parametrization, which requires calculation of inverse function for arc length function. Thus, it is desirable to have a formula for curvature in an arbitrary parametrization. In order to find this formula, we begin with a parametrized curve ($\xi = \xi(t), I$), which is not necessarily a unitspeed parametrized curve. But, according to Theorem 1.1, there exists a unit-speed parametrization of this curve, which will be denoted by $\eta = \eta(s)$. We know that this unit-speed parametrization is constructed by means of the arc length function s = s(t) of ξ , i.e. $\xi(t) = \eta(s(t))$. Thus, it is natural to define the curvature $\kappa(t)$ of a parametrized curve $\xi = \xi(t)$ by putting $\kappa(t) = \kappa(s)|_{s=s(t)}$, where $\kappa(s)$ is the curvature of the unit-speed parametrized curve $\eta(s)$. From this definition and the formula (1.30) we obtain

$$\kappa(t) = \langle \eta''(s), J(\eta'(s)) \rangle \Big|_{s=s(t)}.$$
(1.31)

Differentiating the formula $\xi(t) = \eta(s(t))$ and using the chain rule for parametrized curves, we get

$$\xi'(t) = s'(t) \eta'(s)|_{s=s(t)} \implies \eta'(s)|_{s=s(t)} = \frac{\xi'(t)}{\|\xi'(t)\|}.$$
 (1.32)

Differentiating the above formula one more time, we get

$$s'(t) \eta''(s)|_{s=s(t)} = \frac{\|\xi'(t)\| \xi''(t) - (\|\xi'(t)\|)' \xi'(t))}{\|\xi'(t)\|^2}.$$

Substituting $s'(t) = \|\xi'(t)\|$ we can express the accelerator vector $\eta''(s)$ as follows

$$\eta''(s)|_{s=s(t)} = \frac{1}{\|\xi'(t)\|^2} \xi''(t) - \frac{(\|\xi'(t)\|)'}{\|\xi'(t)\|^3} \xi'(t).$$
(1.33)

Making use of (1.32) and linearity of complex structure J we get

$$J(\eta'(s))|_{s=s(t)} = J\left(\frac{\xi'(t)}{\|\xi'(t)\|}\right) = \frac{J(\xi'(t))}{\|\xi'(t)\|}.$$

Substituting this formula and (1.33) into (1.31) and taking into account that $\xi'(t)$ is perpendicular to $J(\xi'(t))$, i.e. $\langle \xi'(t), J(\xi'(t)) \rangle = 0$, we finally get

$$\kappa(t) = \frac{\langle \xi ''(t), J(\xi '(t)) \rangle}{\|\xi '(t)\|^3}.$$
(1.34)

We can write the above formula for the curvature of a parametrized curve $\xi(t)$ in Cartesian coordinates of the plane E^2 . Let $\xi(t) = (x(t), y(t))$, then

$$\xi'(t) = (\xi(t); x'(t), y'(t)), \quad \xi''(t) = (\xi(t); x''(t), y''(t)),$$

$$J(\xi'(t)) = (\xi(t); -y'(t), x'(t)), \quad \|\xi'(t)\| = \sqrt{(x'(t))^2 + (y'(t))^2}.$$

Substituting these formulae into (1.34) we obtain

$$\kappa(t) = \frac{x'(t) y''(t) - x''(t) y'(t)}{\left[(x'(t))^2 + (y'(t))^2 \right]^{3/2}},$$
(1.35)

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$$\kappa(t) = \frac{1}{\left[(x'(t))^2 + (y'(t))^2 \right]^{3/2}} \begin{vmatrix} x'(t) & y'(t) \\ x''(t) & x''(t) \end{vmatrix}$$

The resulting formula (1.35) is applicable in the case of an arbitrary parametrization of a curve and we do not need previously to find a natural parametrization of a curve.

Let us consider some particular cases of formula (1.35). If a plane curve is defined as a graph of a function y = f(x) then we can parametrize this curve as follows $\xi(x) = (x, f(x))$. Hence $\xi'(x) = (\xi(x); 1, f'(x)), \xi''(x) = (\xi(x); 0, f''(x))$ and

$$\kappa(x) = \frac{f''(x)}{[1 + (f'(x))^2]^{3/2}}.$$
(1.36)

If a plane curve is defined in an implicit form by an equation F(x, y) = c we can use previous formula (1.36) assuming that locally an equation F(x, y) = c is resolved and we have the explicit function y = f(x) satisfying $F(x, f(x)) \equiv c$. Then

$$f'(x) = -\frac{F'_x}{F'_y}.$$
 (1.37)

In order to express the second order derivative f''(x) in terms of partial derivatives of a function F(x, y) we differentiate the relation $F'_x + F'_y f'(x) = 0$, which immediately follows from $dF(x, y) = F'_x dx + F'_y dy = 0$ when y = f(x). Analogously we find

$$\begin{aligned} d(F'_x + F'_y f'(x)) &= dF'_x + (dF'_y) f'(x) + F'_y df'(x) \\ &= F''_{xx} \, dx + F''_{xy} \, dy + (F''_{xy} \, dx + F''_{yy} \, dy) \, f'(x) + F'_y \, f''(x) \, dx. \end{aligned}$$

Taking into account that y is a function of x, i.e. y - f(x), we obtain

$$F'_{y} f''(x) + F''_{xx} + 2 F''_{xy} f'(x) + F''_{yy} dy) (f'(x))^{2} = 0.$$

Substituting from (1.37) and expressing f''(x) in terms of the partial derivatives we get

$$f'' = -\frac{F''_{xx}(F'_y)^2 - 2F''_{xy}F'_xF'_y + F''_{yy}(F'_x)^2}{(F'_y)^3}.$$
(1.38)

Substituting (1.37) and (1.38) into (1.36) we finally find the formula for a curvature of a curve defined by F(x, y) = c

$$\kappa = -\frac{F_{xx}''(F_y')^2 - 2F_{xy}'F_x'F_y' + F_{yy}''(F_x')^2}{[(F_x')^2 + (F_y')^2]^{3/2}}.$$
(1.39)

Suppose we are given a parametrized curve $\xi(t) = (x(t), y(t)), t \in I$ in a plane E^2 . The curvature of this curve can be used to construct a new curve. This new curve is constructed based on the following geometric considerations. Generally the tangent to a curve is a straight line and it has first-order tangency. Thus the tangent to a curve gives us a linear approximation of this curve at some of its points. This means that having expanded the functions x(t), y(t) in a Taylor series we discard

terms of the second order and higher. Let us now pose the question of finding a line such that the tangency would be of the second order. Thus this line would give a quadratic approximation of a curve. In order to find a quadratic approximation of a curve (tangency of the second order) we will use a circle. Let $t_0 \in I$ and $p = \xi(t_0) =$ be a point of a curve ξ . We assume that the curvature $\kappa(t_0)$ of a curve ξ is non-zero at a point p. A circle that passes through a point p and has a second-order tangency with a curve ξ at a point p is called an *osculating circle*. Our aim now is to find a radius and a center of osculating circle at a point p. Let us denote a center of an osculating circle by p_0 , its position vector by \vec{r} and a radius of osculating circle by R. Then we have $\|\vec{\xi}(t_0) - \vec{r}\| = R^2$, where $\vec{\xi}(t_0)$ is a position vector of a point p. Let $\xi(t)$ be a position vector of a point q of a curve ξ close enough to a point p. Let us connect point q with a center of an osculating circle p_0 by a straight line segment and denote the point of intersection of this line segment with an osculating circle by \tilde{q} . The deviation of a curve ξ from osculating circle is equal to the length of the segment $q\tilde{q}$ and can be written as $\|q - \tilde{q}\| = \|q - p_0\| - \|\tilde{q} - p_0\|$. This deviation must be infinitesimal of the third order with respect to $t - t_0$. It is easy to see that $||q-p_0|| - ||\tilde{q}-p_0||$ and $||q-p_0||^2 - ||\tilde{q}-p_0||^2$ are infinitesimals of the same order. Indeed we have

$$\lim_{t \to t_0} \frac{\|q - p_0\|^2 - \|\tilde{q} - p_0\|^2}{\|q - p_0\| - \|\tilde{q} - p_0\|} = \lim_{t \to t_0} (\|q - p_0\| + \|\tilde{q} - p_0\|) = 2R,$$

because $\|\tilde{q}-p_0\|$ is a radius of an osculating circle and $\lim_{t\to t_0} \|q-p_0\| = \|p-p_0\| = R$. Thus we can find a center of an osculating circle \vec{r} and its radius R from the condition that the expansion in the Taylor series of the function

$$f(t) = \|\vec{\xi}(t) - \vec{r}\|^2 - R^2$$

begins with terms of the third order, that is, in the expansion

$$f(t) = f(t_0) + f'(t_0)(t - t_0) + \frac{1}{2}f''(t_0)(t - t_0)^2 + \frac{1}{6}f'''(t_0)(t - t_0)^3 + \dots,$$

we must have $f(t_0) = f'(t_0) = f''(t_0) = 0$. These three conditions are necessary and sufficient for a circle centered at point p and with radius R to have a second-order tangency to a curve ξ , that is, to be an osculating circle. The figure shows that vector $p\vec{p}_0$ is perpendicular to the tangent vector $\xi'(t_0)$. It follows that by turning the unit vector

$$\vec{e} = \frac{\xi'(t_0)}{\|\xi'(t_0)\|}$$

at a right angle counterclockwise with the help of the complex structure J, we can write the position vector of a center of osculating circle \vec{r} in the form

$$\vec{r} = \vec{\xi}(t_0) + R \cdot J(\vec{e}).$$
 (1.40)

Let us show that this vector satisfies the conditions $f(t_0) = f'(t_0) = 0$. We have

$$f(t_0) = \|\vec{\xi}(t_0) - \vec{\xi}(t_0) - R \cdot J(\vec{e})\|^2 - R^2 = R^2 - R^2 = 0.$$
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In the case of the first order derivative of the function f(t) we have

$$f'(t_0) = 2 < \vec{\xi}(t_0) - \vec{r}, \vec{\xi'}(t_0) > = -2\frac{R}{\|\vec{\xi'}(t_0)\|} < J(\vec{\xi'}(t_0)), \vec{\xi'}(t_0) > = 0,$$

because of $J(\vec{\xi'}(t_0)) \perp \vec{\xi'}(t_0)$. The second order derivative of the function f(t) at a point $t = t_0$ can be written as follows

$$f''(t_0) = 2\left(\|\vec{\xi}'(t_0)\|^2 + \langle \vec{\xi}(t_0) - \vec{r}, \vec{\xi}''(t_0) \rangle\right).$$

Substituting (1.40) into the resulting expression and equating it to zero, we obtain

$$\|\vec{\xi'}(t_0)\|^2 - \frac{R}{\|\vec{\xi'}(t_0)\|} < \vec{\xi''}(t_0), J(\vec{\xi'}(t_0)) >= 0.$$

Making use of the formula for curvature (1.35) we find the radius of an osculating circle

$$R = \frac{1}{\kappa(t_0)}.\tag{1.41}$$

Due to the relation (1.41) the radius of an osculating curve is also called a *radius of* curvature of a curve ξ at a point $\xi(t_0)$. The center of an osculating circle is called a center of curvature of a curve ξ at a point $\xi(t_0)$. Substituting (1.41) into (1.40) we find the position vector of a center of curvature

$$\vec{r} = \vec{\xi}(t_0) + \frac{1}{\kappa(t_0)} \cdot J(\frac{\vec{\xi}'(t_0)}{\|\vec{\xi}'(t_0)\|}).$$
(1.42)

Using the affine structure of plane E^2 we can pass from the position vectors $\vec{r}, \vec{\xi}(t_0)$ to the corresponding points p_0, p and write the center of curvature as a point of a plane

$$p_0 = p + \frac{1}{\kappa(t_0)} \cdot J(\frac{\bar{\xi'}(t_0)}{\|\bar{\xi'}(t_0)\|}).$$
(1.43)

Now note that an osculating circle is defined at each point of a curve $\xi = \xi(t)$, that is, for any value of the parameter $t \in I$. Thus we can consider the set of points in a plane E^2 that are the centers of the osculating circles of this curve. Obviously this set of points is a plane curve which is called *evolute* and the parametric equation of the evolute of a curve $\xi = \xi(t)$ follows from (1.42) or (1.43) and has the form

$$\mathbf{c}(t) = \vec{\xi}(t) + \frac{1}{\kappa(t)} \cdot J(\frac{\vec{\xi'}(t)}{\|\vec{\xi'}(t)\|})$$
(1.44)

1.8 PLANE CURVES IN POLAR COORDINATES

So far we have used Cartesian coordinates in Euclidean affine space E^n , which are constructed by means of an orthonormal frame $\mathfrak{E} = \{O; \vec{e_1}, \vec{e_2}, \ldots, \vec{e_n}\}$. However, in Euclidean space E^n , besides Cartesian coordinates, there are another coordinate

systems, which are called curvilinear coordinates. We will consider a general approach to this type of coordinates in subsequent chapters of this book, but now it is useful to consider the special case of such coordinates on a plane E^2 .

A parameterized curve ξ in polar coordinates of a plane can be defined in a similar way to how it is done in Cartesian coordinates. This means that one can define a parameterized curve as a mapping $\xi : I \to E^2$, which in polar coordinates r, ϕ is determined by a parametric equation $\xi(t) = (r(t), \phi(t)), t \in I$, where $r(t), \phi(t)$ are smooth functions. However, it is usually convenient to choose the polar angle ϕ as a parameter. In this case the parametric equation of a parametrized curve has a simple form $r = r(\phi)$ and we will determine a parametrized curve with the help of this equation. If a parametrized curve is given in polar coordinates of a plane by equation $r = r(\phi)$, we can easily derive its parametric equation in Cartesian coordinates, if we construct an orthonormal frame $\mathfrak{E} = \{O; \vec{e_1}, \vec{e_2}\}$ as follows: Origin O coincides with the pole, $\vec{e_1}$ is the unit directional vector of the polar axis (pointing in the positive direction of the polar axis) and $\vec{e_2}$ is obtained by counterclockwise rotation of $\vec{e_1}$ around the origin O by the right angle. Then so constructed Cartesian coordinates can be expressed in terms of polar coordinates as follows

$$x = r \cos \phi, \quad y = r \sin \phi. \tag{1.45}$$

Thus, the parametric equation of a curve $r = r(\phi)$ in Cartesian coordinates is

$$\xi(\phi) = (r(\phi)\cos\phi, r(\phi)\sin\phi), \qquad (1.46)$$

and the parameter is a polar angle ϕ . Hence, we can apply previously developed formulae to calculate the coordinates of velocity vector, acceleration vector, the arc length and the curvature of a curve $r = r(\phi)$.

In the case of polar coordinates, we can construct a basis for the tangent space $T_p E^2$ of a plane at a point p similar to the basis $\{e_{p,1}, e_{p,2}\}$ in Cartesian coordinates. But vectors of this basis will depend on a point of plane. In order to construct a basis for $T_p E^2$ in polar coordinates, we use the coordinate lines. Let p be a point of plane E^2 , whose polar coordinates are r, ϕ . If we fix the polar angle ϕ of a point p and let polar radius r to vary within some interval $(r - \delta, r + \delta)$, then we get the first coordinate line passing through a point p and the parametric equation of this straight line is (t, ϕ) , where $r - \delta < t < r + \delta$. Analogously, the second coordinate line is the circle, whose radius is r and center is the pole O. The parametric equation of the second coordinate line is (r, τ) , where $\phi - \delta < \tau < \phi + \delta$. We will use the velocity vectors of these coordinates lines to construct the orthonormal basis for the tangent space $T_p E^2$. In order to find the coordinates of velocity vectors, we pass from polar coordinates to Cartesian ones, because we know how to calculate the coordinates of velocity vectors only in Cartesian coordinates. The parametric equation of the first coordinate line in Cartesian coordinates is $(t \cos \phi, t \sin \phi)$ and of the second line is $(r\cos\tau, r\sin\tau)$. Thus, the first velocity vector is $(p;\cos\phi,\sin\phi)$ and the second one is $(p; -r\sin\phi, r\cos\phi)$. It is easy to see that the first vector is a unit vector, the length of the second vector is r and vectors are orthogonal to each other. Hence, the vectors

$$\begin{aligned} \mathbf{e}_{p,r} &= (p; \cos \phi, \sin \phi) = \cos \phi \, \mathbf{e}_{p,1} + \sin \phi \, \mathbf{e}_{p,2}, \\ \mathbf{e}_{p,\phi} &= (p; -\sin \phi, \cos \phi) = -\sin \phi \, \mathbf{e}_{p,1} + \cos \phi \, \mathbf{e}_{p,2}, \end{aligned}$$

form the orthonormal basis for tangent space $T_p E^2$ in polar coordinates. The above formulae can be written in the matrix form

$$\begin{pmatrix} \mathbf{e}_{p,r} & \mathbf{e}_{p,\phi} \end{pmatrix} = \begin{pmatrix} \mathbf{e}_{p,1} & \mathbf{e}_{p,2} \end{pmatrix} \begin{pmatrix} \cos\phi & -\sin\phi \\ \sin\phi & \cos\phi \end{pmatrix},$$

and we see that the orthonormal basis in polar coordinates $\mathbf{e}_{p,r}$, $\mathbf{e}_{p,\phi}$ can be obtained by counterclockwise rotation of the canonical orthonormal basis $\mathbf{e}_{p,1}$, $\mathbf{e}_{p,2}$ around a point p by an angle ϕ .

To pass from vectors to the vector fields induced by them, we will omit the point p in the corresponding notation. Hence $\mathbf{e}_1, \mathbf{e}_2$ will stand for the vector fields induced by the canonical basis \vec{e}_1, \vec{e}_2 and $\mathbf{e}_r, \mathbf{e}_{\phi}$ for the vector fields induced by $\mathbf{e}_{p,r}, \mathbf{e}_{p,\phi}$. Thus we have

$$\mathbf{e}_r = \cos\phi \,\mathbf{e}_1 + \sin\phi \,\mathbf{e}_2,\tag{1.47}$$

$$\mathbf{e}_{\phi} = -\sin\phi \,\mathbf{e}_1 + \cos\phi \,\mathbf{e}_2,\,,\qquad(1.48)$$

and

$$\begin{pmatrix} \mathbf{e}_r & \mathbf{e}_\phi \end{pmatrix} = \begin{pmatrix} \mathbf{e}_1 & \mathbf{e}_2 \end{pmatrix} \begin{pmatrix} \cos \phi & -\sin \phi \\ \sin \phi & \cos \phi \end{pmatrix}.$$

Now we can calculate the velocity vector field \mathbf{v} along a curve given in polar coordinates by equation $r = r(\phi)$, and we calculate the coordinates of this velocity vector field in the orthonormal frame field $\mathbf{e}_r, \mathbf{e}_{\phi}$ induced by the polar coordinates. In Cartesian coordinates defined by a canonical frame field $\mathbf{e}_1, \mathbf{e}_2$ this velocity vector field has the coordinates

$$r'\cos\phi - r\,\sin\phi, r'\sin\phi + r\,\cos\phi. \tag{1.49}$$

We can write them in a matrix form as follows

$$\left(\begin{array}{c}r'\cos\phi - r\sin\phi\\r'\sin\phi + r\cos\phi\end{array}\right) = \left(\begin{array}{c}\cos\phi & -\sin\phi\\\sin\phi & \cos\phi\end{array}\right) \left(\begin{array}{c}r'\\r\end{array}\right)$$

Since the matrix

$$\left(\begin{array}{cc}\cos\phi & -\sin\phi\\\sin\phi & \cos\phi\end{array}\right)$$

transforms the coordinates of the velocity vector field \mathbf{v} in frame field $\mathbf{e}_r, \mathbf{e}_{\phi}$ into coordinates in the frame field $\mathbf{e}_1, \mathbf{e}_2$, we conclude that r', r are the coordinates of the velocity vector field in frame field $\mathbf{e}_r, \mathbf{e}_{\phi}$. Hence

$$\mathbf{v} = r' \,\mathbf{e}_r + r \,\mathbf{e}_\phi. \tag{1.50}$$

Hence, $\|\mathbf{v}\| = \sqrt{(r')^2 + r^2}$ and the length of the arc between two points $(r(\phi_1), \phi_1)$, $(r(\phi_2), \phi_2)$ of a curve $r = r(\phi)$, where $\phi_1 < \phi_2$, can be calculated in polar coordinates with the help of the formula

$$s = \int_{\phi_1}^{\phi_2} \sqrt{(r')^2 + r^2} \, d\phi. \tag{1.51}$$

In order to calculate the curvature of a plane curve given in polar coordinates by equation $r = r(\phi)$ we need to calculate the coordinates of acceleration vector field **a** along a curve $r = r(\phi)$. We do this by differentiating the coordinates of the velocity vector field (1.49) with respect ϕ

$$(r''-r)\cos\phi - 2r'\sin\phi, (r''-r)\sin\phi + 2r'\cos\phi,$$

and writing it in the matrix form

$$\begin{pmatrix} (r''-r)\cos\phi - 2r'\sin\phi\\ (r''-r)\sin\phi + 2r'\cos\phi \end{pmatrix} = \begin{pmatrix} \cos\phi & -\sin\phi\\ \sin\phi & \cos\phi \end{pmatrix} \begin{pmatrix} r''-r\\ 2r' \end{pmatrix}.$$

Thus the acceleration vector field in polar coordinates can be written as

$$\mathbf{a} = (r'' - r) \,\mathbf{e}_r + 2 \,r' \,\mathbf{e}_\phi. \tag{1.52}$$

The coordinates of the acceleration vector **a** can be calculated in a different way. To do this, we differentiate both sides of equation (1.50) with respect to polar angle ϕ . We obtain

$$\mathbf{a} = \mathbf{v}' = r'' \,\mathbf{e}_r + r' \,\mathbf{e}_r' + r' \,\mathbf{e}_\phi + r \,\mathbf{e}_\phi'. \tag{1.53}$$

Now the question is what the derivatives of the vector fields \mathbf{e}_r and \mathbf{e}_{ϕ} are equal to. We consider these vector fields as vector fields along a curve $r = r(\phi)$ and differentiate them by the parameter of this curve, which is the polar angle of a point, that is, we differentiate along this curve. However, it is easy to see that these vector fields do not depend on the polar radius (see formulas (1.47,1.48). Thus, their change is completely determined by the movement of the point along the coordinate line ϕ , that is, the circle. From this it follows that the derivatives of these vector fields can be calculated by differentiating the formulae (1.47,1.48) with respect to variable ϕ . This differentiation gives

$$\mathbf{e}_r' = \mathbf{e}_\phi, \mathbf{e}_\phi' = -\mathbf{e}_r.$$

Substituting this result onto (1.48) we get (1.52).

Now applying complex structure transformation to the velocity vector field (at each point of a curve $r = r(\phi)$) we find

$$J(\mathbf{v}) = -r \,\mathbf{e}_r + r' \mathbf{e}_\phi.$$

Making use of the formula (1.34) for a curvature we finally get the curvature of a curve in polar coordinates

$$\kappa = \frac{\langle \mathbf{a}, J(\mathbf{v}) \rangle}{||\mathbf{v}||^{3/2}} = \frac{r^2 + 2(r')^2 - r r''}{[r^2 + (r')^2]^{3/2}}.$$
(1.54)

Example 1.7. One of the most famous examples of a curve defined by a parametric equation in polar coordinates is the logarithmic spiral. It is given by the equation $r = ae^{b\phi}$, where a, b are real numbers such that $a > 0, b \neq 0$. This curve has many wonderful properties. We will calculate the arc length of a logarithmic spiral between two polar angle values. We find

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Figure 1.9 Logarithmic spiral

$$r' = ab e^{b\phi}, \ (r')^2 + r^2 = a^2(b^2 + 1) e^{2b\phi}.$$

Then the length of arc of logarithmic spiral between two its points $r(\phi_1), r(\phi_2)$, where $\phi_1 < \phi_2$ can be calculated by means of the formula (1.51) as follows

$$s = \int_{\phi_1}^{\phi_2} \sqrt{a^2(b^2+1) e^{2b\phi}} \, d\phi = a\sqrt{b^2+1} \int_{\phi_1}^{\phi_2} e^{b\phi} d\phi$$
$$= a\frac{\sqrt{b^2+1}}{b} \int_{\phi_1}^{\phi_2} e^{b\phi} d(b\phi) = a\frac{\sqrt{b^2+1}}{b} e^{b\phi} \Big|_{\phi_1}^{\phi_2} = \frac{\sqrt{b^2+1}}{b} (r(\phi_2) - r(\phi_1)).$$

We calculate the curvature of the logarithmic spiral using formula (1.54). We have

$$r'' = ab^2 e^{b\phi}, \ r^2 + 2(r')^2 - r r'' = a^2(b^2 + 1) e^{2b\phi},$$

and

$$\kappa = \frac{r^2 + 2(r')^2 - r \, r''}{[r^2 + (r')^2]^{3/2}} = \frac{a^2(b^2 + 1) \, e^{2b\phi}}{a^3(b^2 + 1)^{3/2} \, e^{3b\phi}} = \frac{1}{r\sqrt{b^2 + 1}}$$

1.9 BARTELS-FRENET-SERRET EQUATIONS

In this subsection we consider parametrized curves in three dimensional Euclidean space E^3 , construct a moving frame along a parametrized curve and derive Bartels-Frenet-Serret formulas for a moving frame, which express the derivatives of vector fields constituting moving frame, in terms of these vector fields themselves. In the first part of this section, we will use the natural parametrization of a curve.

Let $(\eta = \eta(s), I)$ be a unit-speed parametrized curve in three dimensional Euclidean space E^3 and T be the vector field of velocity along a unit-speed parametrized

curve η . Hence, $T(s) = \eta'(s)$ and ||T(s)|| = 1. In this section we assume that the curvature $\kappa(s)$ of a unit-speed parametrized curve η is non-zero at any point of this curve, i.e. for any $s \in I$ we have $\kappa(s) \neq 0$. The derivative of T is the vector field of acceleration along η and $T' \perp T$, $||T'(s)|| = ||\eta''(s)|| = \kappa(s)$. Thus, we can make T' to be a unit vector field by normalizing it, i.e. we introduce the vector field

$$\mathbf{N}(s) = \frac{1}{\kappa(s)} \mathbf{T}'(s). \tag{1.55}$$

This vector field is orthogonal to the vector field of velocity T and ||N(s)|| = 1 for any $s \in I$. The vector field N will be referred to as a normal vector field along a parametrized curve η . In order to have a three-dimensional frame at each point of a curve η , we introduce one more vector field B along a curve η , which is called a *binormal vector field* and is defined as cross-product of vector field of velocity T and the normal vector field N. Hence $B = T \times N$. Three vector fields $\{T, N, B\}$ form an orthonormal frame for the tangent space $T_{n(s)}E^3$ at each point of a unit-speed parametrized curve $\eta(s)$ and this frame will be referred to as a moving frame of a parametrized curve. Now our aim is to find formulae, which show how the derivatives of the vector fields T', N', B', constituting the moving frame, can be expanded in terms of vector fields themselves. The geometric essence of these formulae is that they show the speed of rotation of the three-dimensional frame $\{T, N, B\}$ when a point moves along a curve. Indeed, all three vectors of this frame are mutually orthogonal unit vectors, and when a point moves along a curve, the lengths of these vectors and their relative position do not change. Geometrically, this means that when a point moves along a curve, these three vectors rotate as a single rigid body. This means that the derivatives of these vector fields will show the speed of this rotation.

It follows from the definition (1.55) of the normal vector field N that

$$\mathbf{T}' = \kappa \, \mathbf{N}.\tag{1.56}$$

Thus, we have expressed the derivative of the vector field of velocity in terms of vector fields of moving frame $\{T, N, B\}$. This equation is called the first Bartels-Frenet-Serret equation.

Now consider the derivative of binormal vector field B. Differentiating the cross-product $T\times N$ we get

$$B' = T' \times N + T \times N'. \tag{1.57}$$

The first term on the right-hand side of (1.57) is equal to the zero vector and this follows from the first Bartels-Frenet-Serret equation (1.56). The second term in (1.57) (and hence the vector field B') is perpendicular to vector fields T and N'. But we have one more vector field, which is also perpendicular to vector fields T and N', this is the normal vector field N. Therefore, the derivative of binormal vector field B' is collinear to the normal vector field N and there exists a uniquely defined scalar function $\tau(s)$ such that

$$\mathsf{B}' = -\tau \, \mathsf{N}.\tag{1.58}$$

This function τ is called a *torsion of the spatial curve* and it is an important geometric

characteristic of a curve, the geometric meaning of which we will explore later in this section. The equation (1.58) is called the third Bartels-Frenet-Serret equation.

Now we consider the derivative of the normal vector field N. As $\{T, N, B\}$ is an orthonormal frame, we can expand the derivative N' as follows

$$\mathbb{N}' = < \mathbb{N}', \mathbb{T} > \mathbb{T} + < \mathbb{N}', \mathbb{N} > \mathbb{N} + < \mathbb{N}', \mathbb{B} > \mathbb{B}.$$

As N is a constant length vector field, it follows from Proposition 1.4 that $\langle N, N' \rangle = 0$ and the middle term in the above expansion vanishes. In order to find the coefficient function $\langle N', T \rangle$ in the first term of expansion, we differentiate the both sides of relation $\langle N, T \rangle = 0$ and get $\langle N', T \rangle = -\langle N, T' \rangle$. Making use of the first Bartels-Frenet-Serret equation $T' = \kappa N$ we find $\langle N', T \rangle = -\kappa$. Thus, we have

$$\mathbf{N}' = -\kappa \,\mathbf{T} + \tau \,\mathbf{B},\tag{1.59}$$

and this equation is called the second Bartels-Frenet-Serret equation.

Summing up these considerations, we can formulate the obtained result as follows. If $(\eta = \eta(s), I)$ is a unit-speed parametrized curve, $\kappa(s)$ is its curvature, which satisfies $\kappa(s) \neq 0$ for any $s \in I$ then the moving frame {T, N, B} of this curve satisfies the Bartels-Frenet-Serret equation

$$\mathbf{T}' = \kappa \,\mathbf{N},\tag{1.60}$$

$$\mathbf{N}' = -\kappa \,\mathbf{T} + \tau \,\mathbf{B},\tag{1.61}$$

$$\mathbf{B}' = -\tau \,\mathbf{N},\tag{1.62}$$

where $\tau = \langle \mathbb{N}', \mathbb{B} \rangle$ is called a torsion of a unit-speed parametrized curve ($\eta = \eta(s), I$). The structure of Bartels-Frenet-Serret formulae becomes very clear if we write them in matrix form

$$\begin{pmatrix} \mathbf{T}' \\ \mathbf{N}' \\ \mathbf{B}' \end{pmatrix} = \begin{pmatrix} 0 & \kappa & 0 \\ -\tau & 0 & \kappa \\ 0 & -\tau & 0 \end{pmatrix} \begin{pmatrix} \mathbf{T} \\ \mathbf{N} \\ \mathbf{B} \end{pmatrix}.$$

We see that the third-order square matrix on the right-hand side of this matrix equation is skew-symmetric. This fact is very important and will be explained later.

It should be noted that the formulas (1.60), (1.61), (1.62) are a powerful tool in the study of spatial curves. Besides, these formulas have several important applications in theoretical physics and mechanics. It is worth to mention a historical aspect of the discovery of these formulas [1]. These formulas were published by F. Frenet in his dissertation in [2] in 1847 and, independently of him, these formulas were obtained by another French mathematician, J.A. Serret, and published in [5] in 1851. In this regard, these formulas are usually called the Frenet-Serret formulas. However, it emerged later that the formulas (1.60), (1.61), (1.62) were discovered by M. Bartels before the French mathematicians F. Frenet and J.A. Serret ever published them. M. Bartels was a professor of mathematics at the University of Dorpat (now Tartu) in Estonia. He used to present his results of studies in differential geometry of curves in lectures for students. One of his students, C.E. Senff, took notes of M. Bartels

lectures and later included these notes into his book "Principal theorems from the theory of curves and surfaces", which was published in 1831. Therefore, from a historical point of view, it would be correct to call the formulas (1.60), (1.61), (1.62) Bartels-Frenet-Serret formulas, as it was proposed in [3].

Example 1.8. As an example, let us find Bartels-Frenet-Serret equations for a helix. Recall that in Example (1.5), a natural (unit-speed) parametrization of a helix was found and we will use it in this example. The unit-speed parametrization of a helix is given by

$$\eta(s) = (a \, \cos \frac{s}{\sqrt{a^2 + b^2}}, a \, \sin \frac{s}{\sqrt{a^2 + b^2}}, \frac{b \, s}{\sqrt{a^2 + b^2}}).$$

To simplify the form of formulae and make them more compact, we introduce the notation $\alpha = \sqrt{a^2 + b^2}$. Hence differentiating with respect to natural parameter s we get the velocity vector field

$$\mathbf{T}(s) = \eta'(s) = (\eta(s); -\frac{a}{\alpha}\sin\frac{s}{\alpha}, \frac{a}{\alpha}\cos\frac{s}{\alpha}, \frac{b}{\alpha}).$$

Differentiating the velocity vector field T we find the acceleration vector field

$$\mathbf{T}'(s) = \eta''(s) = (\eta(s); -\frac{a}{\alpha^2}\cos\frac{s}{\alpha}, -\frac{a}{\alpha^2}\sin\frac{s}{\alpha}, 0).$$

Since the curvature of a spatial curve in unit-speed parametrization is equal to the length of the acceleration vector, we find the curvature of a helix

$$\kappa(s) = ||\eta''(s)|| = \frac{a}{\alpha^2} = \frac{a}{a^2 + b^2}$$

Note that the curvature of a helix is the constant non-zero number. Next we find the normal vector field of a helix

$$\mathbf{N}(s) = \frac{1}{\kappa(s)}\mathbf{T}'(s) = (\eta(s); -\cos\frac{s}{\alpha}, -\sin\frac{s}{\alpha}, 0).$$

The cross product $T \times N$ gives us the binormal vector field B. Hence

$$\mathsf{B}(s) = (\eta(s); \frac{b}{\alpha} \sin \frac{s}{\alpha}, -\frac{b}{\alpha} \cos \frac{s}{\alpha}, \frac{a}{\alpha}).$$

Differentiating the binormal vector field we obtain

$$\mathsf{B}'(s) = (\eta(s); \frac{b}{\alpha^2} \cos \frac{s}{\alpha}, \frac{b}{\alpha^2} \sin \frac{s}{\alpha}, 0)$$

The last formula shows that the derivative of the binormal vector field B' is collinear to the normal vector field N and the corresponding coefficient function, when we express B' in terms of N, taken with a minus, is equal to the torsion $\tau(s)$ of a helix. Hence

$$\tau(s) = \frac{b}{\alpha^2} = \frac{b}{a^2 + b^2}.$$

We leave the verification of the second Bartels-Frenet-Serret equation to reader as a small exercise.

We have derived Bartels-Frenet-Serret formulae based on the assumption that a parametrized curve is regular and it is parameterized by the natural parameter s. Since such a parametrization is always possible, the obtained result is general for the entire class of regular curves. However, we can find a form of Bartels-Frenet-Serret equations in the case of an arbitrary parametrization of a regular curve. To extend these equations to an arbitrary parametrization of a regular curve, we assume that we are given a regular parameterized curve ($\xi, \xi = \xi(t), I$). Let ($\eta = \eta(s), J$) be the unit-speed parametrization of ξ . Then we know that $\xi(t) = \eta(s(t))$, where s(t) is the arc length function of ξ , i.e.

$$s(t) = \int_{t_0}^t ||\xi'(u)|| \, du,$$

where $t_0 \in I$. Thus, using the arc length function s(t), one can switch from a unitspeed parametrization $\eta(s)$ to a given parametrization $\xi(t)$. Hence it follows that we can extend any quantity (function, vector field) defined for a given curve in the natural parametrization and depending on the natural parameter s to an arbitrary parametrization by replacing the natural parameter s with the arc length function s = s(t). According to this approach, we define the moving frame, curvature and torsion of a curve ($\xi = \xi(t), I$) (i.e. in parametrization $\xi(t)$) as follows

$$\begin{split} \mathbf{T}(t) &= \mathbf{T}(s)\big|_{s=s(t)}, \ \mathbf{N}(t) = \mathbf{N}(s)\big|_{s=s(t)}, \ \mathbf{B}(t) = \mathbf{B}(s)\big|_{s=s(t)}, \\ \kappa(t) &= \kappa(s)\big|_{s=s(t)}, \ \tau(t) = \tau(s)\big|_{s=s(t)}. \end{split}$$

It is important to note here that neither the values of the vector fields of the moving frame, nor the values of curvature and torsion at the points of the curve, change, only the value of the parameter that corresponds to them changes.

Now our goal is to find the form of Bartels-Frenet-Serret equations in the case of an arbitrary parametrization $\xi(t)$. We have

$$\mathbf{T}'(t) = s'(t) \, \mathbf{T}'(s) \big|_{s=s(t)} = ||\xi'(t)|| \left(\kappa(s) \, \mathbf{N}(s)\right) \Big|_{s=s(t)} = ||\xi'(t)|| \, \kappa(t) \, \mathbf{N}(t).$$

When deriving this equation, we used the chain rule for parametrized curves and the first Bartels-Frenet-Serret equation in the case of unit-speed parametrization. Passing to vector fields and functions (i.e., omitting parameter t in the above equation) and denoting the function $s'(t) = ||\xi'||$ by v, we get

$$\mathbf{T}' = \upsilon \,\kappa \,\mathbf{N}.\tag{1.63}$$

Similarly one can derive the second and third Bartels-Frenet-Serret equations for a parametrized curve ($\xi = \xi(t), I$). Thus, we got the following result: If ($\xi = \xi(t), I$) is a regular parametrized curve and the curvature of this curve is non-zero at each point of a curve, then the moving frame {T, N, B} satisfies the equations

$$\mathbf{T}' = \boldsymbol{\upsilon} \boldsymbol{\kappa} \, \mathbf{N}, \tag{1.64}$$

$$\mathbf{N}' = -\upsilon\kappa\,\mathbf{T} + \upsilon\tau\,\mathbf{B},\tag{1.65}$$

$$\mathbf{B}' = -\upsilon\tau\,\mathbf{N},\tag{1.66}$$

which will be referred to as generalized Bartels-Frenet-Serret equations.

Our next goal is to derive formulae for the moving frame, curvature and torsion of a parametrized curve ($\xi = \xi(t), I$). To do this, we need the derivatives of $\xi(t)$ up to the third order inclusive. Differentiating the both sides of $\xi(t) = \eta(s(t))$ with respect to t we get

$$\xi'(t) = s' \eta'(s)\big|_{s=s(t)}.$$

Taking into account that $\eta'(s) = T(s)$ and the definition of T(t), we can put the previous formula into the form $\xi' = v T$. Differentiating one more time and applying the first generalized Bartels-Frenet-Serret equation (1.64) we obtain

$$\xi'' = \upsilon' \operatorname{T} + \upsilon T' = \upsilon' \operatorname{T} + \upsilon^2 \kappa \operatorname{N}.$$

Analogously, i.e. differentiating the both sides of the above formula and making use of Bartels-Frenet-Serret equations, we calculate the third order derivative of $\xi(t)$. Hence we have

$$\xi' = \upsilon \mathsf{T}, \tag{1.67}$$

$$\xi'' = \upsilon' \mathsf{T} + \upsilon^2 \kappa \mathsf{N}, \tag{1.68}$$

$$\xi''' = (v'' - v^3 \kappa^2) \mathbf{T} + (3v v' \kappa + v^2 \kappa') \mathbf{N} + v^3 \kappa \tau \mathbf{B}.$$
(1.69)

Using these equations, we can calculate the curvature and torsion of a curve. First of all, we consider the vector product of the vector fields ξ' and ξ'' . We find

$$\xi' \times \xi'' = v^3 \kappa \mathbf{B} \Rightarrow ||\xi' \times \xi''|| = v^3 \kappa.$$

Hence

$$\kappa = \frac{||\xi' \times \xi''||}{||\xi'||^3}.$$
(1.70)

In order to find the torsion of a curve ξ we consider the triple product of the vector fields ξ', ξ'', ξ''' . From (1.67),(1.68), (1.69) it follows that

$$(\xi', \xi'', \xi''') = \upsilon^6 \kappa^2 \tau.$$

$$\tau = \frac{(\xi', \xi'', \xi''')}{||\xi' \times \xi''||^2}.$$
 (1.71)

Hence

O NE of the most important concepts of modern differential geometry is the notion of a vector field. The importance of this concept lies in the fact that it can be used not only in differential geometry, but also in mechanics and theoretical physics. Imagine that we were able to take a snapshot of the flow of a current fluid and at each point we constructed a velocity vector of infinitesimal particles of this fluid. We obtain the distribution of vectors in a certain region of space, which would naturally be called a vector field. Maxwell's great insight was the idea of describing electromagnetic phenomena with the help of a concept of an electromagnetic field. A vector field, considered as the distribution of vectors in a certain region of space, is very clear and easy to understand. However, it turns out that an approach to vector field based on directional derivative is more effective and useful. The concept of a differential form is not so visual, it is more abstract and in order to explain a notion of a differential forms plays an exceptionally important role not only in differential geometry, but also in field theories.

2.1 VECTOR FIELDS IN EUCLIDEAN SPACE

In this section we continue the study of vector fields in an affine Euclidean space E^n started in the previous chapter (Section 1.5). Recall that a vector field in an open subset U of an affine space E^n is defined as a smooth mapping $\mathbf{X} : p \in U \to \mathbf{X}_p \in T_p E^n$, which attributes to each point p of U a tangent vector of an affine space E^n at this point p. We can give this definition by means of the disjoint union of tangent spaces to E^n over an open subset U, i.e. $T(U) = \bigcup_{p \in U} T_p E^n$, as a smooth mapping $\mathbf{X} : U \to T(U)$, which satisfies $\pi \circ \mathbf{X} = \mathrm{id}_U$, where $\pi : T(U) \to U$ is the projection defined by the formula $\pi(\mathbf{v}_p) = p$. The set of all smooth vector fields in an open subset U is denoted by $\mathcal{D}(U)$ and this set becomes the vector space if we define pointwise the addition of vector fields and multiplication by real numbers. One can also define a multiplication of vector fields by smooth functions $f\mathbf{X}$, where f is a smooth function, \mathbf{X} is a vector field and $f\mathbf{X}$ is a vector field, whose value at a point $p \in U$ is the vector

 $f(p)\mathbb{X}_p$. Multiplication by smooth functions together with previously defined addition of vector fields and their multiplication by real numbers makes the set of all vector fields $\mathcal{D}(U)$ a module over the algebra of smooth functions $C^{\infty}(U)$. In the case of an affine Euclidean space E^n we have an inner product and the tangent space T_pE^n at a point p is an Euclidean space and this Euclidean structure induces the inner product of vector fields $\langle \mathbf{X}, \mathbf{Y} \rangle$ as follows: The inner product $\langle \mathbf{X}, \mathbf{Y} \rangle$ is the function on U, whose value at a point $p \in U$ is the inner product of tangent vectors $\langle \mathbf{X}_p, \mathbf{Y}_p \rangle$.

So far we have considered the notion of a vector field in a coordinate free form. This approach has certain advantages, for instance, it makes possible to find invariant structures, that is, structures which are independent of a choice of coordinate system. Such a structure depends directly on the geometry of a space and is not affected by an accidental nature of a choice of coordinate system. However an introduction of a coordinate system allows one to perform calculations and this is the reason why we break this symmetry and choose some specific coordinate system. The most convenient coordinate system is a system of affine coordinates. We will choose it as the basic coordinate system and show how a vector field can be described by means of affine coordinates. First we fix a reference frame $\{O; \vec{e_1}, \vec{e_2}, \ldots, \vec{e_n}\}$ in an affine space E^n and denote the corresponding affine coordinates by x^1, x^2, \ldots, x^n . Then at any point $p \in E^n$ we have the frame $\{\mathbf{e}_{p,1}, \mathbf{e}_{p,2}, \dots, \mathbf{e}_{p,n}\}$ of the tangent space $T_p E^n$. Hence the set of all vectors $\mathbf{e}_{p,i}$, where *i* is fixed and *p* runs over affine space E^n , determines the constant vector field, which will be denoted by D_i . Thus the value of a vector field D_i at a point p (we denote it by $(D_i)_p$) is $(D_i)_p = e_{p,i}$. In the case when E^n is an affine Euclidean space, i.e. there is an inner product, we will use Cartesian coordinates, i.e. we will assume that a reference frame $\{O; \vec{e_1}, \vec{e_2}, \ldots, \vec{e_n}\}$ is orthonormal. Making use of the multiplication of a vector field by a function, one can expand any vector field $X \in \mathcal{D}(U)$ into the sum of the basic vector fields D_1, D_2, \ldots, D_n as follows

$$\mathbf{X} = \sum_{i=1}^{n} X^{i} \mathbf{D}_{i}$$

where X^1, X^2, \ldots, X^n are smooth functions defined on U. Hence in affine coordinates a vector field **X** is uniquely determined by the functions X^i and we will call these functions *components of a vector field* **X** in affine coordinates.

It is easy to see that previously defined operations with vector fields can be expressed in terms of components of these vector fields as follows. If (X^1, X^2, \ldots, X^n) and (Y^1, Y^2, \ldots, Y^n) are the components of vector fields **X**, **Y** respectively, then the components of the sum **X**+**Y** are the functions $(X^1+Y^1, X^2+Y^2, \ldots, X^n+Y^n)$. If E^n is an affine Euclidean space we have the inner product of two vector fields < **X**, **Y** > and this inner product can be written in the Cartesian coordinates as follows

$$< \mathbf{X}, \mathbf{Y} > = \sum_{i,j=1}^{n} X^{i} Y^{i}.$$

If we multiply a vector field **X** by a smooth function f then the components of the vector field $f \mathbf{X}$ are the functions $(f X^1, f X^2, \dots, f X^n)$.

In Cartesian coordinates a vector field X can be expressed in terms of the basic fields D_1, D_2, \ldots, D_n by means of inner product as follows

$$\mathtt{X} = \sum_{i=1}^n < \mathtt{X}, \mathtt{D}_i > \mathtt{D}_i$$

Hence in order to calculate the components of a vector field X in Cartesian coordinates one can use the formula $X^i = \langle X, D_i \rangle$.

2.2 DIRECTIONAL DERIVATIVE

The concept of a vector field is similar to the concept of a function. Indeed a function assigns to each point of a domain a real number and a vector field assigns a vector. Methods for studying the behavior of a function are based on the study of the derivative of a function. Similarly the study of the behavior of a vector field, for instance its rate of change, is based on a differentiation of a vector field. This section is devoted to differentiation of vector fields. But we will begin this section with a concept of directional derivative of a function.

Let $f \in C^{\infty}(U)$ be a smooth function and $\mathbf{v}_p = (p; \vec{v}) \in T_p E^n$ be a tangent vector at a point $p \in U$. A tangent vector \mathbf{v}_p determines the straight line passing through a point p, whose parametric equation can be written in the form $p + t\vec{v}$. The restriction of a function f to the straight line $p + t\vec{v}$ will be denoted by $f(p + t\vec{v})$. It is easy to see that this restriction of a function f is the function of one variable t.

Definition 2.1. A directional derivative of a function f at a point p is defined by the formula

$$\mathbf{v}_p \triangleright f = \frac{d}{dt} \big(f(p + t\vec{v}) \big) \big|_{t=0}.$$

Thus the directional derivative is a mapping $\mathbf{v}_p : C^{\infty} \to \mathbb{R}$. This mapping is linear and satisfies the Leibniz rule, i.e. for any smooth functions f, g and any real numbers $a, b \in \mathbb{R}$ one has the following properties

- 1. $\mathbf{v}_p \triangleright (a f + b g) = a \mathbf{v}_p \triangleright f + b \mathbf{v}_p \triangleright g$,
- 2. $\mathbf{v}_p \triangleright (f g) = (\mathbf{v}_p \triangleright f) g + f (\mathbf{v}_p \triangleright g).$

Definition (2.1) is given in a coordinate free form. If we now assume that an affine space E^n is endowed with a system of affine coordinates x^1, x^2, \ldots, x^n then a function f is the function of n variables $f(x^1, x^2, \ldots, x^n)$ and a vector \vec{v} can be written in coordinates $\vec{v} = (v^1, v^2, \ldots, v^n)$. Now making use of Definition (2.1) we find

$$\mathbf{v}_p \triangleright f = \frac{d}{dt} \left(f(p+t\vec{v}) \right) \Big|_{t=0} = \sum_{i=1}^n \frac{d}{dt} (p^i + t \, v^i) \Big|_{t=0} \frac{\partial f}{\partial x^i} \Big|_p = \sum_{i=1}^n v^i \frac{\partial f}{\partial x^i} \Big|_p. \tag{2.1}$$

One can use this formula to compute a directional derivative of a function in affine coordinates. For instance applying (2.1) we find

$$\mathbf{e}_{p,i} \triangleright f = \frac{\partial f}{\partial x^i}\Big|_p. \tag{2.2}$$

In the case of Cartesian coordinates we can write this formula by means of an inner product and the gradient as follows

$$\mathbf{v}_p \triangleright f = <\mathbf{v}_p, (\operatorname{grad} f)_p > . \tag{2.3}$$

We can extend the notion of directional derivative to vector fields. Let $\mathbf{X} \in \mathcal{D}(U)$ be a vector field. Then according to the definition a vector field \mathbf{X} assigns to each point p of U a tangent vector \mathbf{X}_p . Given a smooth function $f \in C^{\infty}(U)$ one can find a directional derivative $\mathbf{X}_p \triangleright f$. This directional derivative is a real number. Hence one can assign to each point p of U a real number, that is, one has the real-valued function defined on U. Let us denote this function by $\mathbf{X} \triangleright f$. Then according to the previous considerations the value of the function $\mathbf{X} \triangleright f$ at a point p is the real number $\mathbf{X}_p \triangleright f$. For example making use of (2.2) we find $\mathbf{D}_i \triangleright f = \frac{\partial f}{\partial x^i}$. Hence in affine coordinates for any vector field $\mathbf{X} = \sum_{i=1}^n X^i \mathbf{D}_i$ we have

$$\mathbf{X} \triangleright f = \sum_{i=1}^{n} X^{i} \left(\mathbf{D}_{i} \triangleright f \right) = \sum_{i=1}^{n} X^{i} \frac{\partial f}{\partial x^{i}}.$$
(2.4)

It follows from (2.3) that in the case of Cartesian coordinates we can write the last formula in the form

$$\mathbf{X} \triangleright f = \sum_{i=1}^{n} < \mathbf{X}, \text{grad } f > .$$
(2.5)

Formula (2.4) shows that in affine coordinates a vector field X induces the first order differential operator

$$\mathbf{X} \to \sum_{i=1}^{n} X^{i} \frac{\partial}{\partial x^{i}}.$$
 (2.6)

It also shows that if f is a smooth function then $\mathbf{X}[f]$ is also a smooth function. Thus a vector field \mathbf{X} induces the mapping $f \in C^{\infty} \to \mathbf{X}[f] \in C^{\infty}$ which is linear and satisfies the Leibniz rule, that is, for any two smooth functions f, g and any two real numbers a, b we have

1. $X \triangleright (a f + b g) = a (X \triangleright f) + b (X \triangleright g),$

2.
$$X \triangleright (f g) = (X \triangleright f) g + f (X \triangleright g).$$

Definition 2.2. Let \mathcal{A} be a unital associative algebra. Then a linear mapping δ : $\mathcal{A} \to \mathcal{A}$ is said to be an algebra derivation if it satisfies the Leibniz rule

$$\delta(u\,v) = \delta(u)\,v + u\,\delta(v),$$

where $u, v \in \mathcal{A}$.

From this definition it follows that a vector field can be considered as a derivation of the algebra of smooth functions on U. It can be proved that any derivation of the algebra of smooth functions $C^{\infty}(U)$ can be uniquely written in affine coordinates in the form of first order differential operator (2.6). Thus in what follows we will

consider a vector field as a derivation of an algebra of smooth functions and identify it with a first order differential operator (2.6), i.e.

$$\mathbf{X} = \sum_{i=1}^{n} X^{i} \frac{\partial}{\partial x^{i}}.$$
(2.7)

Particularly we have

$$\mathbf{D}_i = \frac{\partial}{\partial x^i}, \ \ (\mathbf{D}_i)_p = \mathbf{e}_{p,i} = \frac{\partial}{\partial x^i}\Big|_p.$$

Now the affine basis $\{\mathbf{e}_{p,1}, \mathbf{e}_{p,2}, \dots, \mathbf{e}_{p,n}\}$ for a tangent space $T_p E^n$ at a point p can be written in the form

$$\left\{\frac{\partial}{\partial x^{1}}\Big|_{p}, \frac{\partial}{\partial x^{2}}\Big|_{p}, \dots, \frac{\partial}{\partial x^{n}}\Big|_{p}\right\}.$$
(2.8)

In the case of Cartesian coordinates this basis is orthonormal and we have

$$< \frac{\partial}{\partial x^i}\Big|_p, \frac{\partial}{\partial x^j}\Big|_p > = \delta_{ij}.$$
 (2.9)

Let us now consider the question of vector fields in curvilinear coordinates. To do this, we formulate in a suitable way a scheme for working with vector fields in affine coordinates, in particular, in Cartesian coordinates. In this case we start with an affine frame $\{O; \vec{e_1}, \vec{e_2}, \ldots, \vec{e_n}\}$ which induces a system of affine coordinates x^1, x^2, \ldots, x^n . Then we move in a parallel way the vectors of basis $\vec{e_1}, \vec{e_2}, \ldots, \vec{e_n}$ to each point p of a space E^n with the help of notion of a vector at a point. Thus we get n constant vector fields D_1, D_2, \ldots, D_n and at each point of a space E^n these vector fields form the basis for a tangent space $T_p E^n$. In order to extend this approach to curvilinear coordinates we can consider the vectors of basis for a tangent space from a slightly different point of view. Let p be a point in space E^n . A straight line $p+t \vec{e_i}$ is called the *i*th coordinate line of an affine coordinate system passing through a point p. Hence ncoordinate lines pass through each point in a space. It is easy to see that the tangent vector to a coordinate line $p + t \vec{e_i}$ at a point p is the *i*th vector of basis $\mathbf{e}_{p,i}$ of a tangent space $T_p E^n$. On the other hand if we compute a directional derivative of a function f along this coordinate line then

$$\frac{d}{dt} \left(f(p+t\,\vec{e_i}) \right) \Big|_{t=0} = \frac{\partial f}{\partial x^i} \Big|_p$$

and we can identify an ith vector of basis with partial derivative with respect to ith coordinate at a point p.

Let us now apply this approach to curvilinear coordinates. Hence we assume that we have a system of curvilinear coordinates $\psi : U \subset E^n \to W \subset \mathbb{R}^n$ which is C^{∞} compatible with affine coordinates x^1, x^2, \ldots, x^n . The curvilinear coordinates will be denoted by $\tilde{x}^1, \tilde{x}^2, \ldots, \tilde{x}^n$ and

$$x^{1} = f^{1}(\tilde{x}^{1}, \tilde{x}^{2}, \dots, \tilde{x}^{n}), \ x^{2} = f^{2}(\tilde{x}^{1}, \tilde{x}^{2}, \dots, \tilde{x}^{n}), \dots, x^{n} = f^{n}(\tilde{x}^{1}, \tilde{x}^{2}, \dots, \tilde{x}^{n}), \ (2.10)$$

will be the functions of transition from an affine coordinates to curvilinear coordinates. Let $p = (\tilde{x}_0^1, \tilde{x}_0^2, \dots, \tilde{x}_0^n)$ be a point of U. Then

$$\xi_i(t) = (\tilde{x}_0^1, \tilde{x}_0^2, \dots, \tilde{x}_0^{i-1}, \tilde{x}_0^i + t, \tilde{x}_0^{i+1}, \dots, x_0^n)$$

is an *i*th coordinate line of a system of curvilinear coordinates passing through a point p. Calculating a directional derivative of a function f along a coordinate line $\xi_i(t)$ we find

$$\frac{d}{dt}(\xi_i(t))\Big|_{t=0} = \sum_{k=0}^n \frac{\partial f}{\partial \tilde{x}^k}\Big|_p \frac{d\tilde{x}^k}{dt}\Big|_{t=0} = \frac{\partial f}{\partial \tilde{x}^i}\Big|_p$$

This formula shows that for any point $p \in U$ the tangent vectors

$$\frac{\partial}{\partial \tilde{x}^1}\Big|_p, \frac{\partial}{\partial \tilde{x}^2}\Big|_p, \dots, \frac{\partial}{\partial \tilde{x}^n}\Big|_p$$
(2.11)

form the basis for a tangent space $T_p E^n$ in curvilinear coordinates $\tilde{x}^1, \tilde{x}^2, \ldots, \tilde{x}^n$. Hence in curvilinear coordinates the vector fields

$$\tilde{\mathsf{D}}_1 = \frac{\partial}{\partial \tilde{x}^1}, \ \tilde{\mathsf{D}}_2 = \frac{\partial}{\partial \tilde{x}^2}, \ \dots, \ \tilde{\mathsf{D}}_n = \frac{\partial}{\partial \tilde{x}^n}$$
 (2.12)

form the basis of a tangent space $T_p E^n$ at each point of an affine space E^n .

It is useful to find how the basic vector fields (2.12) in curvilinear coordinates can be expressed in terms of the basic vector fields in affine coordinates. Consider a smooth function f defined in affine coordinates x^1, x^2, \ldots, x^n . We can pass from affine coordinates to curvilinear coordinates by means of the transition functions (2.10) and we obtain the composite function whose partial derivative with respect to \tilde{x}^i can be written as follows

$$\frac{\partial f}{\partial \tilde{x}^i} = \frac{\partial x^j}{\partial \tilde{x}^i} \frac{\partial f}{\partial x^j}$$

Since f is an arbitrary smooth function we find the transformation rule for basic vector fields

$$\tilde{\mathsf{D}}_i = \sum_{j=1}^n \frac{\partial x^j}{\partial \tilde{x}^i} \mathsf{D}_j.$$
(2.13)

It is worth to mention that this formula describes a transition from one basis for a tangent space $T_p E^n$ to another (at each point p of U) because the Jacobian of this transition (the determinant of the matrix $A = (A_i^j) = (\partial x^j / \partial \tilde{x}^i)$) is non-zero (see Proposition 1.1).

Formula (2.13) is very important because it leads to several important conclusions. First we see that in curvilinear coordinates the basic vector fields $\tilde{D}_1, \tilde{D}_2, \ldots, \tilde{D}_n$ are not constant vector fields. Hence the basis (2.11) for a tangent space $T_p E^n$ depends on a point of a space and it changes when one moves from one point of a space to another. Second formula (2.11) is important when E^n is an affine Euclidean space. Indeed in curvilinear coordinates we can not use formulae of Euclidean geometry for inner product and a length, but formula (2.11) shows a way how to calculate inner product of tangent vectors if we use curvilinear coordinates. Let us consider the case when E^n is an affine Euclidean space. Then

$$\begin{split} <\tilde{\mathbf{D}}_{i},\tilde{\mathbf{D}}_{j}> &= <\sum_{k=1}^{n}\frac{\partial x^{k}}{\partial \tilde{x}^{i}}\,\mathbf{D}_{k},\sum_{m=1}^{n}\frac{\partial x^{m}}{\partial \tilde{x}^{j}}\,\mathbf{D}_{m}> =\sum_{k,m=1}^{n}\frac{\partial x^{k}}{\partial \tilde{x}^{i}}\,\frac{\partial x^{m}}{\partial \tilde{x}^{j}}<\mathbf{D}_{k},\mathbf{D}_{m}> \\ &= \sum_{k,m=1}^{n}\frac{\partial x^{k}}{\partial \tilde{x}^{i}}\,\frac{\partial x^{m}}{\partial \tilde{x}^{j}}\,\delta_{km} =\sum_{k=1}^{n}\frac{\partial x^{k}}{\partial \tilde{x}^{i}}\,\frac{\partial x^{k}}{\partial \tilde{x}^{j}}. \end{split}$$

Let us define the square matrix of *n*th order $\tilde{g} = (\tilde{g}_{ij})$, where

$$\tilde{g}_{ij} = \sum_{k=1}^{n} \frac{\partial x^k}{\partial \tilde{x}^i} \frac{\partial x^k}{\partial \tilde{x}^j}.$$

The matrix \tilde{g} is referred to as a matrix of Euclidean metric in curvilinear coordinates. Obviously the matrix of Euclidean metric in Cartesian coordinates g is the unit matrix $g = (\delta_{ij})$ because $\langle D_i, D_j \rangle = \delta_{ij}$. The square of the arc differential expressed in terms of coordinate differentials is usually referred to as the Euclidean metric. In Cartesian coordinate the Euclidean metric can be written as

$$ds^{2} = (dx^{1})^{2} + (dx^{2})^{2} + \ldots + (dx^{n})^{2} = \sum_{i=1}^{n} \delta_{ij} dx^{i} dx^{j}.$$

Hence in curvilinear coordinates the Euclidean metric has the form

$$ds^{2} = \sum_{i,j=1}^{n} \tilde{g}_{ij} dx^{i} dx^{j} = \sum_{i,j,k=1}^{n} \frac{\partial x^{k}}{\partial \tilde{x}^{i}} \frac{\partial x^{k}}{\partial \tilde{x}^{j}} dx^{i} dx^{j}.$$

2.3 EXTERIOR ALGEBRA OF A VECTOR SPACE

In this section we will use Einstein's convention to write sums containing quantities with indices. This convention is that if the same index occurs twice in an expression, once as a superscript and the second time as a subscript, then one takes the sum over this index, but the symbol of the sum is omitted.

Let V be an n-dimensional vector space. A linear form on a vector space V is a function $\rho: V \to \mathbb{R}$ which satisfies the linearity condition $\rho(a \vec{v}+b \vec{w}) = a \rho(\vec{v})+b \rho(\vec{w})$, where $a, b \in \mathbb{R}$ and $\vec{v}, \vec{w} \in V$. We will also call a linear form a 1-form. The set of all 1-forms defined on a vector space V is the vector space if one defines the addition of 1-forms and their multiplication by real numbers as follows

$$(\varrho + \varsigma)(\vec{v}) = \varrho(\vec{v}) + \varsigma(\vec{v}), \ (a \, \varrho)(\vec{v}) = a \, \varrho(\vec{v}). \tag{2.14}$$

The vector space of 1-forms is called a *dual space* to a vector space V and it will be denoted by V^* . It can be proved that the dual space V^* is *n*-dimensional vector space. If $\{\vec{e}_1, \vec{e}_2, \ldots, \vec{e}_n\}$ is a basis for a vector space V then this basis induces the dual basis $\{e^1, e^2, \ldots, e^n\}$ for the dual vector space and a 1-form e^i of the dual basis is defined by $e^i(\vec{e}_j) = \delta^i_j$. Then any 1-form ϱ can be expressed in terms of the dual basis as follows

$$\varrho = \varrho_i e^i, \ \varrho_i \in \mathbb{R}.$$

Note that in this formula we use the Einstein convention. A *bilinear form* on a vector space V is a function of two vector variables $\vartheta: V \times V \to \mathbb{R}$ which is linear in both variables, i.e.

$$\vartheta(a\,\vec{v}_1 + b\,\vec{v}_2, \vec{w}) = a\,\vartheta(\vec{v}_1, \vec{w}) + b\,\vartheta(\vec{v}_2, \vec{w}),$$

and the analogous property holds for the second variable \vec{w} . We will shortly call a bilinear form a 2-form. The set of 2-forms is the vector space if one defines the sum

of two 2-forms and the multiplication by real numbers similarly to (2.14). A 2-form ϑ is said to be a skew-symmetric 2-form if for any two vectors \vec{v}, \vec{w} it is satisfies

$$\vartheta(\vec{v}, \vec{w}) = -\vartheta(\vec{w}, \vec{v}). \tag{2.15}$$

Evidently the set of skew-symmetric 2-forms is the subspace of the vector space of all 2-forms and this subspace will be denoted by $\wedge^2 V^*$. Generally a *n*-form on a vector space V is a function of n vector variables $\varsigma : V \times V \times \ldots \times V \to \mathbb{R}$ which is linear in every variable. The set of n-forms on a vector space V is a vector space if one defines the sum of two n-forms and the multiplication of n-forms by real numbers similarly to (2.14). A *l*-form is said to be a skew-symmetric if its satisfies

$$\varsigma(\vec{v}_{i_1}, \vec{v}_{i_2}, \dots, \vec{v}_{i_l}) = (-1)^{\mathsf{p}} \varsigma(\vec{v}_1, \vec{v}_2, \dots, \vec{v}_l), \tag{2.16}$$

where **p** is a parity of a permutation (i_1, i_2, \ldots, i_l) . Hence a skew symmetric *l*-form does not change if we perform an even permutation of its vector variables and it changes a sign in the case of odd permutation. The set of skew-symmetric *l*-forms is the subspace in the vector space of all skew-symmetric *l*-forms and this subspace will be denoted by $\wedge^l V^*$.

The wedge product of two 1-forms $\varrho_1, \varrho_2 \in V^*$ is a 2-form $\varrho_1 \wedge \varrho_2$ defined by

$$(\varrho_1 \wedge \varrho_2)(\vec{v}, \vec{w}) = \begin{vmatrix} \varrho_1(\vec{v}) & \varrho_1(\vec{w}) \\ \varrho_2(\vec{v}) & \varrho_2(\vec{w}) \end{vmatrix} = \varrho_1(\vec{v}) \, \varrho_2(\vec{w}) - \varrho_1(\vec{w}) \, \varrho_2(\vec{v}).$$
(2.17)

It is easy to see that for any 1-forms ρ_1, ρ_2, ρ_3 and real numbers a, b we have the following algebraic properties

- 1. $V^* \wedge V^* \to \wedge^2 V^*,$ i.e. the wedge product of two 1-forms is a skew-symmetric 2-form,
- 2. $\rho_1 \wedge \rho_2 = -\rho_2 \wedge \rho_1$, i.e. wedge product is anti-symmetric operation,
- 3. $(a \varrho_1 + b \varrho_2) \wedge \varrho_3 = a \varrho_1 \wedge \varrho_3 + b \varrho_2 \wedge \varrho_3$, i.e. wedge product is bilinear operation.

We can extend the wedge product (2.17) to any k-form ϑ and m-form ς by saying that $\vartheta \wedge \varsigma$ is a *l*-form, where l = k + m, defined by

$$(\vartheta \wedge \varsigma)(\vec{v}_1, \vec{v}_2, \dots, \vec{v}_l) = \sum_{\substack{1 \le i_1 < \dots < i_k \le n \\ 1 \le j_1 < \dots < j_m \le n}} (-1)^{\mathbf{p}} \,\vartheta(\vec{v}_{i_1}, \vec{v}_{i_2}, \dots, \vec{v}_{i_k}) \varsigma(\vec{v}_{j_1}, \vec{v}_{j_2}, \dots, \vec{v}_{j_m}) (2.18)$$

where $(i_1, i_2, \ldots, i_k, j_1, j_2, \ldots, j_m)$ is a permutation of the integers $(1, 2, \ldots, l)$ and **p** is the parity of this permutation. Particularly if k = m = 1 then the general formula (2.18) takes on the form of wedge product of two 1-forms (2.17). It can be verified that the wedge product $\vartheta \wedge \varsigma$ defined in (2.18) is the *l*-ary skew-symmetric form. The wedge product has the following algebraic properties

- 1. if $\vartheta \in \wedge^k V^*$, $\varsigma \in \wedge^m V^*$ then $\vartheta \wedge \varsigma \in \wedge^{k+m} V^*$,
- 2. $(\rho \land \vartheta) \land \varsigma = \rho \land (\vartheta \land \varsigma)$, (associativity)

3. if
$$\vartheta \in \wedge^k V^*$$
, $\varsigma \in \wedge^m V^*$ then $\vartheta \wedge \varsigma = (-1)^{km} \varsigma \wedge \vartheta$,

4. $(a \varrho_1 + b \varrho_2) \wedge \varrho_3 = a \varrho_1 \wedge \varrho_3 + b \varrho_2 \wedge \varrho_3$, where a, b are real numbers.

In order to make the notation of vector space of forms uniform, we will denote the dual space V^* as follows $\wedge^1 V^*$. Hence the sum of vector spaces $V^* + \wedge^2 V^* + \ldots + \wedge^l V^* + \ldots$, i.e. the vector space of all formal finite linear combinations of forms with real coefficients, endowed with the wedge product, is the associative algebra. It is useful to make it a unital algebra by adding the vector space $\wedge^0 V^* = \mathbb{R}$. The real numbers \mathbb{R} can be considered as a one-dimensional vector space spanned by the unity 1. Hence $\wedge^0 V^* = \{a \ 1 : a \in \mathbb{R}\}$. The elements of the one-dimensional vector space $\wedge^0 V^*$ will be referred to as 0-forms. Hence a 0-form here is simply a real number. Naturally we extend the wedge product to the 0-forms by saying that a wedge product of a 0-form and a k-form is multiplication of a k-form by a real number defined in (2.14). It is worth to mention here that due to (2.14) it is natural to define $\rho \wedge a = \rho a = a \rho$, that is, forms commute with real numbers. Now the unity 1 is the identity element of the exterior algebra because $1 \wedge \rho = \rho \wedge 1 = \rho$.

Definition 2.3. The unital associative algebra $(\wedge V^*, \wedge)$, where $\wedge V^*$ is the sum of vector spaces $\wedge^0 V^* + V^* + \ldots + \wedge^k V^* + \ldots$ and \wedge is the wedge product, is called an exterior algebra of a vector space V.

The vector space of the exterior algebra is the sum of vector spaces of k-forms $\wedge^k V^*$. But this sum is not infinite as it may seem at the first glance. Actually this sum ends at the vector space of n-forms $\wedge^n V^*$ (n is the dimension of a vector space V), that is, all vector spaces $\wedge^l V^*$, l > n are trivial, i.e. consist of the zero vector. The reason for this can be understood if we consider the structure of vector spaces $\wedge^k V^*$ with the help of a basis.

Let $\{\vec{e_1}, \vec{e_2}, \ldots, \vec{e_n}\}$ be a basis for V. Then we have the dual basis of 1-forms $\{e^1, e^2, \ldots, e^n\}$ for the dual vector space V^* . Now if we take into account that wedge product is anti-symmetric, i.e. for any integers i, j it holds $e^i \wedge e^j = -e^j \wedge e^i$ then $e^i \wedge e^i = 0$. Thus if a sequence of integers i_1, i_2, \ldots, i_k , where each integer of this sequence can take values from 1 to n, contains two identical integers then, because of anti-symmetry of the wedge product, the product $e^{i_1} \wedge e^{i_2} \wedge \ldots \wedge e^{i_k}$ will be 0. Therefore we will get a non-zero product of basic 1-forms e^i only in the case when all integers in a sequence i_1, i_2, \ldots, i_k are distinct. On the other hand, if we rearrange the integers $i_{\sigma(1)}, i_{\sigma(2)}, \ldots, i_{\sigma(k)}$, where $(\sigma(1), \sigma(2), \ldots, \sigma(k))$ is a permutation of integers $1, 2, \ldots, k$, we get the product, which is expressed as follows

$$e^{i_{\sigma(1)}} \wedge e^{i_{\sigma(2)}} \wedge \ldots \wedge e^{i_{\sigma(k)}} = (-1)^{\mathbf{p}(\sigma)} e^{i_1} \wedge e^{i_2} \wedge \ldots \wedge e^{i_k},$$

where $\mathbf{p}(\sigma)$ is the parity of a permutation σ . These considerations lead to the conclusion that all products $e^{i_1} \wedge e^{i_2} \wedge \ldots \wedge e^{i_k}$, where $1 \leq i_1 < i_2 < \ldots < i_k \leq n$, are independent and can be shown to form a basis for a vector space of k-forms. Thus the dimension of a vector space $\wedge^k V^*$ is C_k^n , i.e. dim $\wedge^k V^* = C_k^n$. Now it becomes clear that the longest non-zero product of basic 1-forms e^i that we can form is the

n-form $e^1 \wedge e^2 \wedge \ldots \wedge e^n$ which span the 1-dimensional vector space $\wedge^n V^*$ and every vector space of k-forms, where k > n, is trivial, i.e. $\wedge^k V^* = \{0\}$ for k > n. Hence

$$\wedge V^* = \wedge^0 V^* + \wedge^1 V^* + \wedge^2 V^* + \ldots + \wedge^n V^*.$$

Any k-form ρ can be uniquely expressed in terms of wedge products of basic 1-forms e^i as follows

$$\varrho = \sum_{1 \le i_1 < i_2 < \dots < i_k \le n} \varrho_{i_1 i_2 \dots i_k} e^{i_1} \wedge e^{i_2} \wedge \dots \wedge e^{i_k}, \qquad (2.19)$$

where $\varrho_{i_1i_2...i_k}$ are real numbers. This formula can be written in a more compact form. Let $\mathcal{I} = \{i_1, i_2, ..., i_k\}$ $(k \leq n)$ be a subset of the set of integers from 1 to n, i.e. if we denote $\mathcal{N} = \{1, 2, ..., n\}$ then $\mathcal{I} \subset \mathcal{N}$. Recall that the concept of a set does not imply ordering of elements. Hence we can choose natural ordering $1 \leq i_1 < i_2 < ... < i_k \leq n$. Now we associate to every subset $\mathcal{I} \subset \mathcal{N}$ the basic k-form as follows

$$\mathcal{I} \subset \mathcal{N} \quad \to \quad e^{\mathcal{I}} = e^{i_1} \wedge e^{i_2} \wedge \ldots \wedge e^{i_k}. \tag{2.20}$$

We assume that a subset \mathcal{I} may be empty. In this case we associate to the empty set the identity element of the exterior algebra $\wedge V^*$, that is, $e^{\emptyset} = 1$. Now formula (2.19) can be written in a compact form as follows

$$\varrho = \sum_{\mathcal{I} \subset \mathcal{N}} \, \varrho_{\mathcal{I}} \, e^{\mathcal{I}}. \tag{2.21}$$

There is one more way to express a k-form in terms of the basic k-forms and in this case we do not need to use ordered sequences of indexes, which makes calculations with forms in some cases easier. If we assume that the coefficients $\varrho_{i_1i_2...i_k}$ in the sum (2.19) are skew-symmetric in subscripts, that is,

$$\varrho_{i_{\sigma(1)}i_{\sigma(2)}\dots i_{\sigma(k)}} = (-1)^{\mathbf{p}(\sigma)} \, \varrho_{i_1,i_2,\dots,i_k},$$

where σ is a permutation and $p(\sigma)$ is its parity, and we take the sum over all sequences i_1, i_2, \ldots, i_k then due to the anti-symmetry of the wedge product we will get every term k! times. Thus we can express any k-form as follows

$$\varrho = \frac{1}{k!} \, \varrho_{i_1 i_2 \dots i_k} \, e^{i_1} \wedge e^{i_2} \wedge \dots \wedge e^{i_k}. \tag{2.22}$$

In the above formula we can use the Einstein convention because now every subscript independently of others takes values from 1 to n.

Let us consider the question of transformation of the coefficients $\varrho_{i_1i_2...i_k}$ of a k-form when one passes from one basis of a vector space V to another. Let $\{\vec{e}'_1, \vec{e}'_2, ..., \vec{e}'_n\}$ be another basis for a vector space V and $\vec{e}'_i = A^j_i \vec{e}_j$, where $A = (A^j_i)$ is a transition matrix. Let $\{\tilde{e}^1, \tilde{e}^2, ..., \tilde{e}^n\}$ be the dual basis for a basis $\{\vec{e}'_1, \vec{e}'_1, ..., \vec{e}'_1\}$. Then $e^i = A^i_i \tilde{e}^j$.

In order to find a transformation law for the coefficients of a k-form ρ it is useful to consider the determinant of a submatrix of a transition matrix A. Let $1 \le i_1 < i_2 < \ldots < i_k \le n$ be numbers of rows and $1 \le j_1 < j_2 < \ldots < j_k \le n$ be numbers

of columns of a matrix A. Then the elements of a transition matrix A located at the intersections of these rows and columns form the submatrix of matrix A. Let us denote this submatrix by $A_{(j_1j_2...j_k)}^{(i_1i_2...i_k)}$. It should be noted that here we mean the following. The elements of the first row of the submatrix are the elements of a matrix A located at the intersection of the row i_1 with columns j_1, j_2, \ldots, j_k , the elements of the second row of the submatrix are the elements of a matrix A located at the intersection of the row i_1 with columns j_1, j_2, \ldots, j_k , the elements of a determinant it follows that Det $(A_{(j_1j_2...j_k)}^{(i_1i_2...i_k)})$ is skew-symmetric in superscripts as well as in subscripts. Indeed if we do a permutation of the superscripts i_1, i_2, \ldots, i_k then this will be equivalent to a permutation of the rows in the determinant Det $(A_{(j_1j_2...j_k)}^{(i_1i_2...i_k)})$. If we do a permutation of the columns j_1, j_2, \ldots, j_k then this will be equivalent to a permutation of the rows in the determinant Det $(A_{(j_1j_2...j_k)}^{(i_1i_2...i_k)})$. If we do a permutation of the columns j_1, j_2, \ldots, j_k then this will be equivalent to a permutation of the rows in the determinant Det $(A_{(j_1j_2...j_k)}^{(i_1i_2...i_k)})$. If we do a permutation of the columns in determinant Det $(A_{(j_1j_2...j_k)}^{(i_1i_2...i_k)})$. Hence

$$\begin{aligned} & \operatorname{Det} \left(A_{(j_1 j_2 \dots i_k)}^{(i_{\sigma(1)} i_{\sigma(2)} \dots i_{\sigma(k)})} \right) &= (-1)^{\mathtt{p}} \operatorname{Det} \left(A_{(j_1 j_2 \dots j_k)}^{(i_1 i_2 \dots i_k)} \right), \\ & \operatorname{Det} \left(A_{(j_{\sigma(1)} j_{\sigma(2)} \dots j_{\sigma(k)})}^{((i_1 i_2 \dots i_k))} \right) &= (-1)^{\mathtt{p}} \operatorname{Det} \left(A_{(j_1 j_2 \dots j_k)}^{(i_1 i_2 \dots i_k)} \right), \end{aligned}$$

where **p** is the parity of a permutation σ .

Let ρ be a k-form. We can write it either in the coordinates of the first basis $\{\vec{e}_1, \vec{e}_2, \ldots, \vec{e}_n\}$ or in the coordinates of the second basis $\{\vec{e}_1', \vec{e}_2', \ldots, \vec{e}_n'\}$, that is,

$$\rho = \frac{1}{k!} \, \rho_{i_1 i_2 \dots i_k} \, e^{i_1} \wedge e^{i_2} \wedge \dots \wedge e^{i_k}, \quad \rho = \frac{1}{k!} \, \tilde{\rho}_{i_1 i_2 \dots i_k} \, \tilde{e}^{i_1} \wedge \tilde{e}^{i_2} \wedge \dots \wedge \tilde{e}^{i_k}. \tag{2.23}$$

Making use of $e^i = A^i_j \tilde{e}^j$ we find for $1 \le j_1 < j_2 < \ldots < j_k \le n$

$$\tilde{\varrho}_{j_1 j_2 \dots j_k} = \sum_{1 \le i_1 < i_2 < \dots < i_k \le n} \operatorname{Det} \left(A^{(i_1 i_2 \dots i_k)}_{(j_1 j_2 \dots j_k)} \right) \varrho_{i_1 i_2 \dots i_k}.$$
(2.24)

Taking into account the skew-symmetry of a determinant $\text{Det}(A_{(j_1j_2...i_k)}^{(i_1i_2...i_k)})$ both with respect to the superscripts and subscripts we can drop the condition for the ordering of indices and then

$$\tilde{\varrho}_{j_1 j_2 \dots j_k} = \frac{1}{k!} \operatorname{Det} \left(A^{(i_1 i_2 \dots i_k)}_{(j_1 j_2 \dots j_k)} \right) \, \varrho_{i_1 i_2 \dots i_k}.$$
(2.25)

Definition 2.4. Assume that in each coordinate system $\{\vec{e}_1, \vec{e}_2, \ldots, \vec{e}_n\}$ of a vector space V we are given a set of real numbers $\{\varrho_{i_1i_2...i_k}\}$ indexed by the set of subscripts $i_1i_2...i_k$, where each subscript runs from 1 to n, and $\varrho_{i_1i_2...i_k}$ are skew-symmetric under a permutation of subscripts. If $\{\varrho_{i_1i_2...i_k}\}$ transform under transition from one coordinate system to another according to the law (2.25) then in a vector space V we have the covariant skew-symmetric tensor of kth order ϱ and $\{\varrho_{i_1i_2...i_k}\}$ are components of this tensor in a coordinate system $\{\vec{e}_1, \vec{e}_2, \ldots, \vec{e}_n\}$.

2.4 DIFFERENTIAL FORMS

In this section the exterior algebra of a vector space described in the previous section will be used to construct differential forms in an affine Euclidean space E^n .

At each point p of an affine Euclidean space E^n we have the tangent space $T_p E^n$.

Let us denote by $T_p^*E^n$ the dual space of a tangent space T_pE^n . This dual space is a vector space of dimension n and will be called a *cotangent space* at a point p. Hence we can consider the exterior algebra $\wedge T_p^*E^n$. According to the considerations of the previous section the exterior algebra $\wedge T_p^*E^n$ is the sum of vector spaces

$$\wedge T_p^* E^n = \wedge^0 T_p^* E^n + \wedge^1 T_p^* E^n + \ldots + \wedge^n T_p^* E^n,$$

where $\wedge^k T_p^* E^n$ is a vector space of k-forms defined on a tangent space $T_p E^n$ and $\wedge^1 T_p^* E^n = T_p^* E^n$. Let $U \subset E^n$ be an open subset in E^n . In analogy with a tangent space we define $\wedge^k T^*(U) = \bigcup_{p \in U} \wedge^k T_p^* E^n, \wedge T^*(U) = \bigcup_{p \in U} \wedge T_p^* E^n$, where \cup is a disjoint union of vector spaces. The projection $\pi : \wedge^k T^*(U) \to U$ is defined in a natural way, i.e. if ϑ is a k-form defined on a tangent space $T_p E^n$ then $\pi(\vartheta) = p$.

Definition 2.5. A smooth mapping $\omega : U \to \wedge^k T^*(U)$, which satisfies $\pi \circ \omega = id_U$, is called a differential k-form defined on an open subset U.

Hence a differential k-form attributes to each point of an open subset of an affine Euclidean space a k-form and this k-form depends smoothly on a point of U. A vector space of differential k-forms defined on an open subset $U \subset E^n$ will be denoted by $\Omega^k(U)$. Here the vector space structure of $\Omega^k(U)$ is defined by pointwise addition of differential k-forms and by pointwise multiplication by real numbers, that is, $(\omega + \theta)_p = \omega_p + \theta_p$ and $(a \omega)_p = a \omega_p$, where $\omega, \theta \in \Omega^k(U), a \in \mathbb{R}$. As in the case of vector fields, we can define one more operation with differential k-forms and this is multiplication of a differential k-form ω by a smooth function $f \in C^{\infty}(U)$ defined pointwise $(f \omega)_p = f(p) \omega_p$. Usually we will multiply a differential k-form by a function f from the left and write $f \omega$, but it is natural to define $\omega f = f \omega$, that is, differential forms commute with smooth functions.

The multiplication of differentials k-forms by smooth functions has the following properties

- $(f+g)\omega = f\omega + g\omega$,
- $f(\omega + \theta) = f\omega + g\theta$,
- $(f g) \omega = f (g \omega),$
- $\mathbf{1} \omega = \omega$, where **1** is the constant function whose value at each point p is 1.

Hence the multiplication of differential k-forms by smooth functions defines the structure of module over the ring of smooth functions on $\Omega^k(U)$.

At each point p of an affine space E^n the tangent space T_pE^n is a vector space of dimension n. Therefore, according to the previous section, at every point p one can construct the exterior algebra $\wedge T_p^*E^n$. This algebra can be extended to the vector space of differential forms if we define a wedge product of two differential forms pointwise. In order to do this we consider the sum of vector spaces $\Omega^k(U)$ whose elements are all possible linear combinations of differential forms with real coefficients, i.e. $\sum_{i=1}^n a_i \,\omega^{(i)}$, where $\omega^{(i)} \in \Omega^i(U), a_i \in \mathbb{R}$. We denote this sum of vector spaces by $\Omega(U)$, that is,

$$\Omega(U) = \Omega^0(U) + \Omega^1(U) + \ldots + \Omega^n(U),$$

and $\Omega^0(U)$ is the vector space of smooth functions $C^{\infty}(U)$. It is clear that the multiplication of differential k-forms by smooth functions induces the structure of module on $\Omega(U)$ if we put $f \sum_{i=1}^{n} \omega^{(i)} = \sum_{i=1}^{n} f\omega^{(i)}$, where $\omega^{(i)} \in \Omega^i(U)$. Now we define the wedge product of any two differential forms pointwise $(\omega \wedge \theta)_p = \omega_p \wedge \theta_p$ and extend it in an obvious way to $\Omega(U)$. Note that the wedge product of a zero-form (function) and a differential k-form (k > 0) is the multiplication of a differential k-form by a function. Evidently the constant function **1**, whose value at any point of an affine space E^n is the unity 1, plays a role of identity element for the wedge product.

Hence $(\Omega(U), \wedge)$ is a unital associative algebra which has the following properties

- if $\omega \in \Omega^k(U), \theta \in \Omega^m(U)$ then $\omega \wedge \theta \in \Omega^{k+m}(U)$,
- $(\omega \wedge \theta) \wedge \rho = \omega \wedge (\theta \wedge \rho)$ (associativity),
- $\omega \wedge \theta = (-1)^{km} \theta \wedge \omega$, where $\omega \in \Omega^k(U), \theta \in \Omega^m(U)$,
- $(f \omega + g \theta) \wedge \rho = f \omega \wedge \rho + g \omega \wedge \rho$, where f, g are smooth functions.

The algebra $(\Omega(U), \wedge)$ will be referred to as an algebra of differential forms defined on an open subset U.

Definition 2.6. A unital associative algebra (\mathcal{A}, \cdot) is said to be a \mathbb{Z} -graded algebra if it is a direct sum of subspaces $\mathcal{A}^i \subset \mathcal{A}$ labelled by integers $i \in \mathbb{Z}$, that is, $\mathcal{A} = \bigoplus_{i \in \mathbb{Z}} \mathcal{A}^i$ and for any $k, m \in \mathbb{Z}, u \in \mathcal{A}^k, v \in \mathcal{A}^m$ it holds $u \cdot v \in \mathcal{A}^{k+m}$. If for any $k, m \in \mathbb{Z}, u \in \mathcal{A}^k, v \in \mathcal{A}^m$ a multiplication of a graded algebra \mathcal{A} satisfies $u \cdot v = (-1)^{km}v \cdot u$ then graded algebra \mathcal{A} is referred to as a graded commutative algebra.

It follows from this definition that the algebra of differential forms $(\Omega(U), \wedge)$ defined on an open subset U of an affine space E^n is a \mathbb{Z} -graded commutative algebra whose subspaces $\Omega^k(U)$ are trivial for k < 0 and k > n.

Let $\psi: U \subset E^n \to W \subset \mathbb{R}^n$ be a system of coordinates (x^1, x^2, \dots, x^n) (not necessarily affine coordinates) C^{∞} -compatible with affine coordinates. Then the vectors at a point $p \in U$

$$\frac{\partial}{\partial x^1}\Big|_p, \frac{\partial}{\partial x^2}\Big|_p, \dots, \frac{\partial}{\partial x^n}\Big|_p \tag{2.26}$$

form the basis for the tangent space $T_p E^n$. The dual basis for the vector space $T_p^* E^n$ (or $\wedge^1 T_p^* E^n$) will be denoted by

$$(dx^1)_p, (dx^2)_p, \dots, (dx^n)_p.$$
 (2.27)

Hence we have

$$(dx^i)_p \left(\frac{\partial}{\partial x^j}\Big|_p\right) = \delta^i_j.$$

Let ω be a differential k-form and $\omega_p, p \in U$ be its value at a point p. According to Definition (2.5) $\omega_p \in \wedge^k T_p^* E^n$. Formula (2.22) shows that we can write a differential k-form ω at a point p as follows

$$\omega_p = \frac{1}{k!} \,\omega_{i_1 i_2 \dots i_k}(p) \,(dx^{i_1})_p \wedge (dx^{i_2})_p \wedge \dots \wedge (dx^{i_k})_p, \tag{2.28}$$

Since p is an arbitrary point of U we can omit point p in formula (2.28) and then we will obtain a formula for a differential k-form ω in coordinates x^1, x^2, \ldots, x^n . In this case the numbers $\omega_{i_1i_2...i_k}(p)$ in (2.28) become functions $\omega_{i_1i_2...i_k}$ of the coordinates x^1, x^2, \ldots, x^n of a point. The requirement of smoothness for a differential form considered as a mapping $U \to \wedge T^*(U)$ now means that all functions $\omega_{i_1i_2...i_k}$ are smooth, i.e. $\omega_{i_1i_2...i_k} \in C^{\infty}(U)$. Hence a differential k-form can be written in coordinates x^1, x^2, \ldots, x^n as the following expression

$$\omega = \frac{1}{k!} \omega_{i_1 i_2 \dots i_k}(x) \, dx^{i_1} \wedge dx^{i_2} \wedge \dots \wedge dx^{i_k}, \qquad (2.29)$$

where x stands for the set of coordinates x^1, x^2, \ldots, x^n of a variable point of U.

At each point p of an open subset $U \subset E^n$ a differential k-form ω is a k-form ω_p , that is, a real-valued multilinear skew-symmetric function of k vectors. On the other hand if we are given k vector fields $\mathbf{X}_1, \mathbf{X}_2, \ldots, \mathbf{X}_k \in \mathcal{D}(U)$ then at each point $p \in U$ we have k tangent vectors $(\mathbf{X}_1)_p, (\mathbf{X}_2)_p, \ldots, (\mathbf{X}_k)_p$. Thus at each point p we can calculate the value of a k-form ω_p on k tangent vectors and this is a real number. Hence we have the function

$$p \in U \to \omega_p((\mathbf{X}_1)_p, (\mathbf{X}_2)_p, \dots, (\mathbf{X}_k)_p).$$

This function is called the value of a differential k-form ω on vector fields X_1, X_2, \ldots, X_k and we will denote this function by $\omega(X_1, X_2, \ldots, X_k)$.

Obviously the value of a differential k-form on k vector fields is skew-symmetric under a permutation of vector fields, that is,

$$\omega(\mathbf{X}_{\sigma(1)},\mathbf{X}_{\sigma(2)},\ldots,\mathbf{X}_{\sigma(k)}) = (-1)^{\mathbf{p}} \ \omega(\mathbf{X}_1,\mathbf{X}_2,\ldots,\mathbf{X}_k),$$

where σ is a permutation of integers $\{1, 2, ..., k\}$ and **p** is the parity of σ . Due to the multilinearity of a k-form we have one more property

$$\omega(\mathbf{X}_1, \mathbf{X}_2, \dots, \mathbf{X}_i + f \mathbf{Y}, \dots, \mathbf{X}_k) = \omega(\mathbf{X}_1, \mathbf{X}_2, \dots, \mathbf{X}_{i-1}, \mathbf{X}_i, \mathbf{X}_{i+1}, \dots, \mathbf{X}_k)$$

+ $f \omega(\mathbf{X}_1, \mathbf{X}_2, \dots, \mathbf{X}_{i-1}, \mathbf{Y}, \mathbf{X}_{i+1}, \dots, \mathbf{X}_k), (2.30)$

where f is a smooth function defined on U and $1 \le i \le k$. If a differential k-form and vector fields are written in coordinates as follows

$$\omega = \frac{1}{k!} \,\omega_{i_1 i_2 \dots i_k}(x) \, dx^{i_1} \wedge dx^{i_2} \wedge \dots \wedge dx^{i_k}, \ \mathbf{X}_1 = X_1^i \frac{\partial}{\partial x^i}, \ \mathbf{X}_2 = X_2^i \frac{\partial}{\partial x^i}, \dots, \mathbf{X}_k = X_k^i \frac{\partial}{\partial x^i}$$

then the value of differential k-form ω on vector fields X_1, X_2, \ldots, X_k can be written as follows

$$\omega(\mathbf{X}_{1}, \mathbf{X}_{2}, \dots, \mathbf{X}_{k}) = \frac{1}{k!} \omega_{i_{1}i_{2}\dots i_{k}}(x) \begin{vmatrix} X_{1}^{i_{1}} & X_{2}^{i_{1}} & \dots & X_{k}^{i_{1}} \\ X_{1}^{i_{2}} & X_{2}^{i_{2}} & \dots & X_{k}^{i_{2}} \\ \dots & \dots & \dots & \dots \\ X_{1}^{i_{k}} & X_{2}^{i_{k}} & \dots & X_{k}^{i_{k}} \end{vmatrix}.$$
 (2.31)

Given a differential k-form ω one can find the expression of this form in coordinates x^1, x^2, \ldots, x^n by means of the basic vector fields $D_i = \partial/\partial x^i$, where $i = 1, 2, \ldots, n$, as follows

$$\omega_{i_1 i_2 \dots i_k} = \omega(\mathsf{D}_{i_1}, \mathsf{D}_{i_2}, \dots, \mathsf{D}_{i_k}) = \omega(\frac{\partial}{\partial x^{i_1}}, \frac{\partial}{\partial x^{i_2}}, \dots, \frac{\partial}{\partial x^{i_k}}).$$
(2.32)

An important structure in the case of differential forms is an exterior differential. It is an exterior differential that underlies the theory of de Rham's cohomologies, which is used to study the global structure of a manifold. We will define an exterior differential step by step and start with 0-forms, i.e. functions. It is worth to mention here that any differential form is uniquely defined if we show how to calculate its value on arbitrary vector fields (the number of vector fields depend on the degree of a differential form). Let f be a smooth function defined on an open subset $U \subset E^n$. The exterior differential of a function f is the 1-form df defined by

$$df(\mathbf{X}) = \mathbf{X} \triangleright f. \tag{2.33}$$

It follows from this definition that in the case of a 0-form the exterior differential d raises the degree of a form by one, i.e. $d : \Omega^0(U) \to \Omega^1(U)$. It also follows from (2.33) that d is a linear mapping. In order to find the expression of the 1-form df in coordinates we calculate the value of this form on basic vector fields D_i . We obtain

$$df(\mathbf{D}_i) = \mathbf{D}_i \triangleright f = \frac{\partial f}{\partial x^i}.$$

Hence

$$df = \frac{\partial f}{\partial x^1} dx^1 + \frac{\partial f}{\partial x^2} dx^2 + \ldots + \frac{\partial f}{\partial x^n} dx^n = \frac{\partial f}{\partial x^i} dx^i.$$
(2.34)

The expression on the right-hand side of the obtained formula is called in the differential calculus of functions the total differential of a function. Thus at the level of functions we can identify the exterior differential with the total differential and this leads to an important conclusion that d satisfies the Leibniz rule

$$d(f h) = (df) h + f (dh).$$
(2.35)

It should be noted that a difference between the exterior differential and the total differential appears when we pass from first order differential to higher-order differentials. Formula (2.34) also explains why the dual basis of a cotangent space $T_p^* E^n$ was denoted as differentials of coordinates $(dx^1)_p, (dx^2)_p, \ldots, (dx^n)_p$. Indeed the coordinates can be considered as functions on space (coordinate functions). Then according to exterior calculus the exterior differential of a coordinate function at some point in space is an element of a cotangent space. However it is worth remembering that in the calculus of differential forms, in contrast to the differential calculus of functions, the differentials at point p are generators of the exterior algebra $\wedge T_p^* E^n$, that is, we can multiply them using exterior multiplication.

Now let ω be a 1-form defined on U. The exterior differential of 1-form ω is the 2-form $d\omega$ defined by

$$d\omega(\mathbf{X}, \mathbf{Y}) = \mathbf{X} \triangleright (\omega(\mathbf{Y})) - \mathbf{Y} \triangleright (\omega(\mathbf{X})) - \omega([\mathbf{X}, \mathbf{Y}]).$$
(2.36)

It follows from this definition that, just as in the case of functions, the exterior differential raises the degree of a 1-form by one, that is, it assigns a 2-form to a 1-form. The linearity of the exterior differential in the case of 1-forms also easily follows

from Formula (??). Hence in this case d is a linear mapping $d : \Omega^1(U) \to \Omega^2(U)$. Thus we have the sequence of vector spaces

$$\Omega^{0}(U) \xrightarrow{d} \Omega^{1}(U) \xrightarrow{d} \Omega^{2}(U).$$
(2.37)

The exterior differential has a very important property $d^2 = d \circ d = 0$. Indeed let f be a smooth function and X, Y be two vector fields defined on U. Then

$$\begin{aligned} d^2 f(\mathbf{X}, \mathbf{Y}) &= (d (df))(\mathbf{X}, \mathbf{Y}) = \mathbf{X} \triangleright (df(\mathbf{Y})) - \mathbf{Y} \triangleright (df(\mathbf{X})) - df([\mathbf{X}, \mathbf{Y}]) \\ &= \mathbf{X} \triangleright (\mathbf{Y} \triangleright f) - \mathbf{Y} \triangleright (\mathbf{X} \triangleright f) - [\mathbf{X}, \mathbf{Y}] \triangleright f = [\mathbf{X}, \mathbf{Y}] \triangleright f - [\mathbf{X}, \mathbf{Y}] \triangleright f = 0. \end{aligned}$$

Any 1-form ω can be expressed in coordinates x^1, x^2, \ldots, x^n as follows $\omega = \omega_i dx^i$. Now our aim is to find the expression for the 2-form $d\omega$ in coordinates x^1, x^2, \ldots, x^n . First of all we know that any 2-form can be written as follows $d\omega = \frac{1}{2}\omega_{ij}dx^i \wedge dx^j$. In order to find the coefficient functions ω_{ij} we use $\omega_{ij} = \omega(\mathsf{D}_i, \mathsf{D}_j)$. Hence

$$\omega_{ij} = d\omega(\mathsf{D}_i,\mathsf{D}_j) = \mathsf{D}_i \triangleright (\omega(\mathsf{D}_j)) - \mathsf{D}_j \triangleright (\omega(\mathsf{D}_i)) - \omega([\mathsf{D}_i,\mathsf{D}_j]).$$

Next we have $\omega(D_i) = \omega_i, \omega(D_j) = \omega_j$ and $[D_i, D_j] = [\frac{\partial}{\partial x^i}, \frac{\partial}{\partial x^j}] = 0$ because partial derivatives commute. We obtain

$$\omega_{ij} = \mathbf{D}_i \triangleright \omega_j - \mathbf{D}_j \triangleright \omega_i = \frac{\partial \omega_j}{\partial x^i} - \frac{\partial \omega_i}{\partial x^j},$$

and

$$\omega = \omega_i \, dx^i \ \to d\omega = \frac{1}{2!} \left(\frac{\partial \omega_j}{\partial x^i} - \frac{\partial \omega_i}{\partial x^j} \right) \, dx^i \wedge dx^j.$$

Now we extend the exterior differential to any k-form ω by the following definition.

Definition 2.7. Let ω be a k-form and $X_1, X_2, \ldots, X_{k+1}$ be vector fields. Then the exterior differential $d\omega$ is the (k + 1)-form defined by

$$d\omega(\mathbf{X}_1, \mathbf{X}_1, \dots, \mathbf{X}_{k+1}) = \sum_i (-1)^{i-1} \mathbf{X}_i \triangleright \left(\omega \left(\mathbf{X}_1, \dots, \hat{\mathbf{X}}_i, \dots, \mathbf{X}_{k+1} \right) \right) \\ + \sum_{i < j} (-1)^{i+j} \omega([\mathbf{X}_i, \mathbf{X}_j], \mathbf{X}_1, \dots, \hat{\mathbf{X}}_i, \dots, \hat{\mathbf{X}}_j, \dots, \mathbf{X}_{k+1}),$$

where a hat $\hat{}$ over a vector field X_i means that this vector field is omitted.

Theorem 2.1. The exterior differential has the following properties:

- 1. the exterior differential is a linear mapping $d: \Omega^i(U) \to \Omega^{i+1}(U)$,
- 2. for any differential k-form ω and any differential form θ it holds

$$d(\omega \wedge \theta) = (d\omega) \wedge \theta + (-1)^k \,\omega \wedge (d\theta),$$

3. $d^2 = d \circ d = 0$.

If $\omega = \frac{1}{k!} \omega_{i_1 i_2 \dots i_k} dx^{i_1} \wedge dx^{i_2} \wedge \dots \wedge dx^{i_k}$ then

$$d\omega = \sum_{j_1 < \dots < j_{k+1}} (-1)^{m-1} \frac{\partial \omega_{j_1 \dots \hat{j}_m \dots j_{k+1}}}{\partial x^{j_m}} dx^{j_1} \wedge \dots \wedge dx^{j_{k+1}}, \qquad (2.38)$$

where \hat{j}_m means that j_m is omitted.

Proof. The linearity of the exterior differential follows from the definition of the sum of two differential k-forms and the multiplication of a differential form by a real number. Indeed if $\omega, \theta \in \Omega^k(U)$ and $a \in \mathbb{R}$ then, by definition,

$$(\omega + \theta)(\mathbf{X}_1, \mathbf{X}_2, \dots, \mathbf{X}_k) = \omega(\mathbf{X}_1, \mathbf{X}_2, \dots, \mathbf{X}_k) + \theta(\mathbf{X}_1, \mathbf{X}_2, \dots, \mathbf{X}_k), (a\,\omega)(\mathbf{X}_1, \mathbf{X}_2, \dots, \mathbf{X}_k) = a\,(\omega(\mathbf{X}_1, \mathbf{X}_2, \dots, \mathbf{X}_k)).$$
(2.39)

Applying Definition (2.7) and using the linearity of directional derivative we obtain $d(\omega + \theta) = d\omega + d\theta$ and $d(a \omega) = a d\omega$.

In order to prove the second property we will use coordinates x^1, x^2, \ldots, x^n . It can be easily checked that the proposed proof can be applied in any coordinate system. Hence the second property takes places in any coordinate system. First of all let us note that due to the linearity of the exterior differential we need to prove the second property only in the case of differential forms

$$\omega = f \, dx^{i_1} \wedge dx^{i_2} \wedge \ldots \wedge dx^{i_k}, \quad \theta = g \, dx^{j_1} \wedge dx^{j_2} \wedge \ldots \wedge dx^{j_l}, \quad f, g \in C^{\infty}(U)$$
(2.40)

because any differential form can be written as the sum of this kind of differential forms. Second it is sufficient to prove the second property only in the case of basic vector fields D_i because any vector field can be written as the sum of basic vector fields multiplied by functions and a differential form has the property (2.30). First we prove that the exterior differential of a k-form ω , where ω is the first form in (2.40), can be written in coordinates as follows

$$d\omega = df \wedge dx^{i_1} \wedge \ldots \wedge dx^{i_k}.$$

Let $D_1, D_2, \ldots, D_{k+1}$ be basic vector fields. Making use of Definition (2.7) we can write

$$d\omega(\mathbf{D}_1, \mathbf{D}_2, \dots, \mathbf{D}_{k+1}) = \sum_{m=1}^{k+1} (-1)^{m-1} \mathbf{D}_m \triangleright (\omega(\mathbf{D}_1, \dots, \hat{\mathbf{D}}_m, \dots, \mathbf{D}_{k+1})).$$
(2.41)

The second sum in the formula of Definition (2.7) vanishes because partial derivatives commute $[D_i, D_j] = 0$. Next

$$\omega(\mathbf{D}_1,\ldots,\hat{\mathbf{D}}_m,\ldots,\mathbf{D}_{k+1}) = f \, dx^{i_1} \wedge dx^{i_2} \wedge \ldots \wedge dx^{i_k}(\mathbf{D}_1,\ldots,\hat{\mathbf{D}}_m,\ldots,\mathbf{D}_{k+1})$$

The value of a k-form $dx^{i_1} \wedge dx^{i_2} \wedge \ldots \wedge dx^{i_k}$ on basic vector fields $D_i, i \neq m$ is a number since according to the definition of the wedge product this value will be equal

to the determinant whose elements are numbers $dx^{i_r}(D_s) = \delta_s^{i_r}$, where $s \neq m$. Thus we have

$$d\omega(\mathbf{D}_{1},\mathbf{D}_{2},\ldots,\mathbf{D}_{k+1}) = \sum_{m=1}^{k+1} (-1)^{m-1} (\mathbf{D}_{m} \triangleright f) \, dx^{i_{1}} \wedge \ldots \wedge dx^{i_{k}} (\mathbf{D}_{1},\ldots,\hat{\mathbf{D}}_{m},\ldots,\mathbf{D}_{k+1})$$

$$= \sum_{m=1}^{k+1} (-1)^{m-1} df(\mathbf{D}_{m}) \, dx^{i_{1}} \wedge \ldots \wedge dx^{i_{k}} (\mathbf{D}_{1},\ldots,\hat{\mathbf{D}}_{m},\ldots,\mathbf{D}_{k+1})$$

$$= df \wedge dx^{i_{1}} \wedge \ldots \wedge dx^{i_{k}} (\mathbf{D}_{1},\ldots,\mathbf{D}_{m},\ldots,\mathbf{D}_{k+1})$$

Now we consider the exterior differential of the wedge product of two forms (2.40). Making use of (2.35) we find

$$\begin{aligned} d(\omega \wedge \theta) &= d(f h \, dx^{i_1} \wedge \ldots \wedge dx^{i_k} \wedge dx^{j_1} \wedge \ldots \wedge dx^{j_l}) \\ &= d(f h) \wedge dx^{i_1} \wedge \ldots \wedge dx^{i_k} \wedge dx^{j_1} \wedge \ldots \wedge dx^{j_l} \\ &= (d(f) h + f d(h)) \wedge dx^{i_1} \wedge \ldots \wedge dx^{i_k} \wedge dx^{j_1} \wedge \ldots \wedge dx^{j_l} \\ &= (df \wedge dx^{i_1} \wedge \ldots \wedge dx^{i_k}) \wedge (h \, dx^{j_1} \wedge \ldots \wedge dx^{j_l}) \\ &+ (-1)^k (f \, dx^{i_1} \wedge \ldots \wedge dx^{i_k}) \wedge (dh \wedge dx^{j_1} \wedge \ldots \wedge dx^{j_l}) \\ &= d\omega \wedge \theta + (-1)^k \omega \wedge d\theta. \end{aligned}$$

Now it is easy to prove that $d^2 = d \circ d = 0$. Previously we proved that this property holds in the case of functions. From previous considerations it follows that it is sufficient to prove $d^2 = 0$ only in the case of forms (2.40). We have

$$d^{2}(\omega) = d(df \wedge dx^{i_{1}} \wedge \ldots \wedge dx^{i_{k}}) = d^{2}(f) \wedge dx^{i_{1}} \wedge \ldots \wedge dx^{i_{k}} - df \wedge d^{2}x^{i_{1}} \wedge \ldots \wedge dx^{i_{k}} + \ldots + (-1)^{k} df \wedge dx^{i_{1}} \wedge \ldots \wedge d^{2}x^{i_{k}} = 0.$$

Inn order to prove formula (2.38) we use the properties of the exterior differential. Hence we have

$$d\omega = d\left(\sum_{i_1 < \dots < i_k} \omega_{i_1 \dots i_k} dx^{i_1} \wedge \dots \wedge dx^{i_k}\right)$$

$$= \sum_{i_1 < \dots < i_k} d\omega_{i_1 \dots i_k} \wedge dx^{i_1} \wedge \dots \wedge dx^{i_k}$$

$$= \sum_{i_1 < \dots < i_k} \frac{\partial \omega_{i_1 \dots i_k}}{\partial x^m} dx^m \wedge dx^{i_1} \wedge \dots \wedge dx^{i_k}.$$
 (2.42)

Since $d\omega$ is the differential (k+1)-form we can write it as follows

$$d\omega = \sum_{j_1 < \dots < j_{k+1}} \omega_{j_1 \dots j_{k+1}} dx^{j_1} \wedge \dots \wedge dx^{j_{k+1}}.$$
 (2.43)

Now let us consider the question of how the wedge products of the basic forms will be obtained from formula (2.42). For this purpose we choose some product $dx^{j_1} \wedge \ldots \wedge dx^{j_{k+1}}$. This product will be obtained from products $dx^{i_1} \wedge \ldots \wedge dx^{i_k}$ by multiplying them from the left by dx^m . To find those products $dx^{i_1} \wedge \ldots \wedge dx^{i_k}$ which will give $dx^{j_1} \wedge$

 $\dots \wedge dx^{j_{k+1}}$ we remove successively 1-forms $dx^{j_1}, dx^{j_2}, \dots, dx^{j_{k+1}}$ from the product $dx^{j_1} \wedge \dots \wedge dx^{j_{k+1}}$. It is obvious that in this way we obtain all the terms that give the form $dx^{j_1} \wedge \dots \wedge dx^{j_{k+1}}$. But since the 1-forms in the product $dx^{j_1} \wedge \dots \wedge dx^{j_{k+1}}$ are ordered we will have to put the differential 1-form dx^m (standing at the first position) to its proper place. Thus there will appear a sign which depends on a position of a corresponding 1-form. Thus we get

$$d\omega = \sum_{j_1 < \dots < j_{k+1}} \sum_{m=1}^{k+1} (-1)^{m-1} \frac{\partial \omega_{j_1 \dots \hat{j}_m \dots j_{k+1}}}{\partial x^{j_m}} \, dx^{j_1} \wedge \dots \wedge dx^{j_{k+1}}.$$
(2.44)

2.5 COVARIANT DERIVATIVE

In this section, we will consider the question of how to measure the rate of change of one vector field in the direction determined by another vector field. Assume that $U \subset E^n$ is an open subset, Y is a vector field defined on U and $\xi : I \to U$ is a parametrized curve. If we restrict a vector field Y to a parametrized curve ξ then we get the vector field $Y|_{\xi}$ along a curve ξ , i.e. a vector field depending on one parameter t. In this case the measure of change of a vector field along a curve ξ is its derivative $(Y|_{\xi})'$ with respect to the parameter t. Note that the derivative $(Y|_{\xi})'$ is a new vector field along a parametrized curve ξ .

Now assume that $p \in U$, $\mathbf{v} = (p; \vec{v}) \in T_p E^n$. Let ξ be a parametrized curve passing through a point p, whose velocity vector is \mathbf{v} . Such a parametrized curve is always exists because we can take, for instance, the straight line $p + t\vec{v}$. We have $\xi(t_0) = p, \xi'(t_0) = \mathbf{v}$. According to the construction described above, we can measure the rate of change of a vector field \mathbf{Y} at a point p in the direction of a tangent vector $\mathbf{v} = (p; \vec{v})$ if we take the derivative $(\mathbf{Y}|_{\xi})'$ of the restriction of a vector field \mathbf{Y} to ξ calculated at a point p as a measure of this change. This derivative of a vector field \mathbf{Y} is called *a covariant derivative of a vector field* \mathbf{Y} at a point p and is denoted by $\nabla_{\mathbf{v}}\mathbf{Y}$. Hence

$$\nabla_{\mathbf{v}}\mathbf{Y} = (\mathbf{Y}|_{\xi})'\big|_{t=t_0} = \frac{d}{dt}(\mathbf{Y}|_{\xi})\big|_{t=t_0}, \ \xi'(t_0) = \mathbf{v}.$$
 (2.45)

It follows from this definition that covariant derivative of a vector field at a point of U is a vector at this point.

We can extend the notion of a covariant derivative at a point to the whole open subset U as follows. Let X, Y be two vector fields defined on U. A vector field X defines the tangent vector X_p at each point $p \in U$. Therefore we can calculate the covariant derivative $\nabla_{X_p} Y$ at each point $p \in U$ and this derivative will determine a vector at each point of an open subset U. Hence we get the new vector field on U which will be denoted by $\nabla_X Y$ and called a *covariant derivative of a vector field* Y with respect to a vector field X. Hence a covariant derivative $\nabla_X Y$ is a vector field defined at each point p of an open subset U as follows

$$(\nabla_{\mathbf{X}}\mathbf{Y})_p = \nabla_{\mathbf{X}_p}\mathbf{Y}.\tag{2.46}$$

Later we will see that covariant derivative $\nabla_X Y$ is a smooth vector field provided that X, Y are smooth vector fields.

Now our aim is to find a formula for a covariant derivative in coordinates. We will do this first of all in affine coordinates and then in an arbitrary curvilinear coordinate system. Let us assume that we have an affine coordinate system x^1, x^2, \ldots, x^n defined in U. This means that we have an affine frame $\{O; \vec{e_1}, \vec{e_2}, \ldots, \vec{e_n}\}$ and the vector fields

$$\mathbf{U}_i = \frac{\partial}{\partial x^i}, \ \ (\mathbf{U}_i)_p = \mathbf{e}_{p,i} = (p; \vec{e_i}), \ \ p \in U.$$

Now any vector field can be expressed in the terms of basic vector fields as follows $\mathbf{Y} = Y^i \mathbf{U}_i$, where Y^i are components of a vector field \mathbf{Y} , i.e. smooth functions. It is important here that the basic vector fields \mathbf{U}_i are constant vector fields. By other words, if we restrict a vector field \mathbf{Y} to a parametrized curve ξ then the components Y^i of a vector field become smooth (composite) functions $Y^i(t)$ of one variable t, but the basic vector fields \mathbf{U}_i remain constant. Thus the change of a vector fields \mathbf{U}_i do not make any contribution to this change. Hence in the case of an affine coordinate system we have

$$(\mathbf{Y}|_{\xi})' = \frac{d}{dt}(\mathbf{Y}|_{\xi}) = \frac{dY^{i}(t)}{dt} \,\mathbf{U}_{i} + Y^{i} \,\frac{d}{dt} \Big(\mathbf{U}_{i}|_{\xi}\Big). \tag{2.47}$$

The second term at the right-hand side of the above formula vanishes because U_i is a constant vector field. Hence for any $p \in U$, $\mathbf{v} \in T_p E^n$ and a parametrized curve ξ such that $\xi(t_0) = p, \xi'(t_0) = \mathbf{v}$ we have

$$\nabla_{\mathbf{v}}\mathbf{Y} = \frac{d}{dt} (Y^i|_{\xi}) \Big|_{t=t_0} \mathbf{e}_{p,i} = (\mathbf{v} \triangleright Y^i) \mathbf{e}_{p,i}.$$
(2.48)

Thus the covariant derivative of a vector field Y at a point p with respect to a tangent vector \mathbf{v} is the tangent vector at a point p whose coordinates in an affine frame are the directional derivatives of the components Y^i of a vector field Y in the direction of a tangent vector \mathbf{v} . Particularly this formula shows that covariant derivative depends linearly on a vector \mathbf{v} and this property does not depend on a coordinate system. Thus we have

$$\nabla_{\mathbf{v}_1+\mathbf{v}_2}\mathbf{Y} = \nabla_{\mathbf{v}_1}\mathbf{Y} + \nabla_{\mathbf{v}_2}\mathbf{Y}.$$
(2.49)

From (2.48) it follows that if $\mathbf{X} = X^i \mathbf{U}_i$ is a vector field then

$$\nabla_{\mathbf{X}} \mathbf{Y} = (\mathbf{X} \triangleright Y^i) \, \mathbf{U}_i = X^j \frac{\partial Y^i}{\partial x^j} \, \frac{\partial}{\partial x^i}. \tag{2.50}$$

It follows from (2.49) that for any vector fields X_1, X_2, Y it holds

$$\nabla_{\mathbf{X}_1+\mathbf{X}_2}\mathbf{Y} = \nabla_{\mathbf{X}_1}\mathbf{Y} + \nabla_{\mathbf{X}_2}\mathbf{Y}.$$
(2.51)

Formula (2.50) also shows that in affine coordinates the components of the covariant derivative are smooth functions and, consequently, the covariant derivative is a smooth vector field.

Now consider a more general situation, that is, a coordinate system defined on an open set U must not be an affine coordinate system, but can be an arbitrary curvilinear coordinate system. The coordinates of this coordinate system will be denoted by x^1, x^2, \ldots, x^n . Then the basic vector fields induced by this coordinate system will be denoted by D_1, D_2, \ldots, D_n , where $D_i = \frac{\partial}{\partial x^i}$. It is important to note that in contrast with affine coordinates now the basic vector fields D_i are need not to be constant vector fields, that is, they can change from point to point. This leads to an important structure of modern differential geometry, which is called a *connection*. Assume that we are given a vector field Y and a parametrized curve ξ . Our aim is to find a rate of change of a vector field Y when a point moves along a curve ξ . We can do this as we did it before, i.e. we restrict a vector field Y to a parametrized curve ξ and then differentiate this restriction $X|_{\mathcal{E}}$ with respect to a parameter of a parametrized curve. But if we write this derivative in curvilinear coordinates then the right-hand side of formula (2.47) will change. Due to a non-constant nature of the basic vector fields D_i the second term at the right-hand side of (2.47) does not vanish. Let $\mathbf{Y} = Y^i \mathbf{D}_i$. Then we have

$$(\mathbf{Y}|_{\xi})' = \frac{d}{dt}(\mathbf{Y}|_{\xi}) = \frac{dY^{i}}{dt} \, \mathbf{D}_{i} + Y^{i} \, \frac{d}{dt} \Big(\mathbf{D}_{i}|_{\xi} \Big) = (\xi' \triangleright Y^{i}) \, \mathbf{D}_{i} + Y^{i} \, \nabla_{\xi'} \, \mathbf{D}_{i}. \tag{2.52}$$

It is useful to consider this formula at a fixed point p of a curve ξ . Without loss of generality we can assume that an open interval I of a parametrized curve $\xi; I \to U$ contains 0 and $\xi(0) = p, \xi'(0) = v$. Then taking t = 0 in (2.52) we get

$$\nabla_{\mathbf{v}}\mathbf{Y} = (\mathbf{Y}|_{\xi})'\Big|_{t=0} = (\mathbf{v} \triangleright Y^i) \ (\mathbf{D}_i)_p + Y^i(p) \ \nabla_{\mathbf{v}} \ \mathbf{D}_i$$
(2.53)

The covariant derivative $\nabla_{\mathbf{v}} \mathbf{D}_i$ is the vector at a point p and it can be expressed as a linear combination of the vectors of basis $(\mathbf{D}_i)_p$, i.e. $\nabla_{\mathbf{v}} \mathbf{D}_i = \Gamma_i^j (\mathbf{D}_j)_p$, where Γ_i^j are numbers defined at a point p, and, due to the linearity (2.49) of the covariant derivative in respect of a vector \mathbf{v} , the coefficients of this combination depend linearly on a vector \mathbf{v} .

Proposition 2.1. Let $p \in U$, $\mathbf{v} \in T_p E^n$ and Λ_i^j be numbers defined at a point p by the formula $\nabla_{\mathbf{v}} \mathbf{D}_i = \Lambda_i^j (\mathbf{D}_j)_p$. Then there exists uniquely determined system of covectors $\varphi_i^j \in T_p^* E^n$ such that $\varphi_i^j(\mathbf{v}) = \Lambda_i^j$.

Proof. First of all let us proof that if a_1, a_2, \ldots, a_n are real numbers defined at a point p and $\mathbf{v}_1, \mathbf{v}_2, \ldots, \mathbf{v}_n$ are linearly independent tangent vectors at a point p then this data uniquely determines a covector $\varphi \in T_p^* E^n$ such that $\varphi(\mathbf{v}_i) = a_i$. We choose a basis $\mathbf{e}_1, \mathbf{e}_2, \ldots, \mathbf{e}_n$ for the tangent space $T_p E^n$ and write each vector \mathbf{v}_i as a linear combination of the vectors of basis $\mathbf{v}_i = v_j^k \mathbf{e}_k$. The dual basis will be denoted by $\mathbf{e}^1, \mathbf{e}^2, \ldots, \mathbf{e}^n$. Obviously the determinant of the matrix (v_i^k) is non-zero number. The condition $\varphi(\mathbf{v}_i) = a_i$ leads to the system of linear equations $v_i^k \varphi_k = a_i$, where $\varphi(\mathbf{v}_k) = \varphi_k$. The matrix of this system is (v_i^k) and this matrix is regular, i.e. $\text{Det}(v_i^k) \neq 0$. Thus this system of equations has the solution $\varphi_1, \varphi_2, \ldots, \varphi_n$, and this solution is unique. Then $\varphi = \varphi_k \mathbf{e}^k$.

Now for every tangent vector \mathbf{v}_k we find the real numbers $(\Lambda_k)_i^j$, where

$$\nabla_{\mathbf{v}_k} \mathbf{D}_i = (\Lambda_k)_i^j \, (\mathbf{D}_j)_p.$$

From the previous considerations it follows that there exists uniquely determined system of covectors φ_i^j such that $\varphi_i^j(\mathbf{v}_k) = (\Lambda_k)_i^j$. Let $\mathbf{v} = a^k v_k$. Then

$$\nabla_{\mathbf{v}} \mathbf{D}_i = \nabla_{a^k \mathbf{v}_k} \mathbf{D}_i = a^k \nabla_{\mathbf{v}_k} \mathbf{D}_i = a^k (\Lambda_k)_i^j (\mathbf{D}_j)_p.$$

Thus $\Lambda_i^j = a^k (\Lambda_k)_i^j$. Hence

$$\varphi_i^j(\mathbf{v}) = \varphi_i^j(a^k \, \mathbf{v}_k) = a^k \, \varphi_i^j(\mathbf{v}_k) = a^k \, (\Lambda_k)_i^j = \Lambda_i^j.$$

In the proof of Proposition (2.1) we used an arbitrary basis \mathbf{v}_k for the tangent space $T_p E^n$ at a point p. But curvilinear coordinates induce the basis \mathbf{D}_k and we can use this basis to define the real numbers $(\Lambda_k)_i^j$. In this case they are denoted by $(\Gamma_k)_i^j$. Hence we have

$$\nabla_{\mathsf{D}_k} \,\mathsf{D}_i = (\Gamma_k)_i^j \,(\mathsf{D}_j)_p. \tag{2.54}$$

We can omit the brackets and write $\Gamma_{ki}^j = (\Gamma_k)_i^j$. The real numbers Γ_{ki}^j defined at a point p of an affine space E^n by the formula $\nabla_{\mathsf{D}_k} \mathsf{D}_i = \Gamma_{ki}^j (\mathsf{D}_j)_p$ are called *Christoffel symbols*.

The following formulas hold at each point $p \in U$

$$\nabla_{\mathbf{v}} \mathbf{D}_i = \varphi_i^j(\mathbf{v}) \ \mathbf{D}_j, \ \nabla_{\mathbf{v}} \mathbf{Y} = (\mathbf{v} \triangleright Y^i) \ (\mathbf{D}_i)_p + \varphi_i^j(\mathbf{v}) \ Y^i(p) \ (\mathbf{D}_j)_p.$$
(2.55)

Now we extend these formulas to the entire open subset U by replacing a tangent vector \mathbf{v} by a vector field \mathbf{X} and covectors φ_i^j by a differential 1-forms ω_i^j . Then the formulas in (2.55) take on the form

$$\nabla_{\mathbf{X}} \mathbf{D}_i = \omega_i^j(\mathbf{X}) \ \mathbf{D}_j, \ \ \nabla_{\mathbf{X}} \mathbf{Y} = (\mathbf{X} \triangleright Y^i) \ \mathbf{D}_i + \omega_i^j(\mathbf{X}) \ Y^i \ \mathbf{D}_j.$$
(2.56)

We see that the covariant derivative of basic vector fields D_i gives rise to the matrix of differential 1-forms $\omega = (\omega_i^j)$, which is referred to as a *matrix of connection* in an affine space E^n . In what follows we will call the elements of the matrix of connection ω_i^j connection 1-forms. If we write $\omega_i^j = \omega_{ki}^j dx^k$ then we have the relation $\omega_i^j(D_k) = \omega_{ki}^j = \Gamma_{ki}^j$. Hence $\omega_i^j = \Gamma_{ki}^j dx^k$.

Let us consider the second formula in (2.56). We can write it either by means of Christoffel symbols or connection 1-forms. If we use Christoffel symbols and take $X = D_k$ then the second formula in (2.56) takes on the form

$$\nabla_k \mathbf{Y} = \nabla_{\mathbf{D}_k} \mathbf{Y} = \left(\frac{\partial Y^i}{\partial x^k} + \Gamma^i_{kj} Y^j\right) \mathbf{D}_i.$$
(2.57)

From this it follows

$$\nabla_{\mathbf{X}}\mathbf{Y} = X^k \,\nabla_k \mathbf{Y} = \left(\frac{\partial Y^i}{\partial x^k} \,X^k + \Gamma^i_{kj} \,X^k \,Y^j\right) \mathsf{D}_i. \tag{2.58}$$

In the case when we prefer to use the calculus of differential forms we can write the second formula in (2.56) as follows

$$\nabla_{\mathbf{X}}\mathbf{Y} = (dY^i + \omega_i^i Y^j)(\mathbf{X}) \mathbf{D}_i.$$
(2.59)

Note that in this formula X is an arbitrary vector field. Since this vector field is both on the left side of the formula and on the right side, we can formally omit it and the formula will remain true. Hence

$$\nabla \mathbf{Y} = (dY^i + \omega^i_j Y^j) \mathbf{D}_i \tag{2.60}$$

However now we have to explain in what sense we understand the obtained formula. First of all we consider this formula as the definition of a new operator ∇ . In order to give a correct definition of this new operator, we should explain how we understand the expression on the right-hand side of formula (2.60). For this purpose we consider two modules $\Omega^1(U), \mathcal{D}(U)$ over the algebra of smooth functions $C^{\infty}(U)$. Here $\Omega^1(U)$ is the module of differential 1-forms and $\mathcal{D}(U)$ is the module of vector fields. We can form the tensor product of these two modules

$$\Omega^1(U,\mathcal{D}) = \Omega^1(U) \otimes_{C^\infty(U)} \mathcal{D}(U).$$

An element of this tensor product is referred to as a differential 1-form with values in vector fields. If we take the basis $\{dx^i\}$ for the module of differential 1-forms $\Omega^1(U)$ and $\{D_k\}$ for the module of vector fields $\mathcal{D}(U)$ then $\{dx^i \otimes D_k\}$ is the basis for the tensor product $\Omega^1(U, \mathcal{D})$. Thus any differential 1-form with values in vector fields θ can be expressed in terms of the basis as follows

$$\theta = \theta_i^k \, dx^i \otimes \mathsf{D}_k,\tag{2.61}$$

where θ_i^k are smooth functions. In order to define the value of θ on a vector field **X** we write θ in the form

$$\theta = \theta_i^k \ dx^i \otimes \mathsf{D}_k = (\theta_i^k \ dx^i) \otimes \mathsf{D}_k = \theta^k \otimes \mathsf{D}_k.$$

If $X \in \mathcal{D}(U)$ then we define the value of θ on a vector field X as follows

$$\theta(\mathbf{X}) = (\theta^k(\mathbf{X})) \, \mathbf{D}_k.$$

Hence we see that the value of θ on a vector field is the vector field and this justifies the term "differential 1-form with values in vector fields".

Now we can give a correct definition for the expression at the right-hand side of (2.60).

Definition 2.8. Let $Y = Y^i D_i$ be a vector field. Then the covariant differential of a vector field Y is a differential 1-form with values in vector fields ∇Y defined by the formula

$$\nabla \mathbf{Y} = (dY^i + \omega_i^i Y^j) \otimes \mathbf{D}_i.$$

Let us study in more detail the structure of the matrix of connection form in curvilinear coordinates. First of all, we note that in affine coordinates of the space E^n , the connection matrix is zero. Indeed, in this case the covariant derivative of the basic vector fields U_i is equal to zero due to the fact these fields are constant. Let us assume that we have two coordinate systems in $U \subset E^n$, where one is the system of affine coordinates x^1, x^2, \ldots, x^n and the other is a system of curvilinear coordinates $\tilde{x}^1, \tilde{x}^2, \ldots, \tilde{x}^n$. The transition functions from one system of coordinates to another will be denoted by $x^i = x^i(\tilde{x}^1, \tilde{x}^2, \ldots, \tilde{x}^n)$ (from curvilinear coordinates to affine coordinates) and $\tilde{x}^i = \tilde{x}^i(x^1, x^2, \ldots, x^n)$ (from affine coordinates to curvilinear). In order to simplify notations we will write them in the compact form $x = x(\tilde{x})$ and $\tilde{x} = \tilde{x}(x)$. The Jacobi matrix of the transition $x = x(\tilde{x})$ will be denoted by $A = (A_i^j)$ and the Jacobi matrix of the inverse transition $\tilde{x} = \tilde{x}(x)$ is the reciprocal matrix of A. Hence we have

$$A_i^j = \frac{\partial x^i}{\partial \tilde{x}^j}, \ (A^{-1})_i^j = \frac{\partial \tilde{x}^i}{\partial x^j}$$

In these formulas the elements of the transition matrix A are functions of curvilinear coordinates $\tilde{x}^1, \tilde{x}^2, \ldots, \tilde{x}^n$, while the elements of the reciprocal matrix A^{-1} are functions of affine coordinates x^1, x^2, \ldots, x^n . We have

$$\frac{\partial}{\partial \tilde{x}^i} = \frac{\partial x^j}{\partial \tilde{x}^i} \frac{\partial}{\partial x^j} = A^j_i \frac{\partial}{\partial x^j}, \quad \frac{\partial}{\partial x^i} = \frac{\partial \tilde{x}^j}{\partial x^i} \frac{\partial}{\partial \tilde{x}^j} = (A^{-1})^j_i \frac{\partial}{\partial \tilde{x}^j}$$

Making use of the notations $U_i = \frac{\partial}{\partial x^i}$, $D_i = \frac{\partial}{\partial \tilde{x}^i}$ we can write the above formulas as in the form

$$\mathsf{D}_i = A_i^j \, \mathsf{U}_j, \quad \mathsf{U}_i = (A^{-1})_i^j \, \mathsf{D}_j.$$

Now our aim is to find the matrix of connection form in curvilinear coordinates \tilde{x}^i . The definition of the matrix of connection in curvilinear coordinates is $\nabla D_i = \omega_i^j D_j$. Substituting $D_i = A_i^j U_j$ into the left-hand side of the definition of the matrix of connection, we get

$$\nabla \left(A_i^j \, \mathbf{U}_j \right) = \left(dA_i^j \right) \mathbf{U}_j + A_i^j \, \nabla \, \mathbf{U}_j.$$

But $\nabla U_j = 0$ and we obtain $\omega_i^j A_j^k U_k = dA_i^k U_k$. Hence $A_j^k \omega_i^k = dA_i^k$ or, in the matrix form $\omega A = dA$. From this we find the matrix of connection

$$\omega_j^i = (A^{-1})_k^i \, dA_j^k \quad \Leftrightarrow \quad \omega = A^{-1} \, dA. \tag{2.62}$$

We mentioned before that the elements of the reciprocal matrix A^{-1} are the functions of affine coordinates x^i . In (2.62) the first factor at the right-hand side is the matrix which is obtained from A^{-1} by substitution $x = x(\tilde{x})$.

The set of dual 1-forms for the set of basic vector fields $\{D_i\}$ in curvilinear coordinates is $\{d\tilde{x}^j\}$, i.e. $d\tilde{x}^j(D_i) = \delta_i^j$. Let us denote $\theta^i = d\tilde{x}^i$. We can find the expressions of these 1-forms in affine coordinates. Indeed let $\theta^i = \theta_i^i dx^j$. Then

$$\delta^i_j = \theta^i(\mathsf{D}_j) = \theta^i(A^k_j\,\mathsf{U}_k) = A^k_j\,\theta^i_m\,\,dx^m(\mathsf{U}_k) = A^k_j\,\theta^i_m\,\delta^m_k = \theta^i_k\,A^k_j.$$

Hence $\theta_k^i = (A^{-1})_k^i$ and $\theta^i = (A^{-1})_k^i dx^k$. Now we can find the exterior differential of the dual 1-forms θ^i . We have $d\theta^i = d(A^{-1})_k^i \wedge dx^i$. In order to express $d(A^{-1})_k^i$ in

the terms of transition matrices we differentiate the both sides of the matrix relation $A^{-1}A = I$, where I is the unit matrix. We get

$$d(A^{-1})A + A^{-1}dA = 0 \implies d(A^{-1}) = -A^{-1}dAA^{-1}.$$

Thus

$$d\theta^{i} = -\left((A^{-1})^{i}_{k} dA^{k}_{m} (A^{-1})^{m}_{j}\right) \wedge dx^{j}$$
$$= -\left((A^{-1})^{i}_{k} dA^{k}_{m}\right) \wedge \left((A^{-1})^{m}_{j} dx^{j}\right)$$
$$= -\omega^{i}_{m} \wedge \theta^{m}.$$

The equation

$$d\theta^i = \theta^m \wedge \omega_m^i$$

is called the first Cartan equation in n-dimensional affine space E^n .

Now our aim is to express the exterior differential of the connection 1-forms in terms of dual 1-forms θ^i . We have $\omega_j^i = (A^{-1})_k^i dA_j^k$ and differentiating the both sides we obtain

$$\begin{aligned} d\omega_{j}^{i} &= -d(A^{-1})_{k}^{i} \wedge dA_{j}^{k} = -\left((A^{-1})_{m}^{i} \, dA_{l}^{m} \, (A^{-1})_{k}^{l}\right) \wedge dA_{j}^{k} \\ &= -\left((A^{-1})_{m}^{i} \, dA_{l}^{m}\right) \wedge \left((A^{-1})_{k}^{l} dA_{j}^{k}\right) \\ &= -\omega_{l}^{i} \wedge \omega_{j}^{l}. \end{aligned}$$

The equation

$$d\omega_j^i = \omega_j^l \wedge \omega_l^i,$$

is called the second Cartan equation in n-dimensional affine space E^n .

The Euclidean structure of the *n*-dimensional affine space E^n makes it possible to single out a very convenient subclass in the class of curvilinear coordinates. This subclass is called orthogonal curvilinear coordinate systems. In such coordinates the coordinate lines are pairwise orthogonal, which means that $\langle D_i, D_j \rangle = 0$ for every pair $i \neq j$. We can rescale the basic vector fields D_i introducing new vector fields

$$\mathbf{E}_i = \frac{1}{\|\mathbf{D}_i\|} \, \mathbf{D}_i, \quad \|\mathbf{E}_i\| = 1,$$

which form the orthonormal basis of the tangent space $T_p E^n$ at each point $p \in U$. Let us find what changes in the previous considerations if we use the set of vector fields $\{\mathbf{E}_i\}$ as the basic vector fields. First of all we note that in the case when we use the Euclidean metric of E^n , it makes no sense to distinguish between upper and lower indices. Indeed, in this case the space is canonically isomorphic to its dual space. Thus we will use only subscripts while the superscript will be put on the first place. The matrix $A = (A_{ji})$ in $\mathbf{E}_i = A_{ji} \mathbf{U}_j$ is a special orthogonal matrix, i.e. $A^T A = I$, Det A = 1 (we assume that a system of curvilinear coordinates has the same orientation as Cartesian coordinates). The condition of orthogonality can be written in the form $A = (A^{-1})^T$. Differentiating both sides of $A^T (A^{-1})^T = I$ we get

$$dA^{T} (A^{-1})^{T} + A^{T} d(A^{-1})^{T} = 0 \implies dA^{T} (A^{-1})^{T} = -A^{T} dA$$

Now we calculate

$$\begin{split} \omega + \omega^T &= A^{-1} \, dA + (A^{-1} \, dA)^T = A^{-1} \, dA + dA^T \, (A^{-1})^T \\ &= A^{-1} \, dA - A^T dA = A^{-1} \, dA - A^{-1} \, dA = 0. \end{split}$$

Hence the matrix of connection ω calculated in the orthonormal basis $\{E_i\}$ is a skewsymmetric matrix. If one looks at the derivation of Cartan equations given above then it can be seen that the Cartan equations will have the same form in the case of orthonormal vector fields $\{E_i\}$.

Example 2.1. As an example, we will consider the parabolic coordinates (u, w, v) in three-dimensional Euclidean space E^3 . The Cartesian coordinates (x, y, z) can be expressed in the terms of parabolic coordinates follows

$$\begin{cases} x = 2 u w \cos v, \\ y = 2 u w \sin v, \\ x = u^2 - w^2. \end{cases}$$
(2.63)

We find

$$\begin{split} \mathsf{D}_1 &= & \frac{\partial}{\partial u} = 2w\cos v \; \mathsf{U}_1 + 2w\sin v \; \mathsf{U}_2 + 2u \; \mathsf{U}_3, \\ \mathsf{D}_2 &= & \frac{\partial}{\partial w} = 2u\cos v \; \mathsf{U}_1 + 2u\sin v \; \mathsf{U}_2 - 2w \; \mathsf{U}_3, \\ \mathsf{D}_3 &= & \frac{\partial}{\partial v} = -2 \; uw\sin v \; \mathsf{U}_1 + 2 \; uw\cos v \; \mathsf{U}_2, \end{split}$$

By straightforward computation we find the matrix of Euclidean metric g in parabolic



coordinates

$$\begin{split} g_{11} &= < \mathsf{D}_1, \mathsf{D}_1 > = (\frac{\partial x}{\partial u})^2 + (\frac{\partial y}{\partial u})^2 + (\frac{\partial z}{\partial u})^2 = 4(u^2 + w^2), \\ g_{22} &= < \mathsf{D}_2, \mathsf{D}_2 > = (\frac{\partial x}{\partial w})^2 + (\frac{\partial y}{\partial w})^2 + (\frac{\partial z}{\partial w})^2 = 4(u^2 + w^2), \\ g_{33} &= < \mathsf{D}_3, \mathsf{D}_3 > = (\frac{\partial x}{\partial v})^2 + (\frac{\partial y}{\partial v})^2 + (\frac{\partial z}{\partial v})^2 = 4u^2w^2, \end{split}$$

and $g_{ij} = \langle D_i, D_j \rangle = 0$ for $i \neq j$. We see that the system of parabolic coordinates is
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an orthogonal system of curvilinear coordinates. Hence we can construct a system of orthonormal vector fields $\{E_1,E_2,E_3\}$ if we define

$$\mathbf{E}_1 = \frac{1}{\sqrt{g_{11}}} \mathbf{D}_1, \ \mathbf{E}_2 = \frac{1}{\sqrt{g_{22}}} \mathbf{D}_2, \ \mathbf{E}_3 = \frac{1}{\sqrt{g_{33}}} \mathbf{D}_3.$$
 (2.64)

The dual 1-forms $\theta_1, \theta_2, \theta_3$ for the orthonormal vector fields $\{E_1, E_2, E_3\}$ are defined by

$$\theta_1 = \sqrt{g_{11}} \, du, \ \ \theta_2 = \sqrt{g_{22}} \, dw, \\ \theta_3 = \sqrt{g_{33}} \, dv.$$
 (2.65)

Hence

$$\begin{aligned} \mathbf{E}_{1} &= \frac{w\cos v}{\sqrt{u^{2}+w^{2}}} \, \mathbf{U}_{1} + \frac{w\sin v}{\sqrt{u^{2}+w^{2}}} \, \mathbf{U}_{2} + \frac{u}{\sqrt{u^{2}+w^{2}}} \, \mathbf{U}_{3}, \quad \theta_{1} = 2\sqrt{u^{2}+w^{2}} \, du \\ \mathbf{E}_{2} &= \frac{u\cos v}{\sqrt{u^{2}+w^{2}}} \, \mathbf{U}_{1} + \frac{u\sin v}{\sqrt{u^{2}+w^{2}}} \, \mathbf{U}_{2} - \frac{w}{\sqrt{u^{2}+w^{2}}} \, \mathbf{U}_{3}, \quad \theta_{1} = 2\sqrt{u^{2}+w^{2}} \, dw \\ \mathbf{E}_{3} &= -\sin v \, \mathbf{U}_{1} + \cos v \, \mathbf{U}_{2}, \quad \theta_{3} = 2 \, uw \, dv. \end{aligned}$$

$$\end{aligned}$$

Now the transition matrix A has the form

$$\begin{pmatrix} \frac{w\cos v}{\sqrt{u^2 + w^2}} & \frac{u\cos v}{\sqrt{u^2 + w^2}} & -\sin v\\ \frac{w\sin v}{\sqrt{u^2 + w^2}} & \frac{u\sin v}{\sqrt{u^2 + w^2}} & \cos v\\ \frac{u}{\sqrt{u^2 + w^2}} & -\frac{w}{\sqrt{u^2 + w^2}} & 0 \end{pmatrix}.$$
 (2.67)

It is easy to check that A is a special orthogonal matrix, i.e. $A \in SO(3)$. Thus we can find the reciprocal matrix A^{-1} by taking transposed matrix A^T . By straightforward calculations we find the elements of the matrix dA. We will present them by columns. So the elements of the first column are

$$dA_{11} = -\frac{uw\cos v}{(u^2 + w^2)^{3/2}}du + \frac{u^2\cos v}{(u^2 + w^2)^{3/2}}dw - \frac{w\sin v}{(u^2 + w^2)^{1/2}}dv,$$

$$dA_{21} = -\frac{uw\sin v}{(u^2 + w^2)^{3/2}}du + \frac{u^2\sin v}{(u^2 + w^2)^{3/2}}dw + \frac{w\cos v}{(u^2 + w^2)^{1/2}}dv,$$

$$dA_{31} = \frac{w^2}{(u^2 + w^2)^{3/2}}du - \frac{uw}{(u^2 + w^2)^{3/2}}dw.$$

The elements of the second column are

$$dA_{12} = \frac{w^2 \cos v}{(u^2 + w^2)^{3/2}} du - \frac{uw \cos v}{(u^2 + w^2)^{3/2}} dw - \frac{u \sin v}{(u^2 + w^2)^{1/2}} dv,$$

$$dA_{22} = -\frac{w^2 \sin v}{(u^2 + w^2)^{3/2}} du - \frac{uw \sin v}{(u^2 + w^2)^{3/2}} dw + \frac{u \cos v}{(u^2 + w^2)^{1/2}} dv,$$

$$dA_{32} = \frac{uw}{(u^2 + w^2)^{3/2}} du - \frac{u^2}{(u^2 + w^2)^{3/2}} dw.$$

Finally the elements of the third column are

$$dA_{13} = -\cos v \, dv,$$

$$dA_{23} = -\sin v \, dv,$$

$$dA_{33} = 0.$$

Making use of the formula $\omega = A^T dA$ we find the connection 1-forms

$$\omega_{12} = \frac{w du - u dw}{u^2 + w^2}, \quad \omega_{23} = -\frac{u dv}{\sqrt{u^2 + w^2}}, \quad \omega_{31} = \frac{w dv}{\sqrt{u^2 + w^2}}$$

2.6 INTEGRAL CURVES OF A VECTOR FIELD

Assume we are given a vector field X in an open subset $U \subset E^n$. Fix a point $p \in U$. Now we can put an interesting question of how to find a curve in U, which passes a point p, such that its velocity vector at any point of a curve is the value of a vector field X at this point. By other words, velocity vector of unknown curve at any of its points is determined by a given vector field X. This curve is called an integral curve of a vector field X through a point p and in this section we will show that integral curve of any smooth vector field exists and its uniquely determined by a point p. In this section we will use Cartesian coordinates of Euclidean space E^n , but actually the results do not depend on a choice of coordinate system and can be applied in any coordinate system.

Let X^1, X^2, \ldots, X^n be components of a vector field X in Cartesian coordinates of E^n . We consider the problem of finding a parametrized curve $(\xi = \xi(t), I)$ such that it passes a point p, whose coordinates will be denoted by p^1, p^2, \ldots, p^n , and its velocity vector at any point is determined by the value of a vector field X at this point. We will also assume that I contains the number zero and a parametrized curve $\xi = \xi(t)$ passes a point p, when t = 0, i.e. $\xi(0) = p$. In order to write down an equation for unknown parametrized curve $\xi = \xi(t)$, we restrict a vector field X to a parametrized curve $\xi = \xi(t)$, that is, $X(t) = X|_{\xi(t)}$. Clearly, X(t) is the vector field along a parametrized curve $\xi = \xi(t)$. Now, the condition, that velocity vector of a parametrized curve $\xi = \xi(t)$ is determined by a vector field X, can be written in the form of equation

$$\boldsymbol{\xi}'(t) = \boldsymbol{X}(t). \tag{2.68}$$

Let $\xi(t) = (\xi^1(t), \xi^2(t), \dots, \xi^n(t))$. Then the equation (2.68), written in coordinates, leads us to the following system of differential equations of first order

$$\frac{d\xi^{1}}{dt} = \mathbf{X}^{1}(\xi^{1}(t), \xi^{2}(t), \dots, \xi^{n}(t)),
\frac{d\xi^{2}}{dt} = \mathbf{X}^{2}(\xi^{1}(t), \xi^{2}(t), \dots, \xi^{n}(t)),
\dots \dots \dots \qquad (2.69)
\frac{d\xi^{n}}{dt} = \mathbf{X}^{n}(\xi^{1}(t), \xi^{2}(t), \dots, \xi^{n}(t)).$$

In this system of differential equations, $\xi^1(t), \xi^2(t), \ldots, \xi^n(t)$ are unknown functions, $X^i(\xi^1(t), \xi^2(t), \ldots, \xi^n(t))$, where $i = 1, 2, \ldots, n$, are given functions and $\xi^i(0) = p^i$ are initial conditions.

A parametrized curve $(\xi = \xi(t), I)$, which satisfies the equation (2.68) or the system of differential equations of first order (2.69) with initial conditions $\xi^i(0) = p^i$, is called an *integral curve* of a vector field **X** through a point p.

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It follows from the theory of differential equations [4] that there exists the solution $(\xi = \xi(t), I)$ of the system of first order equations (2.69), this solution is uniquely determined by initial conditions $\xi^i(0) = p^i$, this solution is smooth if all functions X^1, X^2, \ldots, X^n are smooth functions and it is maximal, that is, if $(\eta = \eta(t), J)$ is also a solution of (2.69) with the same initial conditions, then $J \subset I$ and $\eta(t) = \xi(t)|_J$. Geometrically this means that a vector field X generates the family of parametrized curves in an open subset U such that through each point of U passes one and only one parametrized curve. We will call this family of parametrized curves a flow of integral curves generated by a vector field X.

It is clear that each parametrized curve of this flow, generated by a vector field X, depends not only on parameter t, but also on initial conditions $\xi^i(0) = p^i$. In order to emphasize this dependence, we will denote the solution of (2.69), which satisfies the initial condition $\xi^i(0) = p^i$, by $\xi(p;t) = (\xi^1(p;t),\xi^2(p;t),\ldots,\xi(p;t))$. Thus $\xi^i(p;0) = p^i$. A peculiar property of the flow of integral curves, generated by X, is that it has a structure of a group. Indeed, it follows from the existence and uniqueness of solution $\xi(p;t)$ of the system of differential equations (2.69) that for any point $p \in U$ and $t, s, t + s \in I$, we have

$$\xi(\xi(p;t),s) = \xi(p;t+s).$$
(2.70)

Particularly, for any $t \in I$ such that $-t \in I$, it holds $\xi(\xi(p;t), -t) = \mathrm{id}_U$. Define the family of transformations $g_t : U \to U$ by $g_t(p) = \xi(p;t)$. As any solution of the system (2.69) depends smoothly on initial conditions, a transformation $g_t : U \to U$ for any $t \in I$ is a smooth transformation of U. Then, from (2.70), it follows

$$g_s \circ g_t = g_{t+s}, \ g_{-t} = g^{-1}(t), \ g_0 = \mathrm{id}_U.$$
 (2.71)

The property $g_{-t} = g^{-1}(t)$ shows that each transformation g_t has its inverse, which is also smooth. Thus, g_t is a diffeomorphism of an open subset U for each $t \in I$. Next, the family of diffeomorphisms $g_t : U \to U$, generated by a vector field X, is an Abelian group, where the group operation is a product of diffeomorphisms. This group is called a *one-parameter group of diffeomorphisms of an open subset* U of Euclidean space E^n , generated by a vector field X.

2.7 CONNECTION IN CURVILINEAR COORDINATES

In this section, we will elaborate differential approach, which can be used to measure a change of one vector field Y in the direction determined by another vector field X. This will lead us to such important concepts as the covariant derivative and the covariant differential of a vector field. The peculiarity of the approach in this section is that we will work in curvilinear coordinates. We will first develop this differential calculus at one point of E^n , and then we will extend it to the whole space.

Let us fix an open subset $U \subset E^n$ and two vector fields X, Y defined on U. We assume that U is equipped with a coordinate system $(x^i) = (x^1, x^2, \ldots, x^n)$ and this coordinate system is not necessarily Cartesian, that is, it can also be curvilinear. A coordinate system (x^i) induces the set of basic vector fields $\partial/\partial x^i$, $i = 1, 2, \ldots, n$,

which will be denoted by E_1, E_2, \ldots, E_n . Let us fix a point $p \in U$ and a parametrized curve $\xi: I \to E^n$, which passes through a point p, i.e. $\xi(t_0) = p, t_0 \in I$ and its velocity vector at this point $\xi'(t_0)$ is the value of a vector field X at this point, that is, $\xi'(t_0) = X_p$. In order to simplify notations we will denote this vector by v_p , i.e. $v_p = X_p = \xi'(t_0)$. Our goal is to develop a differential approach to measure a change of a vector field Y at a point p in the direction of a vector field X. Obviously, this change (in infinitesimal) is described by the derivative of a vector field Y along a curve ξ . We can define this derivative in an invariant form, that is, a form that does not use any coordinate system. To do this, we need the increment of a vector field Ywhen shifting to an infinitely close point along the curve. Thus, we should consider the difference $\Upsilon(\xi(t_0 + \Delta t)) - \Upsilon(\xi(t_0))$. However, this is too straightforward and not correct. The fact is that vector $\mathbf{Y}(\xi(t_0 + \Delta t))$ has its origin at point $\xi(t_0 + \Delta t)$, while a vector $Y(\xi(t_0))$ has its origin at point $\xi(t_0)$. Thus, vectors are applied to different points in space E^n and we cannot perform addition or subtraction operations on such vectors. We can solve this problem by dropping an initial point of a vector, that is, by going to its vector part, which is a vector-valued function \vec{Y} (at each point $x \in U$ we have $\mathbf{Y}_x = (x; \mathbf{\vec{Y}}(x))$. In this case, the increment of a vector field Y will be described by the correctly defined difference $\vec{Y}(\xi(t_0 + \Delta t)) - \vec{Y}(\xi(t_0))$. However, at this point it is useful to reflect on a geometric essence of what we have done. Indeed, we can interpret dropping an initial point of a vector as a parallel translation of vector $\vec{Y}(\xi(t_0 + \Delta t))$ from point $\xi(t_0 + \Delta t)$ to point $\xi(t_0)$ (in the sense of Euclidean space E^n) and then subtracting vector $\dot{Y}(\xi(t_0))$. This interpretation is very important and it indicates a deep connection between a derivative we are developing with parallel translation of vectors in Euclidean space E^n . Since our goal is to apply the resulting derivative to vector fields, we must get a vector at a point. It is easy to do as follows. We will first calculate the derivative of a vector field Y along a curve ξ (at a point p), and then apply the resulting vector to point p. Let us denote $\vec{Y}(\xi(t)) = \vec{Y}(t)$. Then we define a covariant derivative of a vector field Y at a point p in direction of X by the formula

$$D_{\mathbf{v}_p}Y = \left(p; \lim_{\Delta t \to 0} \frac{1}{\Delta t} \left(\vec{\mathbf{Y}}(t_0 + \Delta t) - \vec{\mathbf{Y}}(t_0)\right)\right) = \left(p; \frac{d}{dt}\vec{\mathbf{Y}}(t)\Big|_{t=t_0}\right).$$
(2.72)

This definition does not depend on a choice of a coordinate system in U. The next step in the development of a covariant derivative is its calculation in coordinates (x^i) .

EXERCISES

2.1 Define the vector field X in Euclidean plane E^2 by its components as follows: $X^1(x,y) = -y, X^2(x,y) = x$, where x, y are Cartesian coordinates of a point of E^2 . Thus, $X = -y U_1 + x U_2$. Show that the integral curves of this vector field are the circles with the center at the origin and this vector field generates the group of rotations of Euclidean plane E^2 .

2.2

2.3 .

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2.4 . Solution:

2.5 .

Surfaces in 3-dimensional Euclidean space

3.1 HOW TO DEFINE A SURFACE?

In this section, we will consider surfaces in three-dimensional Euclidean space E^3 . Unlike a curve, a surface has its own internal geometry. Figuratively speaking, we could imagine two-dimensional intellectual beings for whom the surface is their universe and who are unaware of the existence of a third dimension, that is, they do not suspect that their universe is in a space of a higher dimension. These beings measure the lengths of lines on the surface, the angles between the lines, that is, they build the internal geometry of a surface without even knowing that this geometry is induced by the Euclidean geometry of the space enclosing their surface.

As before, we assume that in three-dimensional Euclidean space E^3 we have an orthonormal frame $\{O; \vec{e_1}, \vec{e_2}, \vec{e_3}\}$, which defines the Cartesian coordinate system, whose coordinates will be denoted by x, y, z. How can we define a surface analytically, that is, using an equation? Basically we have two different methods to define a surface by means of equation.

- S1) The first one is a method of parametric representation of surfaces by means of two parameters. This method is similar to a parametric representation of a curve, and we used this method in the case of curves in the previous sections.
- S2) The second method is a representation of a surface by an equation f(x, y, z) = c, where f(x, y, z) is a smooth function and c is a constant real number. This method is called implicit representation of a surface, since in this case the points of a surface are solutions to equation f(x, y, z) = c.

We start with a parametric representation of a surface, then describe a method for implicitly representing a surface and point out their relationship. A surface, unlike a curve, which is one-dimensional, is a two-dimensional geometric object and therefore we need two parameters to define a surface parametrically. For this purpose we will consider a Euclidean plane E^2 , which will provide a domain for a surface parameters. As before, we assume that Euclidean plane E^2 is equipped with a Cartesian coordinate system, which we will denote by u, v and these coordinates will be the parameters of

a surface. It is convenient to consider a plane E^2 as existing separately from space E^3 . In other words a plane E^2 is not a plane embedded in space E^3 . To emphasize that plane E^2 is a plane of surface parameters, we will call it a *uv*-plane. Let $U \subset E^2$ be an open subset of a plane E^2 . We will map a subset U to three-dimensional Euclidean space E^3 and in this mapping it will take the shape of a surface. In order to map U into three-dimensional space, we need three functions, let us denote them by x(u, v), y(u, v), z(u, v). We will assume that domain of these functions is an open subset U and these functions are smooth (infinitely differentiable). The mapping determined by functions x(u, v), y(u, v), z(u, v) will be denoted by ϕ , i.e. $\phi : U \to E^3$, where $\phi(u, v) = (x(u, v), y(u, v), z(u, v))$. If we use vector notation $\phi(u, v)$, then this means that $\phi(u, v)$ is the position vector of a point $\phi(u, v)$.

Definition 3.1. A parametrized or local surface in three-dimensional Euclidean space E^3 is a smooth mapping

$$\phi: (u,v) \in U \to \phi(u,v) = (x(u,v), y(u,v), z(u,v)) \in E^3,$$

where $U \subset E^2$ is an open subset of a plane E^2 (domain of parameters). The equation $\phi(u, v) = (x(u, v), y(u, v), z(u, v))$ will be referred to as a parametric equation of a parametrized surface. If $\phi : U \to E^3$ is an injective mapping then ϕ will be called a parametrized coordinate surface. The image of a mapping ϕ will be denoted by \mathfrak{P} , *i.e.* $\mathfrak{P} = Im \phi \subset E^3$.

A parametrized surface will be denoted as a pair (ϕ, U) . The assumption that a mapping $\phi(u, v) = (x(u, v), y(u, v), z(u, v))$ is smooth not yet a guarantee that the image of a subset U under the mapping ϕ will be a smooth surface in E^3 in the sense that at each point $\phi(u, v) \in E^3$ we will have a tangent plane to a surface. But in order to study the shape of a surface in a neighborhood of one of its points we will use a tangent plane. Hence a tangent plane to a surface plays an important role in study of a structure of this surface by methods of differential calculus. Before we formulate an additional condition that will ensure the existence of a tangent plane to a surface at a given point, we must define a concept of a tangent vector to a parametrized surface. Let $q = (u_0, v_0) \in U$ be a point of domain of surface parameters. Then $p = \phi(u_0, v_0) \in E^3$ is the corresponding point on a parametrized surface (ϕ, U) in E^3 . Let $(\xi = \xi(t), I)$ be a parametrized curve in U such that it passes through a point q. Hence if $\xi(t) = (u(t), v(t))$ is a parametric equation of a curve ξ then $\xi(t_0) = q, t_0 \in I$, i.e. $u(t_0) = u_0, v(t_0) = v_0$. Then ϕ maps a parametrized curve ξ to a parametrized curve $\tilde{\xi} = \phi \circ \xi : I \to E^3$ on a parametrized surface ϕ . Obviously the parametrized curve $\tilde{\xi}(t)$ passes through the point p, that is, $\tilde{\xi}(t_0) = p$. The velocity vector $\tilde{\xi}'|_p$ of the parametrized curve $\tilde{\xi}$ calculated at a point p is called a *tangent* vector of a parametrized surface (ϕ, U) at a point p. The set of all tangent vectors to a parametrized surface (ϕ, U) at a point $p = \phi(u, v)$ will be denoted by $T_p \phi$.

As in the case of curves, it is useful to define a notion of a vector field on a parametrized surface. A vector field X on a parametrized surface (ϕ, U) is a mapping which attaches a 3-dimensional vector to each point of a parametrized surface. From this it follows that any vector field X on a parametrized surface (ϕ, U) can be written

in the form

$$\mathbf{X}(u,v) = (\phi(u,v); X^1(u,v), X^2(u,v), X^3(u,v)),$$
(3.1)

where $u, v \in U \subset E^2$ and $X^1(u, v), X^2(u, v), X^3(u, v)$ are smooth functions. In what follows, the case when a vector field **X** on a parametrized surface (ϕ, U) is a tangent vector at every point of this surface, i.e. $\mathbf{X}_p \in T_p \phi$ for any $p \in \text{Im } \phi$, will be very important. In this case a vector field on a parametrized surface will be called a *tangent* vector field.

As an example, we can consider curves on a parametrized surface (ϕ, U) that are induced by the coordinate lines of uv-plane E^2 . Let us choose a value of the second coordinate v and let the first coordinate u change in some sufficient small interval around a value u, that is, u(t) = u + t, where $-\delta < t < \delta$. We get the straight line that passes through point q = (u, v) of uv-plane E^2 and is called the u-coordinate line of a plane E^2 . Its parametric equation is $\xi_u(t) = (u + t, v)$. Here, by a sufficiently small value δ , we mean such a value of t for which the straight line $\xi_u(t)$ does not go beyond the set U. Analogously we construct the v-coordinate line, whose parametric equation is $\xi_v(t) = (u, v + t)$. Parametrized curves $\tilde{\xi}_u = \phi \circ \xi_u$ and $\tilde{\xi}_v = \phi \circ \xi_v$ lie on a parametrized surface ϕ and they pass through a point $p = \phi(u, v)$. A parametrized curve ξ_u is called *u*-line on a parametrized surface and ξ_v is called v-line. Since q = (u, v) is an arbitrary point of a domain U we cover a parametrized surface (ϕ, U) with a grid of coordinate lines. Moreover, one u-line and one v-line pass through each point of a parametrized surface (ϕ, U) . Hence we have the coordinate system of u-lines and v-lines on a parametrized surface (ϕ, U) . This coordinate system is induced by the Cartesian coordinate system of the uv-plane E^2 . If at each point of a parametrized surface the u-line is orthogonal to the v-line then we will call the coordinate system of *u*-lines and *v*-lines orthogonal. According to the previously defined notion of a tangent vector to a parametrized surface the velocity vectors of u-coordinate line and v-coordinate line of a parametrized surface are the tangent vectors of a parametrized surface. In order to find the velocity vector of u-line we have to differentiate its parametric equation $\xi_u(t) = (x(u+t, v), y(u+t, v), z(u+t, v))$ with respect to t. But in this case it will give us the partial derivatives of functions x(u, v), y(u, v), z(u, v) with respect to parameter u. If we denote this velocity vector by ϕ'_u then

$$\phi'_u(u,v) = (\phi(u,v); x'_u(u,v), y'_u(u,v), z'_u(u,v)),$$
(3.2)

where x'_u, y'_u, z'_u stand for partial derivatives with respect u. Analogously we can show that the velocity vector of the v-line has the form

$$\phi'_{v}(u,v) = (\phi(u,v); x'_{v}(u,v), y'_{v}(u,v), z'_{v}(u,v)).$$
(3.3)

Comparing equations (3.2) and (3.3) with (3.1) we see that tangent vectors of coordinate lines give rise to tangent vector fields on a parametrized surface and we will denote these two tangent vector fields by ϕ'_u and ϕ'_v .

Let us now consider an arbitrary tangent vector $\tilde{\xi}'(t)$ at a point p of a parametrized surface ϕ . Here we assume that $\xi(t) = (u(t), v(t))$ is a parametrized curve in the uvplane E^2 and $\tilde{\xi} = \phi \circ \xi$. Then we have

$$\tilde{\xi}' = (\phi(u, v); x'_t(u, v), y'_t(u, v), z'_t(u, v)).$$
(3.4)

Making use of differentiation of a composite functions x(u(t), v(t)), y(u(t), v(t)), z(u(t), v(t)), we get

$$\begin{array}{rcl} x'_t(u,v) &=& x'_u \, u'_t + x'_v \, v'_t, \\ y'_t(u,v) &=& y'_u \, u'_t + y'_v \, v'_t, \\ z'_t(u,v) &=& z'_u \, u'_t + z'_v \, v'_t. \end{array}$$

Substituting these into (3.4) we obtain

$$\tilde{\xi}' = u'_t \phi'_u|_{\xi = \xi(t)} + v'_t \phi'_v|_{\xi = \xi(t)}.$$
(3.5)

The obtained formula is very important. It shows that any tangent vector of a parametrized surface ϕ is a linear combination of tangent vectors ϕ'_u, ϕ'_v and the coefficients of this linear combination are the coordinates of the velocity vector ξ' of a parameterized curve $\xi(t) = (u(t), v(t))$ in the *uv*-plane E^2 . Hence $T_p \phi \subset \text{span} \{\phi'_u, \phi'_v\}$.

Another very important conclusion follows from this. A parametrized surface will have the tangent plane at a given point of a surface if the tangent vectors $\phi'_u(u,v), \phi'_v(u,v)$ are linearly independent at this point. As it follows from (3.5) in this case the vectors form the basis for the tangent plane. Hence the plane $T_p \phi$ spanned by the tangent vectors $\phi'_u(u,v), \phi'_v(u,v)$ will be called the *tangent plane* of a parametrized surface ϕ at a point $p = \phi(u,v)$. A condition for the linear independence of tangent vectors $\phi'_u(u,v), \phi'_v(u,v)$ can be given in the following two equivalent ways:

R1) the tangent vectors $\phi'_u(u, v), \phi'_v(u, v)$ are linearly independent if and only if

$$\phi'_u(u,v) \times \phi'_v(u,v) \neq 0,$$

R2) the tangent vectors $\phi_u'(u,v), \phi_v'(u,v)$ are linearly independent if and only if

$$\operatorname{rank}\left(\begin{array}{cc} x'_u & y'_u & z'_u \\ x'_v & y'_v & z'_v \end{array}\right) = 2.$$

If at each point of a parametrized surface one of these equivalent conditions is satisfied, we will call such a surface *regular*. In what follows, we will always assume that the regularity condition is satisfied and we will omit the word "regular" simply writing "parametrized surface" or "local surface".

The first regularity condition has a simple geometric meaning. It shows that at each point a surface has a non-zero normal vector (orthogonal to a tangent plane). Since we assume that a regularity condition for a parametrized surface is satisfied, a parametrized surface ϕ always has the unit normal vector field N defined by the formula

$$\mathbf{N} = \frac{\phi'_u \times \phi'_v}{||\phi'_u \times \phi'_v||}.\tag{3.6}$$

Example 3.1. As an example of a parameterized surface, we will consider a surface of revolution. Surfaces of revolution form a very important class of surfaces and many well-known surfaces, such as a sphere, torus, catenoid, are surfaces of revolution. The

idea of a surface of revolution is that we fix a plane in space, draw a straight line (axis of rotation) on it and take any curve that lies on a plane on one side of a straight line, without crossing or touching it. Then we start to rotate this curve around a straight line. If, with this rotation, the curve left a trace in space, we would see the surface. This is the surface of revolution.

We will choose xy-coordinate plane as a plane, and x-axis as a straight line. A curve in the xy-coordinate plane will be represented in a parametrized form ($\xi = \xi(t), I$) by a parametric equation $\xi(t) = (f(t), h(t))$. Note that we are considering this curve as a curve in xy-coordinate plane, not in space. Therefore, we omit third coordinate in its parametric equation. We assume that curve ξ lies entirely in the upper half-plane of xy-coordinate plane. Consequently y(t) > 0 for any $t \in I$, i.e. there is no crossing and touching. It should be recalled here that we assume that the curve regularity condition is satisfied, that is, $(f'(t))^2 + (h'(t))^2 > 0$. As the parameters of a surface of revolution u, v, it is convenient to choose the curve ξ parameter t as parameter u and the angle of rotation as parameter v. Then the parametric equation of a surface of revolution has the form

$$\phi(u, v) = (f(u), h(u) \cos v, h(u) \sin v), \quad U = I \times [0, 2\pi]. \tag{3.7}$$

This equation is derived from simple considerations. First of all, it is easy to see that when a point of curve $\xi(t) = (f(t), h(t))$ rotates around x-axis, the first coordinate f(t) of a point does not change. Therefore, in the parametric equation (3.7), the first function is f(u). In a plane parallel to the yz-coordinate plane, a point moves along a circle with a radius of h(u). Therefore, the second and third functions in the parametric equation have the form of a parametrization of a circle with a radius of h(u).

We find

$$\begin{split} \vec{\phi}'_u &= (f', h' \cos v, h' \sin v), \\ \vec{\phi}'_v &= (0, -h \sin v, h \cos v), \\ \vec{\phi}'_u \times \vec{\phi}'_v &= (h h', -f' h \cos v, -f' h \sin v), \\ |\phi'_u \times \phi'_v|| &= h \sqrt{(f')^2 + (h')^2}. \end{split}$$

Since h(t) > 0 and $(f'(t))^2 + (h'(t))^2 > 0$ it follows that $\|\phi'_u \times \phi'_v\| > 0$ and (3.7) is a regular parametrization of a surface of revolution. Finally we find the unit normal vector field on a surface of revolution

$$\vec{N} = \frac{1}{\sqrt{(f')^2 + (h')^2}} \left(h', -f' \cos v, -f' \sin v \right)$$

Let us now consider the second way of representing a surface in three-dimensional Euclidean space, the so-called implicit way of representing a surface. Although this method is based on equation f(x, y, z) = c, the set of solutions of which gives a set of surface points, we start with a more general definition, which later leads to the concept of manifold.

Let $\mathfrak{S} \subset E^3$ be a subset of a three-dimensional Euclidean space E^3 .

Definition 3.2. A subset \mathfrak{S} of a three-dimensional Euclidean space E^3 is called a surface if for any point $p \in \mathfrak{S}$ there exists a neighborhood V of this point in E^3 and a local coordinate surface (ϕ, U) such that $\phi : U \to \mathfrak{S} \cap V$ is onto mapping.

The geometric meaning of this definition is that a surface consists of several local coordinate surfaces (sometimes called patches), which in some places overlap (places of the so-called gluing) and together they form a surface. This definition says what the structure of a set called a surface should be, but it does not say anything about how a surface in three-dimensional space E^3 can be represented. Therefore, we will consider a method to represent a surface by means of equation f(x, y, z) = c and prove that, under a certain condition on a function f, the set of solutions to this equation forms a surface in the sense of Definition 3.2.

Let $W \subset E^3$ be an open subset, $f \in C^{\infty}(W)$ be a smooth function and $c \in \mathbb{R}$ be a real number. We consider the equation f(x, y, z) = c. The set of all solutions of this equation will be denoted by $f^{-1}(c)$, that is, this set consists of all points in space E^3 at which the value of a function f is equal to c. We will assume that this set is non-empty, i.e. an equation f(x, y, z) = c has at least one solution and call this set *level surface* of a function f. Obviously $f^{-1}(c) \subset W$.

Proposition 3.1. Let $\mathfrak{S} = f^{-1}(c)$ be a level surface of a function f and $p \in \mathfrak{S}$ be a point of this level surface such that the gradient of a function f is non-zero at p, that is, $\nabla f|_p \neq 0$. If $\xi : I \to \mathfrak{S}$ is a parametrized curve on a level surface \mathfrak{S} , which passes through a point p, that is, $\xi(t_0) = p, t_0 \in I$ then the velocity vector $\xi'|_p$ of this curve at a point p is orthogonal to the gradient $\nabla f|_p$.

Proof. Let $\xi = \xi(t) = (x(t), y(t), z(t)), \xi(t_0) = p$. Since a parametrized curve $\xi(t)$ lies on a level surface \mathfrak{S} we have $f(\xi(t)) = c$ for any $t \in I$. Differentiating this equation with respect to t and putting $t = t_0$ we get

$$\frac{d}{dt}f(\xi(t))\Big|_{t=t_0} = \frac{\partial f}{\partial x}\Big|_p \frac{dx}{dt}\Big|_{t=t_0} + \frac{\partial f}{\partial y}\Big|_p \frac{dy}{dt}\Big|_{t=t_0} + \frac{\partial f}{\partial z}\Big|_p \frac{dz}{dt}\Big|_{t=t_0} = 0.$$
(3.8)

Taking into account that

$$\xi'|_p = (p; \frac{dx}{dt}\Big|_{t=t_0}, \frac{dy}{dt}\Big|_{t=t_0}, \frac{dz}{dt}\Big|_{t=t_0}),$$

is the velocity vector of a curve ξ at a point p we can consider the middle part of (3.8) as a scalar product of two vectors, where first is the vector of gradient

$$\nabla f|_p = (p; \frac{\partial f}{\partial x}\Big|_p, \frac{\partial f}{\partial y}\Big|_p, \frac{\partial f}{\partial z}\Big|_p),$$

and the second is the velocity vector $\xi'|_p$. Hence

$$\langle \nabla f|_p, \xi'|_p \ge 0.$$

Thus $\nabla f|_p \perp \xi'|_p$.

The proved proposition shows that a tangent vector to a level surface $f^{-1}(c)$ at a point $p \in f^{-1}(c)$ is orthogonal to the gradient of a function f at this point. Thus, the gradient of a function f can be considered as a normal vector to a level surface $f^{-1}(c)$ at some point, and the plane, which is orthogonal to the gradient ∇f at this point, can be considered as a tangent plane to a level surface. We will denote a tangent plane of a level surface $\mathfrak{S} = f^{-1}(c)$ at a point p by $T_p \mathfrak{S}$. Thus

$$T_p\mathfrak{S} = \{\mathbf{v}_p \in T_p E^3 : < \mathbf{v}_p, \nabla f|_p >= 0\}.$$

Thus, we see that the condition for the existence of a tangent plane to a level surface of a function f is that the gradient of this function differs from zero. Therefore, a point on a level surface $f^{-1}(c)$ will be called *regular (nonsingular)* if the gradient ∇f at this point is nonzero. If this condition is satisfied for any point of a level surface, we will call such a level surface *regular*.

Theorem 3.1. A regular level surface is a surface.

Proof. Let $f^{-1}(c)$ be a regular level surface and $p(x_0, y_0, z_0) \in f^{-1}(c)$. This means that $f(x_0, y_0, z_0) = c$. It follows from the assumption that the gradient of a function f is nonzero that at least one of the partial derivatives of this function f'_x, f'_y, f'_z at a point p is nonzero. Without loss of generality, we assume that $f'_z|_p \neq 0$. Now we can apply the implicit function theorem, since all the conditions of this theorem are satisfied. Therefore, there exists in E^3 a neighborhood V of point p, in which the equation f(x, y, z) = c defines z as a single-valued smooth function of the variables x, y, that is, $z = \varphi(x, y)$. This means that if we substitute the function $\varphi(x, y)$ instead of z in a function f(x, y, z), we get the identical equality $f(x, y, \varphi(x, y)) \equiv c$. The value of $\varphi(x, y)$ at a point $q = (x_0, y_0)$ of xy-coordinate plane is z_0 , i.e $\varphi(x_0, y_0) = z_0$. Let \tilde{U} denote the intersection of sets V and $f^{-1}(c)$, i.e. $\tilde{U} = V \cap f^{-1}(c)$. Let U be the orthogonal projection of \tilde{U} onto the xy-coordinate plane. Now consider a parametrized surface (ϕ, U) , where U is an open subset of the plane (in this case, it is the xycoordinate plane), and $\phi: U \to E^3$ is a mapping of U into three-dimensional space E^3 defined by $\phi(x,y) = (x,y,\varphi(x,y))$. Obviously, ϕ is a bijective smooth mapping of U onto $\tilde{U} = V \cap f^{-1}(c)$. Thus, (ϕ, U) is a local coordinate surface for $f^{-1}(c)$ and $\phi: U \to \tilde{U} = V \cap f^{-1}(c)$ is onto mapping and the theorem is proved.

3.2 FIRST FUNDAMENTAL FORM OF A SURFACE

Let us remind that E^3 is a three-dimensional Euclidean space, which means that we have the concept of length of a segment, the concept of angle between straight lines and the formula for calculating the lengths and angles in Cartesian coordinates. A surface \mathfrak{S} is in this three-dimensional space E^3 and, therefore, we can use the concepts of length and angle between straight lines in the case of a surface. But we can not just apply them to surface points, we must first adapt them for use on a surface. In other words, if we have two points on a surface, we can connect them with a straight line and calculate the length of the resulting segment. But this is not very interesting, because our goal is to measure the length of a segment lying on the

surface. A surface is generally not a linear object, such as a plane. Therefore, the line segments connecting points of a surface and lying on the surface will be curves. We already know that the calculation of the arc length of a curve should start with the differential of an arc, that is, with an infinitesimal element of an arc, and then, by integration, calculate the finite lengths. Since the concept of a length of a segment comes to our surface from the surrounding Euclidean space E^3 , we must start with an infinitesimal length in space E^3 .

Let $d\vec{r} = (dx, dy, dz)$ be an infinitesimal vector in E^3 . Then the square of its length is $||d\vec{r}||^2 = dx^2 + dy^2 + dz^2$. This expression is often called the infinitesimal metric (or simply the metric) of Euclidean space E^3 in Cartesian coordinates x, y, z. Suppose that a surface \mathfrak{S} is parametrized in a neighborhood of some of its point pby a local coordinate surface (ϕ, U) , where

$$\phi(u, v) = (x(u, v), y(u, v), z(u, v))$$

is a parametric equation of a part $\phi(U) \subset \mathfrak{S}$ of a surface. Then the differentials of coordinates x, y, z on a local coordinate surface (ϕ, U) can be expressed in terms of differentials of parameters u, v as follows

$$dx = x'_u du + x'_v dv,$$

$$dy = y'_u du + y'_v dv,$$

$$dz = z'_u du + z'_v dv.$$

Substituting the obtained expressions in $dx^2 + dy^2 + dz^2$, we get

$$dx^{2} + dy^{2} + dz^{2} = ||\phi'_{u}||^{2} du^{2} + 2 < \phi'_{u}, \phi'_{v} > du \, dv + ||\phi'_{v}||^{2} \, dv^{2},$$

where

$$\begin{aligned} ||\phi'_{u}||^{2} &= (x'_{u})^{2} + (y'_{u})^{2} + (z'_{u})^{2}, \\ <\phi'_{u}, \phi'_{v} > &= x'_{u}x'_{v} + y'_{u}y'_{v} + z'_{u}z'_{v}, \\ ||\phi'_{v}||^{2} &= (x'_{v})^{2} + (y'_{v})^{2} + (z'_{v})^{2}. \end{aligned}$$

We will use the following notations introduced by K. F. Gauss

$$E = ||\phi'_u||^2, \quad F = \langle \phi'_u, \phi'_v \rangle, \quad G = ||\phi'_v||^2.$$
(3.9)

Using these notations, we can write the square of the length of an infinitesimal vector $d\vec{r}$ as follows

$$dx^{2} + dy^{2} + dz^{2} = E \, du^{2} + 2 F \, du \, dv + G \, dv^{2}.$$
(3.10)

The expression on the right-hand side of equality (3.10) is called a *first fundamental* form of a surface or Riemannian metric of a surface. From an algebraic point of view, this expression is a quadratic form of the differentials du, dv of the surface parameters and we will denote this quadratic form by

$$\Psi_1(du, dv) = E \, du^2 + 2 F \, du \, dv + G \, dv^2. \tag{3.11}$$

The symmetric matrix of the first fundamental form will be denoted by Ψ_1 , i.e.

$$\Psi_1 = \left(\begin{array}{cc} E & F \\ F & G \end{array}\right).$$

Let us clarify the geometric meaning of the first fundamental form of a surface. Let $(\xi = \xi(t), I), \xi(t) = (u(t), v(t))$ be a parametrized curve in $U \subset E^2$ and ξ' its velocity vector field. Then $(\tilde{\xi} = \tilde{\xi}(t), I)$, where $\tilde{\xi} = \phi \circ \xi$, is the induced parametrized curve in a parametrized surface (ϕ, U) and $\tilde{\xi}'$ is velocity vector field along $\tilde{\xi}$. Hence

$$\tilde{\xi}' = u' \,\phi'_u + v' \,\phi'_v$$

and

$$\|\tilde{\xi}'\|^2 = \|\phi'_u\|^2 (u')^2 + 2 < \phi'_u, \phi'_v > u'v' + \|\phi'_v\|^2 (v')^2$$

Multiplying both sides of this equality by dt^2 and making use of du = u' dt, dv = v' dt, we get

$$\|\tilde{\xi}'\|^2 dt^2 = E \, du^2 + 2 \, F \, du \, dv + G \, dv^2.$$

On the left-hand side of this equality we have the square of the differential of an arc (1.19) of parametrized curve $\tilde{\xi}$ in a surface \mathfrak{S} . Thus, we obtain

$$ds^2 = E \, du^2 + 2 F \, du \, dv + G \, dv^2.$$

Therefore, using integration, i.e.

$$s_{\tilde{\xi}} = \int_{t_0}^{t_1} \sqrt{E \, du^2 + 2 F \, du \, dv + G \, dv^2}$$

we can calculate the lengths of curves on surface \mathfrak{S} . This means that the first fundamental form belongs to the intrinsic geometry of a surface.

The first fundamental form of a surface can be viewed from a slightly different point of view. It is well known that a scalar product in Euclidean space is a positive definite symmetric bilinear form on vectors. If we restrict this bilinear form to a tangent plane to a surface, we get a bilinear form on the tangent vectors and this bilinear form can also be called the first fundamental form of a surface. Thus, in order to compute the value of the first fundamental form on tangent vectors $\mathbf{v}, \mathbf{w} \in T_p \mathfrak{S}$, we must compute their scalar product $\langle \mathbf{v}, \mathbf{w} \rangle$. If we compute this scalar product in the basis $\{\phi'_u, \phi'_v\}$, where $\mathbf{v} = a_1 \phi'_u + a_2 \phi'_v, \mathbf{w} = b_1 \phi'_u + b_2 \phi'_v$, we get

$$< v, w >= E a_1 b_1 + F (a_1 b_2 + a_2 b_1) + G a_2 b_2.$$

Using the first fundamental form of a parametrized surface, one can calculate the areas on this surface. Let (ϕ, U) be a parametrized surface and V be an open or closed subset of U. Our aim is to calculate the area of $\phi(V)$. We cover an open subset U with a grid of coordinate lines such that a distance between two parallel u-coordinate lines is Δu and a distance between two parallel v-coordinate lines is Δv . Then an open subset V will be divided into a set of rectangles with sides $\Delta u, \Delta v$. Some of these rectangles will lie entirely inside V, and part of them will lie on the

boundary of this subset. To calculate the area of $\phi(V)$ on a parametrized surface (ϕ, U) , we use a limiting process $\Delta u, \Delta v \to 0$. In this case, the total area of the rectangles crossing the boundary of a subset V will tend to zero. Let us consider a point $q = (u_0, v_0) \in V$ on the *uv*-plane such that the rectangular with vertices $(u_0, v_0), (u_0 + \Delta u, v_0), (u_0, v_0 + \Delta v), (u_0 + \Delta u, v_0 + \Delta v)$ lies entirely inside an open subset V. Let us denote this rectangular by $\sigma_{(q,\delta)} \subset V$. Then the image of this rectangular on a surface $\phi(\sigma_{(q,\delta)})$ is a curvilinear quadrilateral. In order to find a suitable linear approximation for the area of this curvilinear quadrilateral we use the Taylor expansions

$$\phi(u_0 + \Delta u, v_0) = \phi(u_0, v_0) + \phi'_u(u_0, v_0) \Delta u + \dots,$$

$$\phi(u_0, v_0 + \Delta v) = \phi(u_0, v_0) + \phi'_u(u_0, v_0) \Delta v + \dots.$$

These expansions show that we can replace the area of a curvilinear quadrilateral $\phi(\sigma_{(q,\delta)})$ with the area of a parallelogram constructed on vectors $\phi'_u \Delta u$ and $\phi'_v \Delta v$. This gives us an approximation up to second order infinitesimals. The area of the parallelogram can be calculated by means of the cross product of the vectors $\phi'_u \Delta u$, $\phi'_v \Delta v$ and making use of the identity

$$\|\vec{a}\times\vec{b}\|^2 = \|\vec{a}\|^2 \|\vec{b}\|^2 - <\vec{a}, \vec{b}>^2,$$

where \vec{a}, \vec{b} are vectors in 3-dimensional Euclidean space, we get

$$\begin{split} \|(\phi'_u\,\Delta u)\times(\phi'_v\,\Delta v)\| &= \sqrt{\|\phi'_u\,\Delta u\|^2\|\phi'_v\,\Delta v\|^2 - \langle \phi'_u\,\Delta u,\phi'_v\,\Delta v\rangle^2} \\ &= \sqrt{EG - F^2}\,\Delta u\Delta v. \end{split}$$

Summing up the areas of these parallelograms and taking the limit of these sums when $\Delta u, \Delta v \rightarrow 0$, we get the formula for calculating the areas on a parametrized surface

$$S = \int \int_{(V)} \sqrt{EG - F^2} \, du \, dv.$$
 (3.12)

The structures on a surface that are determined by the first fundamental form of a surface belong to the *internal geometry of a surface*. Hence the length of a curve on a surface, the angle between two curves on a surface and the area of a part of a surface belong to the internal geometry of a surface.

3.3 SHAPE OPERATOR OF A SURFACE

An important tool in the study of a local structure of a surface is a shape operator. A shape operator underlies concepts such as a normal curvature of a surface and principal curvatures of a surface. A notion of shape operator is based on the idea of studying the rate of change of a unit normal vector field of a surface along a curve lying on this surface. A logical reasoning in this case is similar to the reasoning in the case of the natural parametrization of a curve, that is, since the length of a vector does not change when moving along a curve, only its direction changes and the rate of change of direction of a vector gives us information about a curvature of a surface along this curve.

Let \mathfrak{S} be a surface, $p \in \mathfrak{S}$ be a point of this surface and (ϕ, U) be a local coordinate surface in a neighborhood of p. Then we have the unit normal field on a parametrized surface (ϕ, U) induced by cross product of tangent vector fields ϕ'_u and ϕ'_{v} . This unit normal vector field will be denoted by N. The tangent plane of a surface \mathfrak{S} will be denoted by $T_p\mathfrak{S}$. Assume that we have a curve $(\xi = \xi(t), I)$ lying on a local coordinate surface $\phi(U) \subset \mathfrak{S}$ and passing through a point p when t = 0. The velocity vector of this curve at a point p will be denoted by v. Hence $\xi'(t)|_{t=0} = v$ and $\mathbf{v} \in T_p \mathfrak{S}$. At each point of a curve ξ , we have a unit normal vector of surface \mathfrak{S} . The restriction of the unit normal vector field N to a curve ξ is denoted by $\mathbb{N}(t)$, that is, $\mathbb{N}|_{\mathcal{E}(t)} = \mathbb{N}(t)$. Obviously $\mathbb{N}(t)$ is a vector field along a curve ξ . Now our goal is to investigate how fast the vector field N(t) changes in a neighborhood of a point p along a curve ξ . To do this, we use the derivative of vector field N(t) with respect to parameter t. The derivative $\mathbb{N}'(t)$ is also a vector field along a curve ξ . As $\mathbb{N}(t)$ is a unit vector field it follows from Proposition 1.4 that N'(t) is orthogonal to N(t). This means that $\mathbb{N}'(t)$ is a tangent vector field to a surface \mathfrak{S} . The derivative $\mathbb{N}'(t)$ calculated at a point p, i.e. when t = 0, will give us a characteristic of instantaneous rate of change of direction of the normal vector field at a point p in the direction of tangent vector v. This derivative is referred to as a *covariant derivative* of a normal vector field N at a point p in the direction of a tangent vector v. We will denote covariant derivative by $\nabla_{\mathbf{v}} \mathbb{N}$. Hence

$$\nabla_{\mathbf{v}}\mathbf{N} = \frac{d}{dt} \Big(\mathbf{N}(t)\Big)\Big|_{t=0}$$

There are two important points to note here. The first is that $\nabla_{\mathbf{v}} \mathbb{N}$ is a vector, and the second is that it is a tangent vector to surface \mathfrak{S} at a point p, i.e. $\nabla_{\mathbf{v}} \mathbb{N} \in T_p \mathfrak{S}$.

Thus, the covariant derivative of the normal vector field of a surface \mathfrak{S} determines the transformation of the tangent plane

$$\mathbf{v} \in T_p \mathfrak{S} \to \nabla_{\mathbf{v}} \mathbb{N} \in T_p \mathfrak{S}.$$

First of all, let us show that this transformation is linear. Let

$$(\xi = \xi(t), I), \xi(t) = (u(t), v(t)), \xi(I) \subset U,$$

be a parametrized curve in parameters plane E^2 . Assume this curve passes a point $q = \xi(t_0), t_0 \in I$ and $\mathbf{w} = \xi'(t)|_{t=t_0} = (w_1, w_2)$, where $w_1 = u'(t)|_{t=t_0}, w_2 = v'(t)|_{t=t_0}$, is the velocity vector of this curve at a point q. Denote by $(\tilde{\xi} = \tilde{\xi}(t), I)$, where $\tilde{\xi} = \phi \circ \xi$, induced parametrized curve on a surface \mathfrak{S} and $p = \phi(q)$. Then the coordinates of the velocity vector $\mathbf{v} = \tilde{\xi}'(t)|_{t=t_0}$ of induced curve $\tilde{\xi}$ at the point in the basis $\{\phi'_u|_p, \phi'_v|_p\}$ p are w_1, w_2 , i.e.

$$\mathbf{v} = w_1 \, \phi'_u |_p + w_2 \, \phi'_v |_p.$$

The covariant derivative can be expressed as follows

$$\nabla_{\mathbf{v}}\,\mathbf{N} = w_1\,\mathbf{N}'_u|_p + w_2\,\mathbf{N}'_v|_p.$$

The above formula shows that the covariant derivative $\nabla_{\mathbf{v}} \mathbb{N}$ of the normal vector field \mathbb{N} in the direction of a tangent vector \mathbf{v} linearly depends on the coordinates w_1, w_2 of vector \mathbf{v} and therefore it has the property

$$\nabla_{a\mathbf{v}+\mathbf{w}}\,\mathbf{N} = a\,\nabla_{\mathbf{v}}\,\mathbf{N} + \nabla_{\mathbf{w}}\,\mathbf{N},$$

where $a \in \mathbb{R}$ and $\mathbf{v}, \mathbf{w} \in T_p \mathfrak{S}$.

Definition 3.3. The linear transformation of tangent plane $T_p\mathfrak{S}$ of a surface

$$S: \mathbf{v} \in T_p \mathfrak{S} \to S(\mathbf{v}) = -\nabla_{\mathbf{v}} \mathbb{N} \in T_p \mathfrak{S}$$

is called a shape operator of a surface \mathfrak{S} at a point p.

Theorem 3.2. If $v, w \in T_p \mathfrak{S}$ are tangent vector to a surface \mathfrak{S} at a point p then

$$< S(\mathbf{v}), \mathbf{w} > = < \mathbf{v}, S(\mathbf{w}) >,$$

i.e. the shape operator is symmetric.

Proof. We parametrize a surface \mathfrak{S} in a neighborhood of a point p using a local coordinate surface (ϕ, U) . Then the tangent vectors ϕ'_u, ϕ'_v to a surface \mathfrak{S} at a point p form the basis for the tangent plane $T_p\mathfrak{S}$. Hence tangent vectors \mathbf{v}, \mathbf{w} at a point p is a linear combination of the vectors of basis $\mathbf{v} = a \, \phi'_u + b \, \phi'_v, \mathbf{w} = c \, \phi'_u + d \, \phi'_v$, where a, b, c, d are real numbers. One side we have

$$\begin{aligned} < S(\mathbf{v}), \mathbf{w} > &= < S(a \, \phi'_u + b \, \phi'_v), c \, \phi'_u + d \, \phi'_v > \\ &= ac < S(\phi'_u), \phi'_u > + ad < S(\phi'_u), \phi'_v > \\ &+ bc < S(\phi'_v), \phi'_u > + bd < S(\phi'_v), \phi'_v > . \end{aligned}$$

On the other side we have

$$\begin{split} < \mathbf{v}, S(\mathbf{w}) > &= < a \, \phi'_u + b \, \phi'_v, S(c \, \phi'_u + d \, \phi'_v) > \\ &= ac < \phi'_u, S(\phi'_u) > + ad < \phi'_u, S(\phi'_v) > \\ &+ bc < \phi'_v, S(\phi'_u) > + bd < \phi'_v, S(\phi'_v) > . \end{split}$$

Subtracting the second from the first, we get

 $< S(\mathbf{v}), \mathbf{w} > - < \mathbf{v}, S(\mathbf{w}) > = (ad - bc) \big(< S(\phi'_u), \phi'_v > - < \phi'_u, S(\phi'_v) > \big).$

Thus, the assertion of the theorem will be proved if we show that

$$< S(\phi'_u), \phi'_v > = < \phi'_u, S(\phi'_v) > .$$

Consider the left-hand side of this formula. It is easy to see that $S(\phi'_u) = -\mathbb{N}_u$. Indeed, by the definition of the shape operator, we should calculate the derivative of the normal vector field N in the direction of the tangent vector ϕ'_u , that is, along the *u*-coordinate line. But this will be the partial derivative of the normal vector field N with respect to parameter *u*. Similarly $S(\phi'_v) = -\mathbb{N}_v$. Since the normal vector field is orthogonal to a tangent plane we have $\langle \mathbb{N}, \phi'_v \rangle \ge 0$. Differentiating both sides with respect to u, we get

$$< \mathbb{N}_u, \phi'_v) > + < \mathbb{N}, \phi'_{vu}) > = 0.$$

So, from the formulas $- \langle \mathbb{N}_u, \phi'_v \rangle > = \langle \mathbb{N}, \phi'_{vu} \rangle > =$ and $S(\phi'_u) = -\mathbb{N}_u$ it follows that

$$< S(\phi'_u), \phi'_v > = - < \mathbb{N}_u, \phi'_v > = < \mathbb{N}, \phi'_{vu} > .$$
 (3.13)

Analogously differentiating $\langle N, \phi'_u \rangle \ge 0$ with respect to v we obtain

$$<\mathbf{N}_v,\phi'_u>+<\mathbf{N},\phi'_{uv}>=0.$$

Hence $- \langle \mathbb{N}_v, \phi'_u \rangle > = \langle \mathbb{N}, \phi'_{uv} \rangle > =$ and $S(\phi'_v) = -\mathbb{N}_v$ give

$$\langle \phi'_u, S(\phi'_v) \rangle = -\langle \phi'_u, \mathbb{N}_v \rangle = \langle \mathbb{N}, \phi'_{uv} \rangle.$$
 (3.14)

Comparing (3.13) with (3.14) and taking into account that the result of differentiation does not depend on the order in which the partial derivatives are taken, we get

$$< S(\phi'_u), \phi'_v > = < \phi'_u, S(\phi'_v) > .$$

3.4 PRINCIPAL CURVATURES OF A SURFACE

At the beginning of the previous section, we noted that a shape operator of a surface refers to a curvature of a surface in some direction. However, it is important to have a numerical characteristic of the curvature. First of all, note that the covariant derivative $\nabla_{\mathbf{v}} \mathbf{N}$ is a vector that lies in the tangent plane of the surface $T_p \mathfrak{S}$. It can be decomposed into two components, one of which will be an orthogonal projection to the direction of vector \mathbf{v} , and the other to an orthogonal direction \mathbf{v}^{\perp} . The projection to the orthogonal direction \mathbf{v}^{\perp} has no relation to a curvature of a surface in the directly related to a curvature of a surface in direction \mathbf{v} and the length of this projection is a numerical characteristic of a curvature of a surface. An orthogonal projection is easiest to calculate if we have a unit vector of a given direction, then a scalar product with such a vector will give an orthogonal projection onto that direction. This prompts us to consider the set of unit vectors in the tangent plane $T_p \mathfrak{S}$. Obviously, the end points of these vectors will lie on a unit circle centered at point p. We will denote this unit circle by \mathbb{S}_p^1 . Thus

$$\mathbb{S}_p^1 = \{ \mathbf{v} \in T_p \mathfrak{S} : ||v|| = 1 \}.$$

Definition 3.4. A normal curvature of a surface \mathfrak{S} at a point p in the direction of a tangent vector \mathbf{v} is called $\kappa(\mathbf{v}) = \langle S(\mathbf{v}), \mathbf{v} \rangle$, where $\mathbf{v} \in \mathbb{S}_n^1$.

A normal curvature of a surface \mathfrak{S} at point p can be viewed as a function on the unit circle \mathbb{S}_p^1 , that is, $\kappa : \mathbb{S}^1 \to \mathbb{R}$. Geometrically, the meaning of this function

is quite clear, it calculates the projection of the covariant derivative of the normal vector field $\nabla_{\mathbf{v}} \mathbb{N}$ (taken with minus) onto a tangent vector \mathbf{v} . Since this function is defined using a linear transformation and scalar product, this function is smooth. In particular, it is continuous. But a continuous function on a compact set, in this case the unit circle \mathbb{S}_p^1 , reaches its maximum and minimum value.

Definition 3.5. Those unit vectors in \mathbb{S}_p^1 on which function κ reaches its maximum or minimum value are called principal directions at a point p of a surface \mathfrak{S} . The values of the function κ on these vectors (or at the points of the circle \mathbb{S}_p^1 determined by the endpoints of these vectors) are called the principal curvatures of a surface \mathfrak{S} at a point p. In the case when a normal curvature of a surface \mathfrak{S} at a point p is constant, that is, does not depend on a tangent vector \mathbf{v} , then point p is called an umbilical point.

It should be noted here that a normal curvature has the property

$$\kappa(-\mathbf{v}) = \langle S(-\mathbf{v}), -\mathbf{v} \rangle = \langle S(\mathbf{v}), \mathbf{v} \rangle = \kappa(\mathbf{v}). \tag{3.15}$$

Therefore, a normal curvature of a surface does not depend on a tangent vector \mathbf{v} , but on the straight line it defines in a tangent plane. It is for this reason that the term principal direction is used. The definition of the principal curvatures of a surface has a transparent geometric meaning, which consists in the fact that the principal curvatures give the maximal and minimal values of a normal curvature of a surface at some point. All other values of a normal curvature of a surface lie between these two values. However, this definition is difficult to use to calculate the principal curvatures of a surface. Therefore, in the next theorem, we will prove that the principal curvatures are the eigenvalues of the shape operator of a surface and this will provide a practical method for calculating the principal curvatures.

Theorem 3.3. Let \mathfrak{S} be a surface, p a point of this surface. Then the principal curvatures and the principal directions of a surface \mathfrak{S} at a point p are the eigenvalues and the eigenvectors of the shape operator $S: T_p\mathfrak{S} \to T_p\mathfrak{S}$ respectively.

Proof. In the proof, it is convenient to use the identification of vectors of a tangent plane $T_p\mathfrak{S}$ with complex numbers. First of all, we will construct the axis of the Cartesian coordinate system on a tangent plane $T_p\mathfrak{S}$. As x-axis, we choose one of the principal directions in a tangent plane $T_p\mathfrak{S}$, namely, the one on which the value of the normal curvature κ is minimal. The unit directional vector of x-axis will be denoted by \mathbf{e}_1 and $\kappa(\mathbf{e}_1) = \kappa_1$ is the minimal value of normal curvature. Hence κ_1 is one of principal curvatures and for any $\mathbf{v} \in \mathbb{S}_p^1$ it holds $\kappa(\mathbf{v}) \geq \kappa_1$. As the y-axis, we take a straight line which passes through a point p and whose directional vector is $\mathbf{e}_2 = J(\mathbf{e}_1)$, where J is the complex structure of a tangent plane. Now, to each vector $\mathbf{v} = x \, \mathbf{e}_1 + y \, \mathbf{e}_2$ of a tangent plane with the origin at point p, we associate a complex number z = x + iy. Obviously this mapping $\mathbf{v} \to z$ is an isomorphism of vector spaces. If $\mathbf{v}_1 \to z_1 = x_1 + iy_1$ and $\mathbf{v}_2 \to z_2 = x_2 + iy_2$ then the scalar product $\langle \mathbf{v}_1, \mathbf{v}_2 \rangle$ can be expressed in terms of corresponding complex numbers as follows

$$z_1 \cdot z_2 = \frac{1}{2} (z_1 \, \bar{z}_2 + \bar{z}_1 \, z_2).$$

The shape operator S of a surface \mathfrak{S} at a point p can be given in terms of complex numbers as follows

$$S(z) = \zeta \, z - \theta \, \bar{z},\tag{3.16}$$

where $\zeta, \theta \in \mathbb{C}$ are constant complex numbers. The matrix of the shape operator S (we denote it by the same letter S) in the basis $\{\mathbf{e}_1, \mathbf{e}_2\}$ has the form

$$S = \begin{pmatrix} \Re(\zeta - \theta) & -\Im(\zeta + \theta) \\ \Im(\zeta - \theta) & \Re(\zeta + \theta) \end{pmatrix},$$
(3.17)

where \Re, \Im stand for the real and imaginary components of a complex number. As the shape operator is symmetric one can easily show that this leads to $\Im \zeta = 0$, that is, ζ in the case of a symmetric operator is real number $\zeta \in \mathbb{R}$. Thus the matrix of the shape operator takes on the form

$$S = \begin{pmatrix} \zeta - \Re \theta & -\Im \theta \\ -\Im \theta & \zeta + \Re \theta \end{pmatrix}.$$
(3.18)

Using a complex coordinate of a plane makes it easy to parametrize the unit circle \mathbb{S}_p^1 as follows $z = e^{i\varphi}$. With this parametrization of the unit circle \mathbb{S}_p^1 , the normal curvature becomes an explicit function of φ of the form

$$\kappa_p(\varphi) = \langle S(e^{i\varphi}), e^{i\varphi} \rangle = \frac{1}{2} \left((\bar{\zeta} e^{-i\varphi} - \bar{\theta} e^{i\varphi}) e^{i\varphi} + (\zeta e^{i\varphi} - \theta e^{-i\varphi}) e^{-i\varphi} \right)$$
$$= \frac{1}{2} (\zeta + \bar{\zeta}) - \frac{1}{2} (\bar{\theta} e^{2i\varphi} + \theta e^{-2i\varphi}) = \zeta - \Re(\theta e^{-2i\varphi}).$$

Due to the property (3.15) of normal curvature, it suffices to consider this function on the segment $0 \leq \varphi < \pi$. Recall that, according to the assumption made at the beginning of the proof, at point $\varphi = 0$ the function $\kappa_p(\varphi)$ reaches its minimum on the segment $[0, \pi)$. Hence the derivative of the function $\kappa_p(\varphi)$ at this point should be zero. We have

$$\kappa_p'(\varphi) = 2 \Re(i \,\theta \, e^{-2i\varphi}), \quad \kappa_p'(\varphi)\big|_{\varphi=0} = 2 \Re(i \,\theta).$$

From the equality to zero of the derivative $\kappa'_p(\varphi)$ at the point $\varphi = 0$ it follows that $\Re(i\,\theta) = 0$. Consequently the imaginary part of the number θ must be equal to zero, that is, $\Im\theta = 0$. Hence $\theta \in \mathbb{R}$ is a real number. Taking into account the obtained result, we see that the matrix of the shape operator (3.38) becomes diagonal, i.e.

$$S = \left(\begin{array}{cc} \zeta - \theta & 0\\ 0 & \zeta + \theta \end{array} \right).$$

This means that the vectors $\mathbf{e}_1, \mathbf{e}_2$ of the tangent plane $T_p\mathfrak{S}$ are eigenvectors of the shape operator operator S. Hence the principal direction \mathbf{e}_1 is an eigenvector of the shape operator S. From the form of matrix of the shape operator it also follows that the eigenvalue of the shape operator in the case of eigenvector \mathbf{e}_1 is equal to $\zeta - \theta$. Hence $\kappa_1 = \langle S(\mathbf{e}_1), \mathbf{e}_1 \rangle = \langle \zeta - \theta \rangle \|\mathbf{e}_1\| = \zeta - \theta$ and this is the minimum of the

function $\kappa(\varphi)$ on the segment $[0, \pi)$. Taking into account that θ is a real number, we get that function $\kappa(\varphi)$ has the form

$$\kappa(\varphi) = \zeta - \theta \,\Im(e^{-2i\varphi}) = \zeta - \theta \,\cos 2\varphi, 0 \le \varphi < \pi. \tag{3.19}$$

It is easy to show that $\theta > 0$. Indeed, $\kappa_1 = \zeta - \theta$ is a minimum value of $\kappa(\varphi)$ when $\varphi = 0$, hence $\zeta - \theta \cos 2\varphi > \zeta - \theta$ for any $\varphi \in (0, \pi)$ and, consequently, $\theta(1 - \cos 2\varphi) > 0$. Making use of $1 - \cos 2\varphi = 2 \sin^2 \varphi$, we get $\theta \sin^2 \varphi > 0$, thus $\theta > 0$. Taking this into account, as well as the graph of function $\cos 2\varphi$, we see that function $\kappa(\varphi)$ has a maximum at point $\varphi = \pi/2$ of the open segment $(0, \pi)$ and it is equal to $\zeta + \theta$. But point $\varphi = \pi/2$ corresponds to the second basis vector \mathbf{e}_2 . This means that this is the second principal direction of the shape operator and at the same time it is the eigenvector of the shape operator. Thus $\kappa_2 = \zeta + \theta$ is the second principal curvature (maximal).

In the proof of Theorem (3.3) we obtained the formula for normal curvature function $\kappa(\varphi) = \zeta - \theta \cos 2\varphi$. We write this function as

$$\kappa(\varphi) = \zeta \left(\cos^2 \varphi + \sin^2 \varphi\right) - \theta \left(\cos^2 \varphi - \sin^2 \varphi\right) = (\zeta - \theta) \cos^2 \varphi + (\zeta + \theta) \sin^2 \varphi.$$

Taking into account that $\kappa_1 = \zeta - \theta$, $\kappa_2 = \zeta + \theta$ are principal curvatures we get

$$\kappa(\varphi) = \kappa_1 \cos^2 \varphi + \kappa_2 \sin^2 \varphi, \qquad (3.20)$$

and this is called *Euler curvature formula*. Thus, the normal curvature $\kappa(\varphi)$ of a surface \mathfrak{S} in the direction that is determined by an angle ϕ , counted counterclockwise from the principal direction corresponding to the minimal value of normal curvature, is expressed by formula (3.20).

3.5 SECOND FUNDAMENTAL FORM OF A SURFACE.

Definition 3.6. Let \mathfrak{S} be a surface, p be a point on this surface, $T_p\mathfrak{S}$ be the tangent plane of \mathfrak{S} at a point p and \mathbf{v}, \mathbf{w} be two tangent vectors. Second fundamental form of a surface \mathfrak{S} at a point p is a bilinear form $\Psi_2: T_p\mathfrak{S} \times T_p\mathfrak{S} \to \mathbb{R}$ defined by

$$\Psi_2(\mathbf{v},\mathbf{w}) = < S(\mathbf{v}), \mathbf{w} > .$$

Let us derive the form of Ψ_2 in local parameters u, v of a surface \mathfrak{S} , that is, we assume that in a neighborhood of a point p a surface \mathfrak{S} is parametrized with the help of a local coordinate surface (ϕ, U) . To express the second fundamental form Ψ_2 in terms of the local parameters u, v and their differentials du, dv, we use an infinitesimal tangent vector

$$d\phi = du \,\phi'_u + dv \,\phi'_v$$

We have

$$\begin{split} \Psi_2(d\phi, d\phi) &= \langle S(du \, \phi'_u + dv \, \phi'_v), du \, \phi'_u + dv \, \phi'_v \rangle \\ &= \langle du \, S(\phi'_u) + dv \, S(\phi'_v), du \, \phi'_u + dv \, \phi'_v \rangle \\ &= du^2 \langle S(\phi'_u), \phi'_u \rangle + 2 \, du dv \, \langle S(\phi'_u), \phi'_v \rangle + \langle S(\phi'_v), \phi'_v \rangle \,. \end{split}$$

Let us introduce the notations

$$l = \langle S(\phi'_u), \phi'_u \rangle, \ m = \langle S(\phi'_u), \phi'_v \rangle, \ n = \langle S(\phi'_v), \phi'_v \rangle.$$
(3.21)

Using these notation, the second fundamental form can be written as follows

$$\Psi_2(du, dv) = l \, du^2 + 2m \, du dv + n \, dv^2. \tag{3.22}$$

The symmetric matrix of the second fundamental form will be denoted by Ψ_2 , i.e.

$$\Psi_2 = \begin{pmatrix} l & m \\ m & n \end{pmatrix} \tag{3.23}$$

3.6 THE WEINGARTEN EQUATIONS OF A SURFACE AND GEODESICS

In this section, we derive equations that can be considered as analogues of the Frenet-Serret equations in the theory of curves. Suppose we are given a parametrized surface (ϕ, U) . Then three vector fields $\{\phi'_u, \phi'_v, N\}$, where ϕ'_u, ϕ'_v are tangent vector fields and N is the unit normal vector field, are defined on this surface. These three vector fields form the frame of the three-dimensional Euclidean space E^3 at each point of a parametrized surface (ϕ, U) . This means that any smooth vector field on a surface can be represented as a linear combination of $\{\phi'_u, \phi'_v, N\}$, where the coefficients of this linear combination are some functions. In particular, if we take the partial derivatives of these vector fields with respect to the surface parameters u, v, we obtain new vector fields, and the goal of this section is to find the coefficients of the expansion of these derivatives with respect to vector fields $\{\phi'_u, \phi'_v, N\}$. To make the calculations easier to write, we will use indices and Einstein's rule of summation over repeated indices. Note that the formulas that we obtain in this section are of a general nature and are valid in Riemannian geometry.

Let \mathfrak{S} be a surface, p be a point of this surface and (ϕ, U) be a coordinate parametrized surface in a neighborhood of p. Parameters of a parametrized surface (ϕ, U) will be denoted by u_1, u_2 or by $u_i, i = 1, 2$. The partial derivatives with respect to parameters u_1, u_2 will be denoted by ∂_1, ∂_2 , that is,

$$\partial_1 = \frac{\partial}{\partial u_1}, \ \partial_2 = \frac{\partial}{\partial u_2}, \ \partial_{11}^2 = \frac{\partial^2}{\partial u_1^2}, \ \partial_{12}^2 = \frac{\partial^2}{\partial u_1 \partial u_2}, \ \partial_{22}^2 = \frac{\partial^2}{\partial u_2^2}$$

The tangent vector fields induced by curvilinear coordinates of a surface u_1, u_2 can be written as follows $\partial_1 \phi, \partial_2 \phi$. The coefficients of the first fundamental form of a surface will be denoted by $g_{11}, g_{12} = g_{21}, g_{22}$, where $g_{11} = E, g_{12} = g_{21} = F, g_{22} = G$. Hence the Riemannian metric of a surface in curvilinear coordinates u_1, u_2 can be expressed as follows $g = g_{ij} du_i du_j = g_{11} du_1^2 + 2 g_{12} du_1 du_2 + g_{22} du_2^2$, where

$$g_{ij} = \langle \partial_i \phi, \partial_j \phi \rangle$$
.

The reciprocal matrix for the second order square matrix (g_{ij}) will be denoted by (g^{ij}) , that is,

$$\left(\begin{array}{cc}g_{11} & g_{12}\\g_{21} & g_{22}\end{array}\right)^{-1} = \left(\begin{array}{cc}g^{11} & g^{12}\\g^{21} & g^{22}\end{array}\right).$$

Hence we have the relations $g_{ik}g^{kj} = \delta_i^k$. The coefficients of the second fundamental form of a surface will be denoted by $h_{11} = l, h_{12} = h_{21} = m, h_{22} = n$. Thus

$$h_{ij} = \langle S(\partial_i \phi), \partial_j \phi \rangle = - \langle \partial_i N, \partial_j \phi \rangle = \langle N, \partial_{ij} \phi \rangle.$$
(3.24)

Now a partial derivative ∂_j of a tangent vector field $\partial_i \phi$ and a partial derivative ∂_i of the unit normal vector field N are new vector fields $\partial_{ij}^2 \phi$, $\partial_i N$ on a parametrized surface (ϕ, U) and we can express them in the terms of the vector fields $\partial_1 \phi$, $\partial_2 \phi$, N. Before writing down these expressions, let us consider in more detail the derivative of the tangent vector field. We have curvilinear coordinates 1,2 on a surface 3. Tangent vector fields $\partial_1 \phi$, $\partial_2 \phi$ on a surface (ϕ, U) are tangent vector fields to the coordinate lines of the curvilinear coordinates u_1, u_2 of a surface. Thus the differentiation of these vector fields with respect to curvilinear coordinates u_1, u_2 can be considered as covariant derivatives $\nabla_{D_i}D_j$, where $D_i = \partial_i\phi$, on a surface (ϕ, U) . Thus we have a situation similar to that which we considered in the previous chapter in the section devoted to connection in curvilinear coordinates in the Euclidean space E^3 . This is the reason why we will use similar notation for the expansion coefficients, that is, we denote them by Γ_{ij}^k , where i, j, k = 1, 2. Note that in this case Γ_{ij}^k are functions of the curvilinear coordinates u_1, u_2 of a surface \mathfrak{S} . Hence we have

$$\partial_{ij}^2 \phi = \Gamma_{ij}^k \,\partial_k \phi + \sigma_{ij} \,N, \quad \partial_i N = \gamma_i^k \,\partial_k \phi + \tilde{\gamma}_i \,N, \tag{3.25}$$

where $\Gamma_{ij}^k, \sigma_{ij}, \gamma_i^k, \tilde{\gamma}_i$ are functions of curvilinear coordinates u_1, u_2 of a surface. Our aim in this section is to show that these functions can be expressed in terms of the first and second fundamental forms of a surface \mathfrak{S} . It is clear that we can consider a partial derivative $\partial_i N$ of a unit normal vector field as a covariant derivative with respect to tangent vector field $\partial_i \phi$. But we showed in (3.3) that covariant derivative of a unit normal vector field of a surface in direction of any tangent vector is a tangent vector. Hence $\tilde{\gamma}_i$ in (3.25) is zero and we have

$$\partial_{ij}^2 \phi = \Gamma_{ij}^k \partial_k \phi + \sigma_{ij} N, \qquad (3.26)$$

$$\partial_i N = \gamma_i^k \,\partial_k \,\phi. \tag{3.27}$$

Taking the scalar product of the both sides of equation (3.26) with a unit normal vector field N and making use of $\langle N, \partial_i \phi \rangle = 0$ we obtain

$$\langle N, \partial_{ij}^2 \phi \rangle = \sigma_{ij}$$

Then it follows from (3.24) that $\sigma_{ij} = h_{ij}$.

Now we scalarly multiply the both sides of equation (3.26) by tangent vector field $\partial_l \phi$ and apply $\langle \partial_k \phi, \partial_l \phi \rangle = g_{kl}$. We obtain

$$<\partial_{ij}^2\phi, \partial_l\phi>=\Gamma_{ij}^k g_{kl}.$$

We denote $\Gamma_{ij}^k g_{kl} = \Gamma_{ij,l}$. Hence $\langle \partial_{ij}\phi, \partial_l\phi \rangle = \Gamma_{ij,l}$. Now we will express the lefthand side of this equation in the terms of components of the Riemannian metric g of a surface \mathfrak{S} . For this purpose we take a partial derivative ∂_j of the both sides of relation $g_{il} = \langle \partial_i \phi, \partial_l \phi \rangle$. We obtain

$$\partial_j g_{il} = <\partial_{ij}^2 \phi, \partial_l \phi> + <\partial_i \phi, \partial_{jl}^2 \phi> .$$

Thus

$$\Gamma_{ij,l} + \Gamma_{jl,i} = \partial_j g_{il}. \tag{3.28}$$

By cyclically rearranging the subscripts i, j, l in (3.28), we get two more relations

$$\Gamma_{jl,i} + \Gamma_{li,j} = \partial_l g_{ji}, \tag{3.29}$$

$$\Gamma_{li,j} + \Gamma_{ij,l} = \partial_i g_{lj}. \tag{3.30}$$

Subtracting (3.28) from the sum of relations (3.29) and (3.30), taking into account the symmetry $\Gamma_{il,j} = \Gamma_{li,j}$, we find

$$\Gamma_{il,j} = \frac{1}{2} (\partial_i g_{lj} + \partial_l g_{ij} - \partial_j g_{il}).$$

In order to find Γ_{il}^k we multiply the both sides of $\Gamma_{il}^s g_{sr} = \Gamma_{il,r}$ by the reciprocal matrix g^{rk} . We get

$$\Gamma_{il,r} g^{rk} = \Gamma_{il}^s g_{sr} g^{rk} = \Gamma_{il}^s \delta_s^k = \Gamma_{il}^k.$$

Thus

$$\Gamma_{il}^{k} = \Gamma_{il,r} g^{rk} = \frac{1}{2} g^{rk} (\partial_{i} g_{lr} + \partial_{l} g_{ir} - \partial_{r} g_{il}).$$
(3.31)

In order to find functions γ_i^k in (3.27) we scalarly multiply the both sides of (3.27) by a tangent vector field $\partial_i \phi$. We get

$$<\partial_i N, \partial_j \phi >= \gamma_i^k g_{kj}.$$

Multiplying the both sides of this relation by g^{jl} (and taking sum over j) and using $\langle \partial_i N, \partial_j \phi \rangle = -h_{ij}$, we get $\gamma_i^l = -h_{ij}g^{jl}$. Hence

$$\partial_i N = -h_{ij} g^{jl} \,\partial_l \phi.$$

We can summarize our calculations in the form of the following theorem.

Theorem 3.4. Let \mathfrak{S} be a surface, p be a point of this surface and (ϕ, U) be a parametrized coordinate surface in a neighborhood of a point p with curvilinear coordinates u_1, u_2 . Then we have the following system of differential equations

$$\partial_{il}^2 \phi = \Gamma_{il}^k \partial_k \phi + h_{ij} N. \tag{3.32}$$

$$\partial_i N = -h_{ij} g^{jl} \partial_l \phi, \qquad (3.33)$$

where h_{ij} are components of the second fundamental form of a surface and

$$\Gamma_{il}^{k} = \frac{1}{2} g^{rk} (\partial_{i} g_{lr} + \partial_{l} g_{ir} - \partial_{r} g_{il}).$$

The second equation (3.33) of this theorem is called *Weingarten equation of a surface*.

The Weingarten equations can be derived in a more explicit way. Note that actually a matrix γ_i^k in equation (3.27) is a matrix of a shape operator and we actually solved a system of two equations by making use of the reciprocal matrix of this system g^{ij} . Let \mathfrak{S} be a surface, $p \in \mathfrak{S}$ be a point of this surface, (ϕ, U) be a local coordinate surface in a neighborhood of this point. Then tangent vectors ϕ'_u, ϕ'_v form the basis for a tangent plane $T_p\mathfrak{S}$. Our aim is to calculate the matrix of the shape operator in the basis $\{\phi'_u, \phi'_v\}$. We have

$$S(\phi'_u) = S_{11} \phi'_u + S_{21} \phi'_v, \qquad (3.34)$$

$$S(\phi'_v) = S_{12} \phi'_u + S_{22} \phi'_v, \qquad (3.35)$$

where

$$S = \left(\begin{array}{cc} S_{11} & S_{12} \\ S_{21} & S_{22} \end{array}\right),$$

is the matrix of the shape operator of a surface. Let us multiply, using scalar multiplication, the first equation (3.34) first by vector ϕ'_u , and then by vector ϕ'_v . We get

$$< S(\phi'_u), \phi'_u > = S_{11} < \phi'_u, \phi'_u > +S_{21} < \phi'_v, \phi'_u >, < S(\phi'_u), \phi'_v > = S_{11} < \phi'_u, \phi'_v > +S_{21} < \phi'_v, \phi'_v >.$$

Making use of the definitions (3.9)(3.21) of the coefficients of first and second fundamental form of a surface we get the system of linear equations

$$E S_{11} + F S_{21} = l,$$

$$F S_{11} + G S_{21} = m$$

where S_{11}, S_{21} are unknowns. Solving this system by means of determinants we get

$$S_{11} = \frac{\begin{vmatrix} l & F \\ m & G \end{vmatrix}}{\begin{vmatrix} E & F \\ F & G \end{vmatrix}} = \frac{lG - mF}{EG - F^2}, \ S_{21} = \frac{\begin{vmatrix} E & l \\ F & m \end{vmatrix}}{\begin{vmatrix} E & F \\ F & G \end{vmatrix}} = \frac{mE - lF}{EG - F^2}.$$
 (3.36)

Similarly, multiplying (with the help of scalar multiplication) the second equation (3.35) alternately by vectors ϕ'_u and ϕ'_v , then using the definitions of the coefficients of the first and second fundamental forms and solving the resulting system of linear equations, we get

$$S_{12} = \frac{mG - nF}{EG - F^2}, \ S_{22} = \frac{nE - mF}{EG - F^2}.$$
(3.37)

Thus, we have found the matrix of the shape operator of a surface in the basis $\{\phi'_u, \phi'_v\}$ of the tangent plane $T_p\mathfrak{S}$

$$S = \frac{1}{EG - F^2} \begin{pmatrix} lG - mF & mG - nF \\ mE - lF & nE - mF \end{pmatrix}$$
(3.38)

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Considering that $S(\phi'_u) = -N'_u, S(\phi'_v) = -N'_v$, we can write equations (3.34), (3.35) in the form

$$N'_{u} = \frac{mF - lG}{EG - F^{2}}\phi'_{u} + \frac{lF - mE}{EG - F^{2}}\phi'_{v}, \qquad (3.39)$$

$$N'_{v} = \frac{nF - mG}{EG - F^{2}} \phi'_{u} + \frac{mF - nE}{EG - F^{2}} \phi'_{v}.$$
 (3.40)

Equations (3.39), (3.40) are the Weingarten equations of a surface written with the help of coefficients E, F, G of the first fundamental form of a surface.

3.7 GAUSS AND PETERSON-MAINARDI-CODAZZI EQUATIONS

In this section we study the question of integrability of the equations in Theorem 3.4. Assume that we are given the system of differential equations in partial derivatives

$$\frac{\partial f_{\mu}}{\partial u} = \sum_{\nu=1}^{n} A_{\mu\nu} f_{\nu}, \quad \frac{\partial f_{\mu}}{\partial v} = \sum_{\nu=1}^{n} B_{\mu\nu} f_{\nu}, \quad (3.41)$$

where $f_{\mu}(u,v), \mu = 1, 2, ..., n$ are unknown functions of two variables u, v and $A_{\mu\nu}(u,v), B_{\mu\nu}(u,v)$ are given smooth functions. This system of differential equations is called completely integrable if

$$\frac{\partial}{\partial v} \left(\sum_{\nu=1}^{n} A_{\mu\nu} f_{\nu} \right) = \frac{\partial}{\partial u} \left(\sum_{\nu=1}^{n} B_{\mu\nu} f_{\nu} \right).$$
(3.42)

If the integrability condition (3.42) is satisfied then for any initial conditions $f_{\mu}(u_0, v_0) = a_{\mu}$, where $a_{\mu} \in \mathbb{R}$, there exists a solution for a system of differential equations (3.41) and this solution is unique.

Now we will find an integrability condition for the system of differential equations in Theorem 3.4. In the case of the first equation the integrability condition has the form $\partial_{ijk}\phi = \partial_{ikj}\phi$. Hence, taking a partial derivative ∂_k of the both sides of the first equation, we obtain the relation

$$\partial_{ijk}\phi = \partial_k \Gamma^m_{ij} \ \partial_m \phi + \Gamma^m_{ij} \ \partial^2_{mk}\phi + \partial_k h_{ij} \ N + h_{ij} \ \partial_k N$$

Substituting $\partial_{km}^2 \phi = \Gamma_{km}^r \partial_r \phi + h_{km} N$, $\partial_k N = -h_{ks} g^{sr} \partial_r \phi$ into the right-hand side of the above relation and collecting together terms containing tangent vector fields $\partial_m \phi$ and unit normal vector field N, we get

$$\partial_{ijk}\phi = (\partial_k\Gamma^m_{ij} + \Gamma^r_{ij}\Gamma^m_{rk} - h_{ij}h_{ks}g^{sm})\partial_m\phi + (\partial_kh_{ij} + \Gamma^r_{ij}h_{rk})N.$$
(3.43)

Swapping the subscripts j and k in this relation, we get

$$\partial_{ikj}\phi = (\partial_j\Gamma^m_{ik} + \Gamma^r_{ik}\Gamma^m_{rj} - h_{ik}h_{js}g^{sm})\partial_m\phi + (\partial_jh_{ik} + \Gamma^r_{ik}h_{rj})N.$$
(3.44)

Subtracting (3.44) from (3.43) we get zero at the left-hand side. Taking into account that the vector fields $\{\partial_1\phi, \partial_2\phi, N\}$ form a frame, i.e. they are linearly independent at each point of a surface, we get two equations

$$(h_{ij}h_{ks} - h_{ik}h_{js})g^{sm} = \partial_k \Gamma^m_{ij} - \partial_j \Gamma^m_{ik} + \Gamma^r_{ij} \Gamma^m_{rk} - \Gamma^r_{ik} \Gamma^m_{rj}, \qquad (3.45)$$

$$\partial_j h_{ik} - \partial_k h_{ij} = \Gamma^r_{ij} h_{rk} - \Gamma^r_{ik} h_{rj}. \tag{3.46}$$

3.8 GAUSS AND MEAN CURVATURES

In this section we will consider two important characteristics of a surface, which are a mean and Gaussian curvature of a surface. For this we use the shape operator of a surface. From the theory of linear operators, we know that by choosing a basis of a finite-dimensional linear space, we can associate each linear operator with its matrix. Obviously, this matrix depends on the choice of the basis; when passing to another basis, the matrix of the operator will change. However, there are real numbers associated with a matrix that do not change when a basis is changed. In the case of a second-order matrix, these are the determinant and the trace of a matrix. Recall that the trace Tr A of an *n*th order square matrix A is the sum of its elements on the main diagonal of a matrix, that is, $\text{Tr } A = \sum_{i=1}^{n} a_{ii}$. Since the determinant and the trace of a matrix are invariants of a transformation of a basis, we can talk about the determinant and trace of a linear operator.

Definition 3.7. The Gaussian curvature K of a surface is the determinant of the shape operator of a surface, that is, K = DetS The mean curvature H of a surface is the half of the trace of the shape operator, that is, $H = \frac{1}{2}TrS$.

We can easily find formulae to express the Gaussian K and mean H curvatures of a surface in terms of the coefficients of the first and second fundamental forms. Indeed, the Weingarten equations give us the matrix of the shape operator of a surface and, calculating the determinant and trace of this matrix, we obtain formulae for the Gaussian and mean curvatures. Hence for Gaussian curvature we get

$$\begin{split} K &= \frac{1}{(EG - F^2)^2} \Big((lG - mF)(nE - mF) - (mG - nF)(mE - lF) \Big) \\ &= \frac{lnEG - lmFG - mnEF + m^2F^2 - m^2EG + lmFG + mnEF - lnF^2}{(EG - F^2)^2} \\ &= \frac{ln(EG - F^2) - m^2(EG - F^2)}{(EG - F^2)^2} = \frac{ln - m^2}{EG - F^2}. \end{split}$$

Making use of the matrices of the first and second fundamental forms of a surface we can write

$$K = \frac{\text{Det }\Psi_2}{\text{Det }\Psi_1}.$$
(3.47)

Analogously calculating the trace of the matrix (3.38) of the shape operator we find the mean curvature of a surface

$$H = \frac{lG - 2mF + nE}{2(EG - F^2)}.$$
(3.48)

EXERCISES

3.1 A chain rule for parametrized curves. Let $(\xi = \xi(t), I)$ be a parametrized curve and $(\eta = \eta(\tau), J)$ be its reparametrization by means of a function $t = t(\tau)$. Show that the velocity vector of parametrized curve η can be expressed through the velocity vector of a parametrized curve ξ as follows $\eta'(\tau) = t'(\tau)\xi'(t)|_{t=t(\tau)}$. From this formula it follows that reparametrization preserves the regularity of a parametrized curve.

- 3.2 Length of a parametrized curve. Let $(\xi = \xi(t), I)$ be a parametrized curve and $(\eta = \eta(\tau), J)$ be its reparametrization by means of a function $t = t(\tau)$. Prove that the length of the arc of a parametrized curve ξ between two points $t_0, t_1 \in I, t_0 < t_1$ is equal to the arc of its reparametrization η between two points τ_0, τ_1 , where $t_0 = t(\tau_0), t_1 = t(\tau_1)$, i.e. the length of an arc of a parametrized curve is invariant under reparametrization.
- 3.3 Curvature of ellipse. An ellipse is a parametrized curve $(\xi = \xi(t), I)$, where $\xi(t) = (a \cos t, b \sin t), a > b > 0$ are semi-axes of ellipse and $I = [0, 2\pi]$. Find the curvature of an ellipse. Analyze the curvature formula of ellipse to find its maximum and minimum value.
- 3.4 Logarithmic spiral. Fix a point p on a ray originating from a point O. A ray starts to rotate at a constant angular velocity ω around a point O, and at the same time point p begins to move along a rotating ray (away from O) with a speed proportional to its distance from a point O. Find an equation of the trajectory of a point p in polar coordinates, where an initial position of a ray is the polar axis. This trajectory is called a *logarithmic spiral*.

Solution: $r = r_0 e^{a\phi}$, where $a = k/\omega$, k is a proportionality factor of increase of a speed with distance, ω is an angular velocity.

3.5 Curvature of a plane curve in polar coordinates. Let $r = r(\phi)$ be an equation of parametrized curve in polar coordinates. Find the formula for the curvature of this parametrized curve.

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